

RELEVANCE OF MIDLATITUDE DYNAMICS DURING BREAK-ACTIVE CYCLES OF SUMMER MONSOON, WITH EMPHASIS ON INDIA AND BEYOND

**A thesis submitted during 2015 - 2021 to the University of Hyderabad in partial fulfilment
of the award of a Ph. D. degree in Earth and Space Sciences**

by

Govardhan Dandu

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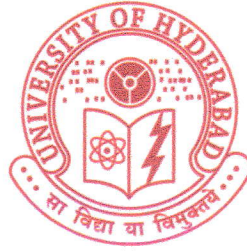
Under the supervision of

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CERTIFICATE

This is to certify that the thesis entitled **“RELEVANCE OF MIDLATITUDE DYNAMICS DURING BREAK-ACTIVE CYCLES OF SUMMER MONSOON, WITH EMPHASIS ON INDIA AND BEYOND”** submitted by **Govardhan Dandu** bearing Reg. No. **15ESPE05**, in partial fulfilment of the requirements for the award of Doctor of Philosophy in Earth & Space Sciences, is a bonafide work carried out by him under my supervision and guidance.

This thesis is not been submitted previously in part or in full to this or any other University or Institution for the award of any degree or diploma.

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This thesis is free from plagiarism and has not been submitted previously in the part of in full to this or any other University or Institution for award of any degree or diploma.

Further, the student has the following publications before submission of the thesis/monograph for adjudication and has produced evidence for the same in the form of acceptance letter or the reprint in the relevant are of this research: (Note: at least one publication in referred journal is required)

1. Govardhan, D., V. B. Rao and K. Ashok, 2017: Understanding the dynamics of breaks in the Indian Summer Monsoon and its revival. *J. Atmos. Sci.*, 74, <http://dx.doi.org/10.1175/JAS-D-16-0325.1> I.F. 3.578 3 citations. Chapter 5 of dissertation where there this publication appears.
2. Rao, V. B., Govardhan, D, Ashok, K., Mary Toshie Kayano, Julio Pablo Reyes Fernandez and Marcelo Barbio Rosa, (2020): A unified view of breaks in Indian, South America and North American monsoons. (*JGR, under review*). Chapter 3 & chapter 4 of dissertation where there this publication appears.
3. Feba, F., D. Govardhan, C. T. Tejavath, Karumuri Ashok, 2021, ENSO Modoki teleconnections to Indian summer monsoon rainfall—A review, *Elsevier*, Book title: Indian Summer Monsoon Variability ElNiño-Teleconnections and Beyond, 1st Edition, ISBN: 9780128224021. Chapter 7 of dissertation where there this publication appears.

and has made presentations in the following conferences:

1. Presented a poster at IMSP, Annual Monsoon Workshop - 2015 and National Symposium on "Understanding and Forecasting the Monsoon Extremes", at IITM, Pune during February 23-24, 2016.
2. Presented a poster at International Conference at CLIVAR Early Career Scientists Symposium at First Institute of Oceanography/State Oceanic Administration of China, Qingdao, China during September 18-25, 2016.
3. Presented a poster at TROPMET – 2016; National Symposium on Tropical Meteorology: Climate Change & Coastal Vulnerability at Indian Institute of Technology, Bhubaneswar during December 18-21, 2016.

Further, the student has passed the following course towards fulfilment of coursework requirement for Ph.D. / was exempted from doing coursework (recommended by Doctoral Committee) on the basis of the following courses during his Ph. D. program

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ES805	Research Methodology	3	Pass
ES 806	Mathematics for Earth Sciences	4	Pass
ES 807	Interdisciplinary Course	3	Pass
ES 811-830	Special paper on a specified Research Topic	2	Pass


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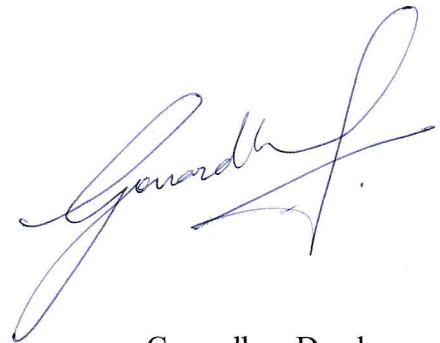

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DECLARATION

I, Govardhan Dandu, hereby declare that this thesis entitled **“RELEVANCE OF MIDLATITUDE DYNAMICS DURING BREAK-ACTIVE CYCLES OF SUMMER MONSOON, WITH EMPHASIS ON INDIA AND BEYOND”**, submitted by me under the guidance and supervision of Prof. Karumuri Ashok, Centre for Earth, Ocean and Atmospheric Sciences, School of Physics, University of Hyderabad, is a bonafide research work. I also declare that it has not been submitted previously in part or in full to this University or any other University or Institution for the award of any degree or diploma.

Date: 08/11/2021

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5	15ESPE03	K. MALLESH	Pass	Pass	Pass	Exempted	Pass
6	15ESPE04	CHENNU RAMU	Pass	Pass	Pass	Pass	Pass
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10	15ESPE10	BILLINGI TARUN KUMAR	Pass	NA	Pass	NA	Pass
11	16ESPE01	RESHMA M.R.	Pass	NA	Pass	Exempted	Pass
12	16ESPE02	SEEDABATTULA V BALAJI MANASA RAO	Pass	NA	Pass	NA	Pass
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ES 806 Mathematics for Earth Sciences -4
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07/06/2017

Dedicated to 20 ” ‘* _ _ *` “ OZ...

*“One day, in retrospect, the years of struggle will strike you as
the most beautiful.”*

— Sigmund Freud

CONTENTS

<i>Acknowledgements</i>	<i>I</i>
<i>List of Publications</i>	<i>III</i>
<i>Conference & Workshop proceedings</i>	<i>IV</i>
<i>List of Acronyms</i>	<i>VI</i>
<i>List of Symbols</i>	<i>IX</i>
<i>List of Figures</i>	<i>X</i>
<i>List of Tables</i>	<i>XXV</i>
 <i>Abstract</i>	 <i>XXVI</i>
 <i>1. Introduction</i>	 <i>1</i>
1.1 Pioneering studies of Indian summer monsoon	2
1.2 Intra-seasonal Variability of Indian summer monsoon	5
1.3 Interactions between Indian summer monsoon and Midlatitude	15
1.3.1 The impact of Jet streams on Indian summer monsoon	15
1.3.2 Relevance of Midlatitude Blocking High with Indian summer monsoon	19
1.3.3 Quasi Resonant Amplification	23
1.4 Intraseasonal variability of North and South American monsoons	25
1.4.1 Intraseasonal variability of North American Monsoon	28
1.4.2 Intraseasonal variability of South American Monsoon	30

1.5 Objectives and Scope of the study	33
1.5.1 Objectives	33
1.5.2 Scope of the present study	34
2. Data & Methodology	35
2.1 Datasets used	36
2.1.1 Observed data	36
2.1.2 Reanalysis data	36
2.2 Dates of break periods, and study region	37
2.3 Methodology applied	41
2.3.1 Dynamical analysis carried out	43
2.3.1.1 Eliassen-Palm Fluxes	43
2.3.1.2 Energetics computations	44
2.3.1.3 Condition for unstable waves	46
2.3.1.4 Condition for Barotropic instability	48
2.3.2 Statistical methods	48
2.3.2.1 One sample Student's <i>t</i> test for composite analysis	48
2.3.2.2 Students <i>t</i> test for two unequal variances	49
2.3.2.3 Pearson's correlation coefficient	49
2.4 A brief introduction to the Weather Research and Forecasting Model	50
3. Breaks in Indian summer monsoon	52
3.1 Introduction	53

3.2 Breaks in the Indian Monsoon Season: Rainfall characteristics and associated synoptic situation	54
3.3 Genesis and Maintenance of the Blocking High Episode	71
3.4 Summary of the chapter	82
4. Breaks in North and South American summer monsoons	83
4.1 Introduction	84
4.2 Breaks in North American Monsoon	85
4.3 Breaks in South American Monsoon	97
4.4 Summary of the chapter	114
5. Understanding the revival of Indian summer monsoon rainfall after breaks	115
5.1 Introduction	116
5.2 Barotropic instability in the aftermath of breaks	117
5.3 Wavelength Threshold for manifestation of a post-break synoptic disturbance	131
5.4 Summary of the chapter	137
6. Assessment of a dynamical mechanism for the revival of active summer rainfall conditions over India following a break phase	139
6.1 Introduction	140
6.2 WRF model description and sensitivity experiments performed	140
6.3 Results and Discussion	143

6.3.1	<i>Wavelength characteristics and energetics exchange for the revival of active conditions over India</i>	143
6.3.2	<i>Eddy momentum flux and rainfall characteristics during a break phase</i>	147
6.4	<i>Summary of the chapter</i>	153
7.	<i>Summary, conclusion and future scope of the thesis</i>	155
7.1	<i>Conclusion</i>	156
7.2	<i>Future scope</i>	163
	<i>Appendix-I</i>	158
	<i>Appendix-II</i>	170
	<i>References</i>	182
	<i>Similarity Index Report</i>	200

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Govardhan Dandu

LIST OF JOURNAL PUBLICATIONS

1. **Govardhan, D.**, V. B. Rao and K. Ashok, 2017: Understanding the dynamics of breaks in the Indian Summer Monsoon and its revival. *J. Atmos. Sci.*, **74**, <http://dx.doi.org/10.1175/JAS-D-16-0325.1> **I.F. 3.578** **3 citations**
2. Rao, V. B., **Govardhan, D.**, Ashok, K. , Mary Toshie Kayano, Julio Pablo Reyes Fernandez and Marcelo Barbio Rosa, (2020): A unified view of breaks in Indian, South America and North American monsoons. (**JGR, under review**)

BOOK CHAPTER

1. Feba, F., **D. Govardhan**, C. T. Tejavath, Karumuri Ashok, 2021, ENSO Modoki teleconnections to Indian summer monsoon rainfall—A review, *Elsevier*, *Book title: Indian Summer Monsoon Variability ElNiño-Teleconnections and Beyond, 1st Edition, ISBN: 9780128224021.*

OTHER PUBLICATIONS

1. Boyaj A., K. Ashok, S. Ghosh, P. Devanand, and **D. Govardhan**, 2017: The Chennai extreme rainfall event in 2015: The Bay of Bengal connection, *Climate Dynamics*. DOI: 10.1007/s00382-017-3778-7 **I.F. 4.048.** **7 citations**
2. Rao, V. B., Koteswara Rao, K, Mahendranath, B., Lakshmi Kumar, T. V., **Govardhan, D.**, (2020): Large-scale connection to deadly Indian heat waves, *Quart. J. Royal. Met. Society* (<https://doi.org/10.1002/qj.3985>)

CONFERENCE & WORKSHOP PROCEEDINGS

1. Attended an international workshop at INCOIS under the ITCO-ocean program entitled **"Indian Ocean Dynamics: From the Large-scale Circulation to Small-scale Eddies and Fronts"** during November 16-27, 2015 under Julian P. McCreary, Professor and Director, IPRC, University of Hawaii, Manoa. He also familiarized himself with the WRF model installation and preliminary analysis in this semester.
2. Presented a poster at IMSP, Annual Monsoon Workshop - 2015 and National Symposium on **"Understanding and Forecasting the Monsoon Extremes"**, at IITM, Pune during February 23-24, 2016.
3. Presented a poster in international conference at **CLIVAR** Early Career Scientists Symposium at First Institute of Oceanography/State Oceanic Administration of China, Qingdao, China during September 18-25, 2016.
4. Presented a poster at **TROPMET – 2016**; National Symposium on Tropical Meteorology: Climate Change & Coastal Vulnerability at Indian Institute of Technology, Bhubaneswar during December 18-21, 2016.
5. Participated in a national workshop on **"Extreme Weather and Climate Variability: Observation, Understanding and Prediction Training Workshop"** conducted by GIAN during December 22-31, 2016 at Indian Institute of Technology, Bhubaneswar.

6. Participated in the **“TERI-NORCE Research School: Global Climate Anomalies, Teleconnections and Regional Implications”**, during the period August 05th -08th 2019 held at TERI Institute, New Delhi.
7. Participated in a workshop on **“Thermodynamics in Earth System”**, by Prof. Axel Kleidon, Max Plank Institute, during the period 10th-15th October 2019 held at IISER, Pune.

LIST OF ACRONYMS

ASM	:	Asian Summer Monsoon
BoB	:	Bay of Bengal
CBI	:	Barotropic instability
CMKE	:	Rate of conversion of mean kinetic energy to eddy kinetic energy
CMR	:	Core Monsoon Region
CTL	:	Control experiment
DCW	:	Daily Climatological Winds
DCWNC	:	Daily Climatological Winds with convection scheme switched off
ENSO	:	El Niño-Southern Oscillation
EP	:	Eliassen - Palm
FAA	:	Federal Aviation Administration
FGGE	:	First Global Experiment
GARP	:	Global Atmospheric Research Programme
HIF	:	High-index flow
IMD	:	Indian Meteorological Department
ISM	:	Indian Summer Monsoon
ISO	:	Intraseasonal oscillations
ITCZ	:	Inter-tropical convergence zone

IV	:	Intaseasonal variability
JJAS	:	June, July, August, and September
LBA	:	Large-Scale Biosphere Atmosphere
LBC	:	Lateral boundary conditions
LIF	:	Low-index flow
LLJ	:	Low-level jet
MCZ	:	Maximum cloud zone
MJO	:	Madden Julian Oscillation
MT	:	Monsoon trough
NAM	:	North American Monsoon
NCAR	:	National Centre for Atmospheric Research
NCEP	:	National Centres for Environmental Prediction
NCEP-FNL	:	National Centers for Environmental Prediction-Final Analysis
NE-SW	:	Northeast to Southwest
NH	:	Northern Hemisphere
NW-SE	:	Northwest to South east
QRA	:	Quasi-Resonant Amplification
SACZ	:	South Atlantic Convergence Zone
SAH	:	South Asian High
SAM	:	South American Monsoon

SLP	:	Sea level pressure
SMO	:	Sierra Madre Occidental
SWJ	:	Subtropical Westerly Jet
TEJ	:	Tropical Easterly Jet
TRMM	:	Tropical Rainfall Measuring Mission
USA	:	United States of America
WETAMC	:	Wet Season Atmospheric Mesoscale Campaign
WRF	:	Weather Research & Forecasting model

LIST OF SYMBOLS

Symbol	Definition
Φ	Geopotential Height
U	Zonal wind
V	Meridional wind
$D/2$	Latitudinal distance between Jets
L_c	Critical wavelength
$\overline{K}$	Mean Kinetic energy
K'	Eddy Kinetic energy
ζ	Absolute vorticity
f	Coriolis force
m	mass
u'	eddy zonal wind
v'	eddy meridional wind

LIST OF FIGURES

Figure 1.1:	Spatial distribution of climatological mean sea level pressure (mb) during the June through September months (JJAS), over the 1979-2007 period.	3
Figure 1.2:	Climatology of rainfall (mm/day) over the Indian region for the JJAS season during the period 1901-2019.	3
Figure 1.3:	Interannual variability of the Indian summer rainfall over the Indian region during 1901-2019, provided as example of its temporal variations.	4
Figure 1.4:	Spatial distribution of mean sea level pressure and rainfall averaged over the duration of a break event (01-09 August 2000) and corresponding pre-break conditions (23-31 Jul 2000). The break event has been identified Based on the Rajeevan et al. (2010) criterion.	6
Figure 1.5:	Mean JJAS Wind vector and magnitude (m/s) at a height of 200 hPa for the 1979 - 2019 period.	16
Figure 1.6:	(a) The black dashed line box denotes 1. North American, 2. African (northern western parts of the continent) 3. Indian and 4. East Asian monsoonal regions in the northern hemisphere and the shading represents the climatology mean sea level pressure of the JJAS season for the 1979-2007 period. (b) The black dashed line box denotes the 5. South American, 6. African (middle and southern African countries), and 7. Australian monsoonal regions in the southern hemisphere, and	27

the shading represents the DJF season climatology mean sea level pressure for the 1979-2007 period.

Figure 3.1: Composite daily zonal anomalies of Geopotential height (m) over the monsoon season of 1979-2007 in the Indian region during **(a)** pre-break periods, **(b)** break periods, and **(c)** post-break periods. The hatching in the Figures 1 a-c indicates the regions where the composite Geopotential height is significantly different from zero at 95% confidence level. Statistical significance has been obtained using a two-tailed one sample Student's t test. Note that significance test has not been applied to negative values. The shaded region varies between -120 m to 120 m, with an interval of 20 m. 56

Figure 3.2: Difference between the composite daily zonal anomalies of Geopotential height (m) during the 1979-2007 period over the Indian region between **(a)** break and pre-break periods **(b)** break periods and post-break periods. Difference in composites of zonal anomalies of daily mean sea level pressure (hPa) over the Indian region during the 1979-2007 period between **(c)** during-break and pre-break periods, and **(d)** during-break and post-break periods. The shaded region in the panels (a) & (b) vary between -50 m to 50 m, with an interval of 10 m and those in panels (c) & (d) vary between -2.4 hPa to 2.4 hPa, with an interval of 0.4 hPa. The hatched regions indicate locations of significant differences, at 95% confidence level, in daily zonal anomalies of Geopotential height in panels (a) & (b), and that of daily 57

mean sea level pressure in panels (c) & (d). Statistical significance has been obtained using a two-tailed two sample Student's t-test with unequal variances.

Figure 3.3:	Difference between the composite daily zonal anomalies of Geopotential height at 1000 hPa (m) during the 1979-2007 period over the Indian region between (a) break and pre-break periods (b) break periods and post-break periods. The shaded region in panels (a) & (b) vary between -20 hPa to 20 hPa, with an interval of 5 hPa. The hatched regions indicate locations of significant differences, at 95% confidence level.	58
Figure 3.4:	Figure 3.4a: Daily geopotential height distribution at 200 hPa level for a pre-break period (23-31 July 2000) over the Indian region.	61
	Figure 3.4b: Daily geopotential height distribution at 200 hPa level for a break period (01-09 August 2000) over the Indian region.	61
	Figure 3.4c: Daily geopotential height distribution at 200 hPa level for a post-break period after (10-18 August 2000) over the Indian region.	62
Figure 3.5:	Figure 3.5a: Daily geopotential height distribution at 1000 hPa level for a pre-break period (23-31 July 2000) over the Indian region.	63
	Figure 3.5b: Daily geopotential height distribution at 1000 hPa level for a during-break period (01-09 August 2000) over the Indian region.	63
	Figure 3.5c: Daily geopotential height distribution at 1000 hPa level for a post-break period (10-18 August 2000) over the Indian region.	64

Figure 3.6:	Figure 3.6a: Daily spatial rainfall distribution for a pre-break period (23-31 July 2000) over the Indian region.	65
	Figure 3.6b: Daily spatial rainfall distribution during (01-09 August 2000) a break period over the Indian region.	65
	Figure 3.6c: Daily spatial rainfall distribution for the post-break period (10-18 August 2000) over the Indian region.	66
Figure 3.7:	Composite vertical wind zonal anomalies ($\times 10^{-3}$) (ω) (Pa/s) over Indian region and longitudinally averaged along 60°E to 100°E during the summer monsoon for the 1979-2007 period, (a) pre-break periods (b) during-break periods, and (c) post-break periods, (d) difference between the composite break and composite pre-break periods, and (e) difference between the composite break and composite post-break periods. The shaded region in panels (a) (b) & (c) vary between -120 hPa to 30 hPa, with an interval of 10 hPa and those in panels (d) & (e) vary between -6 hPa to 3 hPa, with an interval of 0.5 hPa. The hatched regions indicate composite ω zonal anomalies are significant at 95% confidence from a two-tailed Student's t-test.	67
Figure 3.8:	Composite Rainfall (mm) over Indian region during the Indian summer monsoon for the 1979-2007 period, (a) pre-break periods (b) during-break periods, and (c) post-break periods, (d) difference between the composite break and composite pre-break periods, and (e) difference between the composite break and composite post-break periods. The shaded region in the panels (a) (b) & (c) vary between 3 mm/day to 15	70

mm/day, with an interval of 3 mm/day. The shaded region in panels (d) & (e) vary between 2 to 6 mm/day & -2 to -6 mm/day, with an interval of 1 mm/day. The hatched regions indicate composite rainfall are significant at 95% confidence from a two-tailed Student's t-test.

Figure 3.9: Globally zonal averaged U wind (m/s); **(a) (upper panel):** for the composite pre-break, composite during-break & composite post-break periods of ISM (Table 1), cases available during the 1979-2007 period; **(b) (lower panel):** Similarly, for the case study, pre-break 23-31 July 2000, during-break: 01-09 Aug 2000 and post-break 10-17 Aug 2000 of ISM. 73

Figure 3.10: Wavenumber spectrum using meridional wind (m/s) at 200 hPa level for; the case study pre-break: 23-31 July 2000, composite pre-break periods during 1979-2007 of the Indian summer monsoon, climatological mean of month July during the 1979-2018 period at latitude 45° N & 50° N, and standard deviation for all the ten harmonics of July at 45° N for the period 1979-2018. 74

Figure 3.11 **(a):** Eliassen-Palm fluxes cross-section of composite during-break periods of Indian summer monsoon; **(b):** Climatological Eliassen-Palm fluxes for the summer season, JJAS during the 1979 – 2007 period. **(c):** Eliassen-Palm fluxes for the case study during the break period 01-09 August 2000 of ISM. In all the figures, the contour 76

interval is 50 m^3 . The dashed lines represent convergence and continuous lines represent divergence.

- Figure 3.12:** Vertical-latitude structure of zonal anomalies of daily U wind (m/s) which are longitudinally averaged between 65°E to 100°E , for **(a)** ccomposite pre-break, **(b)** composite during-break and **(c)** composite post-break spells of Indian summer monsoon. The hatched region is significant at 95% confidence using a two-tailed Student's t test. In all the figures, contouring vary between -5 to 5 (m/s) with an interval of 1 m/s 77
- Figure 3.13:** Structure of daily zonal wind anomalies which are zonally averaged between 60°E to 100°E , before a break **(a)** pre-break (23-31 July 2000); **(b)** during-break (01-09 August 2000) and **(c)** post-break (10-18 August 2000). 79
- Figure 3.14:** **(a)** Latitudinal variation of zonal wind (m/s), eddy momentum (m^2/s^2) transport and condition for the barotropic instability (CBI), which are zonally averaged between 60°E to 100°E ; for composite during-break periods at 200 hPa during the 1979-2007 period of ISM. **(b)** Latitudinal variation of zonal wind (m/s), eddy momentum (m^2/s^2) transport and condition for the barotropic instability (CBI) for the case study during the break period 01-09 August 2000 at 200 hPa height of ISM. 81

Figure 4.1:	Composite daily zonal anomalies of Geopotential height (m) over the monsoon season of 1979-2007 in the North American region during (a) pre-break periods (b) break periods, and (c) post-break periods. The hatching in the Figures 10 a-c indicates the regions where the composite Geopotential height is significantly different from zero at 95% confidence level. Statistical significance has been obtained using a two-tailed one sample Student's t test. In all the figures, positive values are shaded, and negative values are shown in contours. Note that significance test has not been applied to negative values. Contours vary between -120 m to 0 m, with an interval of 20 m. Same interval in shading, as evidenced by the greyscale shown, is used for positive values.	87
Figure 4.2:	Figure 4.2: Daily rainfall during dry/break period, 16-24 July 2000 over the North American region.	88
Figure 4.3:	Difference between the composite daily zonal anomalies of Geopotential height (m) during the 1979-2007 period over the North American region between (a) break and pre-break periods (b) break periods and post-break periods. Difference in composites of zonal anomalies of daily mean sea level pressure (hPa) over the North American region during the 1979-2007 period between (c) during-break and pre-break periods, and (d) during-break and post-break periods. The hatched regions indicate locations of significant differences, at 95% confidence level, in daily zonal anomalies of	90

Geopotential height in panels (a) & (b), and that of daily mean sea level pressure in panels (c) & (d). Statistical significance has been obtained using a two-tailed two sample Student's t-test with unequal variances. In all the figures, positive values are shaded, and negative values are shown in contours. The contours in panels (a) & (b) vary between -120 m to 0 m, with an interval of 20 m and those in panels (c) & (d) vary between -10 hPa to 0 hPa, with an interval of 1 hPa. Same intervals as used for contours are used for the shadings.

Figure 4.4: Composite vertical wind zonal anomalies ($\times 10^{-3}$) (ω) (Pa/s) over North American region and longitudinally averaged along 60° W - 125° W during the summer monsoon for the (a) pre-break periods (b) during-break periods, and (c) post-break periods, (d) difference between the composite break and composite pre-break periods, and (e) difference between the composite break and composite post-break periods. The contours in panels (a) (b) (c) (d) & (e) vary between -30 Pa/s to 0 Pa/s, with an interval of 10 Pa/s. The hatched regions indicate composite ω zonal anomalies are significant at 95% confidence from a two-tailed Student's t test. **91**

Figure 4.5: Wavenumber spectrum using meridional wind (m/s) at 200 hPa level; for composite pre-break periods of North American monsoon (Table 2.3); case study pre-break period, 10-15 Jul 1995; the climatological mean of month July during 1979-2018 periods at 45° N and 50° N, and **92**

standard deviation for all the ten harmonics of July at 45° N for the period 1979-2007.

- Figure 4.6:** (a) Eliassen-Palm fluxes cross-section of composite during-break periods of NAM; (b): Eliassen-Palm fluxes for the case study during the break period 16-24 July 2000 of NAM. In all the figures, the contour interval is 50 m^3 . The dashed lines represent convergence and continuous lines represent divergence. **94**
- Figure 4.7:** (a)(Top) Composite globally zonal averaged U wind (m/s) of composite pre-break, composite during-break and composite post-break periods of North American monsoon; (b)(Bottom) Globally zonal averaged U wind of pre-break periods in three different cases: 10-16 July 1981, 20 - 24 July 1995 and 10-15 July 2000 of NAM. **95**
- Figure 4.8:** Zonal wind profiles during the break phase for 1981, 1995 and 2000 events. **96**
- Figure 4.9:** Composite structure of daily U wind zonal anomalies (m/s) for the break cases of NAM, 10-16 July 1981, 20 - 24 July 1995 and 10-15 July 2000; which are zonally averaged in between 60° W - 125° W. The hatched region indicates 95% confidence using a two-tailed Student's t test. In this figure positive values are shaded and negative values are shown in contours. Contouring between -5 to 0 (m/s) with an interval of 1 m/s **96**
- Figure 4.10:** Composite daily zonal anomalies of Geopotential height (m) over the monsoon season of 1979-2007 in the South American region during **99**

(a) pre-break periods (b) break periods, and (c) post-break periods. The hatching in the Figures 17 a-c indicates the regions where the composite Geopotential height is significantly different from zero at 95% confidence level. Statistical significance has been obtained using a two-tailed one sample Student's t test. In all the figures, positive values are shaded, and negative values are shown in contours. Note that significance test has not been applied to negative values. Contours vary between -120 m to 0 m, with an interval of 20 m. Same interval in shading is used for positive values.

Figure 4.11: (a) Geopotential Height anomalies (m) at 200 hPa during the break spell, 21-26 January 1980 over the South American region. (b) Geopotential Height anomalies (m) at 200 hPa during the break spell, 06-11 January 1981 over the South American region. **100**

Figure 4.12: Difference between the composite daily zonal anomalies of Geopotential height (m) during the 1979-2007 period over the South American region between (a) break and pre-break periods (b) break periods and post-break periods. Difference in composites of zonal anomalies of daily mean sea level pressure (hPa) over the South American region during the 1979-2007 period between (c) during-break and pre-break periods, and (d) during-break and post-break periods. The hatched regions indicate locations of significant differences, at 95% confidence level, in daily zonal anomalies of Geopotential height in panels (a) & (b), and that of daily mean sea level **102**

pressure in panels (c) & (d). Statistical significance has been obtained using a two-tailed two sample Student's t-test with unequal variances. In all the figures, positive values are shaded, and negative values are shown in contours. The contours in panels (a) & (b) vary between -100 m to 0 m, with an interval of 20 m and those in panels (c) & (d) vary between -25 hPa to 0 hPa, with an interval of 5 hPa. Same intervals as used for contours are used for the shadings.

Figure 4.13: Composite vertical wind zonal anomalies ($\times 10^{-3}$) (ω) (Pa/s) longitudinally averaged along 60° W to 125° W over South American region during the summer monsoon for the 1979-2007 period, **(a)** pre-break periods **(b)** during-break periods, and **(c)** post-break periods, **(d)** difference between the composite break and composite pre-break periods, and **(e)** difference between the composite break and composite post-break periods. The contours in panels (d) & (e) vary between -30 Pa/s to 0 Pa/s, with an interval of 10 Pa/s. The hatched regions indicate composite ω zonal anomalies are significant at 95% confidence from a two-tailed Student's t-test. **103**

Figure 4.14: Wavenumber vs. Amplitude spectrum using meridional wind (m/s) at 200 hPa, for an individual case 1981 at 45° N, composite of pre-break periods of South American monsoon (Table 3), the climatological mean of month July during 1979-2018 periods at 45° N and 50° N, and standard deviation for all the ten harmonics of July at 45° N for the period 1979-2007. **105**

- Figure 4.15** **(a):(Top)** Globally zonal averaged U wind (m/s) of composite pre-break, composite during-break and composite post-break periods of SAM **(b): (Bottom)** Globally zonal averaged U wind (m/s) for pre-break periods of six different cases of SAM: 18-21 January 1980, 01-04 January 1981, 10-15 January 1988, 16-20 January 1990, 25 December-01 January 1993, 19-24 January 1996. **106**
- Figure 4.16:** **(a)** (Top) Composite of globally zonal averaged U wind (m/s) for during-break in six different cases. **(b)** (Bottom) Globally zonal averaged U wind (m/s) for post-break periods in six different cases: 18-21 January 1980, 01-04 January 1981, 10-15 January 1988, 16-20 January 1990, 25 December-01 January 1993, 19-24 January 1996. **107**
- Figure 4.17:** **(a)(Top)** Pre-break composite structure of U wind zonal anomalies of South American monsoon (Table 3) (m/s), **(b)(Bottom)** During-break composite structure of zonal wind anomalies of South American monsoon (m/s) (Table 3), which are zonally averaged between 60° W to 125° W. The hatched region indicates 95% confidence using a two-tailed Student's t test. In both the figures positive values are shaded and negative values are shown in contours. Contouring between -5 to 0 (m/s) with an interval of 1 m/s. **109**
- Figure 4.18:** Structure of zonal wind anomaly on 4 January 1981, which is zonally averaged over 270° E- 330° E. **110**

Figure 4.19:	Eliassen Palm fluxes for (a) 18-26 July 1980 & (b) 05-10 July 1981. The dotted lines denote convergence and continues lines denote divergence and contouring at 50 m ³ .	111
Figure 4.20:	Eliassen -Palm flux cross-sections of southern hemisphere (a) Composite during-break periods (Table 2.4) of SAM; (b) during the season (austral summer) DJF for the 1979-2007 period. The contour interval is 50 m ³ . In both the figures dashed lines represent convergence and continuous lines represent divergence.	112
Figure 5.1:	Observed Sea Level Pressure (SLP) distribution on 10 Aug., 2000, after a break period.	117
Figure 5.2:	Conversion of Kinetic Energy (J/s) anomaly values of 41 break periods during the 1979-2007 period.	120
Figure 5.3:	(a-e): Spatial distriubution of zonal wind at 200 hPa for 41 break events (Rajeevan et al. 2008) during the 1979–2007 period for the region 0°–50°N, 20°–120°E.	125
Figure 5.4:	(a-e): Spatial distribution of zonal wind on a peak days of of 41 break events (Rajeevan et al. 2008) during the 1979–2007 period for the region 0°–50°N, 20°–120°E.	128
Figure 5.5:	(a) Zonal wind at 200hpa on 4august, 2000, a typical break day; (b) the corresponding Geopotential distribution (in Km). The black line shows the NE-SW tilt orientation in both the figures.	128
Figure 5.6:	(a) Eddy momentum flux transfer during a break (1-9 August); before the break (14-23 July); and after the break (10-15 August), in relation	129

to the 1-9 August break in the year 2000. **(b)** zonal wind profiles (82.5° E) at 200hpa during a break day (4th August); on a day in the pre-break period; (23rd July); on a post-break day (11th August), all for the same event presented in Fig. 5.6(a).

- Figure 5.7:** **(a)** Meridional distribution of the composite absolute vorticity (200 hPa), obtained by compositing it over all the break periods for the period of 1979-2007. **(b)** Same as Fig. 5.7(a) but composited over each break periods during 1979-2007. **130**
- Figure 5.8:** Latitudinal distance (degrees) between westerlies and easterlies at 200 hPa for various phases associated with observed break events. **133**
- Figure 5.9:** Wavelength anomalies (Km) of the zonal wind at 200 hPa, averaged over the span of each break event **136**
- Figure 6.1:** Daily mean C_k , the rate of conversion of mean kinetic energy to eddy kinetic energy (J/s) during break event from 31 July-09 August 2000 over core monsoon region of India, from the NCEP II reanalysis data, and from the CTL, DCW, and DCWNC experiments. The abscissa shows the dateline from 31st July-12th August 2000 and the ordinate shows the rate of conversion of energy. **145**
- Figure 6.2:** Sea level pressure on 10 August 2000 over the Indian region from (a)NCEP II reanalysis data, (b) DCW, (c) DCWNC, and (d) CTL experiments **146**
- Figure 6.3:** Evolution of the eddy momentum flux (m^2/s^2) through the break event 01-09 August 2000 over the core monsoon region of India using NCEP **147**

II, DCW, DCWNC, and CTL. The abscissa shows the dateline from 31st July-12th August 2000 and the ordinate shows the rate of conversion of energy.

- Figure 6.4:** Zonal wind (u; Black) and Condition for barotropic instability (CBI) (red) (/s) for break event 01-09 August 2000 over core monsoon region of India using (a)NCEP II, (b) DCW, (c) DCWNC, and (d) CTL **149**
- Figure 6.5:** Geopotential height (Km) at 200 hPa on break event 01-09 August 2000 of (a) NCEP II, (b) DCW, (c) CTL, and (d) DCWNC **150**
- Figure 6.6:** Geopotential height (m) at 200 hPa on break event 01-09 August 2000 of (a) difference between DCW-NCEP II, (b) difference between CTL-NCEP II, and (c) difference between DCWNC-NCEP II **150**
- Figure 6.7:** Zonal wind (m/s) at 200 hPa on break event 01-09 August 2000 of (a) NCEP II, (b) DCW, (c) CTL, and (d) DCWNC **151**
- Figure 6.8:** Zonal wind (m/s) at 200 hPa on break event 01-09 August 2000 of (a) difference between DCW-NCEP II, (b) difference between CTL-NCEP II, and (c) difference between DCWNC-NCEP II **151**
- Figure 6.9:** Daily Rainfall over core monsoon region of India, using NCEP II, DCW, CTL, and DCWNC. The abscissa shows the dateline from 31st July-12th August 2000 and the ordinate shows the rate of conversion of energy. **152**

LIST OF TABLES

Table 2.1:	A list of source/weblinks of datasets used in the present study.	37
Table 2.2:	A list of break days based on Rajeevan et al., 2010 for the period 1979-2007 (J-July & A-August) during Indian summer monsoon.	38
Table 2.3:	List of North American break monsoon days	40
Table 2.4:	List of South American break monsoon days	41
Table 5.1:	Gives the conditions after break period for every event of break period during 1979-2007 period. ' *** ' represents the revival of ISM without low formation.	118
Table 5.2:	Conversion of Mean Kinetic Energy Values (Joule/second) at 200 hPa.	121
Table 5.3:	The channel width (D/2) between the subtropical westerly jet and tropical easterly jet (°) at 200 hPa for 41 break periods during the 1979-2007 period.	132
Table 5.4:	Critical wavelength (L_c Km) for pre-break, break and post-break periods.	135
Table 6.1:	List of physics schemes/parametrizations used in the sensitivity experiments of WRF model	141
Table 6.2:	A list of Experiments carried out using WRF model.	142
Table 6.3:	Latitudinal distance (D/2) (°) and Critical wavelength (L_c) (Km) during the break period B2000 in the NCEP II reanalysis data and conducted sensitivity experiments.	144

ABSTRACT

The summer monsoon Intraseasonal variability is primarily associated with active and break spells of rainfall. In this thesis, I investigate a few noteworthy aspects of the intraseasonal variability of the Indian summer monsoon in connection to midlatitude upper level circulation. I conducted several dynamical and statistical analyses, and sensitivity simulations with a regional model, to this end. The Indian summer monsoon rainfall (ISMR) is essentially a primary source of basic human life and a long-standing source of reliance for the Indian economy. My primary focus in this research is on the initiation and development of a break phase, as well as the revival of active conditions following a break phase over the Indian subcontinent. Furthermore, I investigated the break monsoon dynamics of North and South American summer monsoon systems and explored commonalities to the Indian summer monsoon.

The monsoon breaks in the core monsoon region or central region of India, lasting a few days-rarely even up to two weeks, are periods of low to no rainfall. Throughout summer monsoon, a seasonal high pressure region, extending from the Tibetan plateau, prevails over north India at 200 hPa. Our analysis of various observed and reanalyses data sets for the 1979-2007 period shows that breaks are characteristically preceded by an upper-tropospheric transient blocking high over north India, which extends to low levels. The formation of this blocking high is found to be due to the penetration of mid-latitude baroclinic waves into lower latitudes. Importantly, this transient blocking high is weaker than the seasonal High pressure at this level. The weakening of the high pressure due to the presence of this transient blocking leads to a sinking motion relative to the prebreak or after-break period which in turn leads to a filled up surface low, and an eventual

reduction in daily rainfall, Importantly, this mechanism is found to be associated with monsoon breaks in the North America, and South America as well.

The blocking high over the core monsoon region during breaks is associated with the convergence of Eliassen-Palm fluxes, which causes the upper-level zonal westerlies to decelerate. This holds the blocking high in place over the core monsoon region without being advected by the zonal wind. Also, I find that the resonance and amplification of planetary Rossby waves play a vital role in the block formation over the northwest of India. This sequence of events associated with the Indian breaks, that is, the initiation due to the formation of upper level blocking high due to resonance of planetary Rossby waves at mid-latitudes, followed by the break event and a relative sinking over the core monsoon region, also explains the North and South American summer monsoon breaks.

Following the initiation of a break phase over the Indian region, I observe an anomalous southward shift of a subtropical westerly jet stream during the peak of a significant break phase, followed by an anomalous northward shift of a stronger-than-expected tropical easterly jet stream. These major changes during a break facilitate an instability mechanism, which evidently contributes to the development of a synoptic disturbance, and subsequent revival of the Indian summer monsoon into the active phase in about 61% of the cases.

Our computations of energetics indicate an increase in the eddy kinetic energy at the expense of the mean kinetic energy during the breaks, in agreement with the formation of the synoptic disturbance. This demonstrates that barotropic instability in the presence of a monsoon basic flow is the primary physical mechanism that controls the subsequent revival of the monsoon. This entire life cycle of monsoon break formation and reversal to normal monsoon can be seen as the

barotropic adiabatic cycle, similar to the index cycle of midlatitudes in a barotropic atmosphere. I propose these oscillations as ‘Monsoon Index Cycles’.

Apart from the above dynamical analysis carried out using NCEP reanalysis 2 data and gridded Indian summer monsoon rainfall data sets for the 1979-2007 period on the revival phase of monsoon from the break, I conducted a few sensitivity experiments using WRF regional climate model in order to ascertain the aforementioned dynamical mechanism for the revival of active conditions after a break phase of Indian summer monsoon, for one typical break case 01-09 August 2000. The sensitivity experiments essentially examine the role of the break-related circulation, particularly the role of the horizontal shear at upper level between the subtropical Westerly jet and tropical Easterly jet, eddy momentum flux and associated energetics. The results from the model confirm my interpretation of the mechanisms from analysis of observational and reanalysis data.

Chapter 1

Introduction

Outline of the chapter:

In this chapter, I discuss the available literature on the intraseasonal variability of the Indian summer monsoon with main focus on active to break conditions. I also discuss the intraseasonal variability of South and North American summer monsoonal regions.

Based on the scientific gaps identified during the review of the literature, we developed four objectives and further discussed the scope of the thesis. It will also be interesting to see if some of the mechanisms that cause the Indian summer monsoon, North American monsoon, and South American monsoon breaks have little in common and can be unified into a theory.

1.1 Pioneering studies of Indian Summer Monsoon

Scientific interest in the monsoon is not new. The earliest theory explaining the monsoons, which is based on the importance of the differential heating between ocean and land, was espoused by Edmund Halley (1686). Eliot (1884) has been credited with the identification of “minor cyclones of the Southwest monsoon season” which we now refer to as monsoon depression. Since then, there has been tremendous interest in studying the monsoon, its variability, and its relevance for applications.

Fig 1.1 shows the mean sea level pressure during June, July, August, and September (JJAS) months, which together comprise the Indian summer monsoon (ISM) season, for the 1979-2007 period. The low-pressure zone across the Indian region is conventionally referred to as a monsoon trough (MT) (Fig. 1.1.), which is a part of the Inter-tropical convergence zone (ITCZ), and is associated with the large-scale monsoon flow. To distinguish the presence of the ITCZ at the equatorial region for a few days and times during the same season, Gadgil (2003) proposed the nomenclature of convective tropical convergence zone. Fig. 1.2, based on the Indian Meteorological Department (IMD) observations for the 1901-2019 period, shows the mean seasonal rainfall over India. It is clear that there is a lot of spatial variability of rainfall distribution during the Indian summer monsoon. The Indian summer monsoon also varies on different time scales. Fig. 1.3, for example, shows the interannual variability of rainfall anomalies over the Indian region, indicating normal, above normal, and below-normal rainfall years.

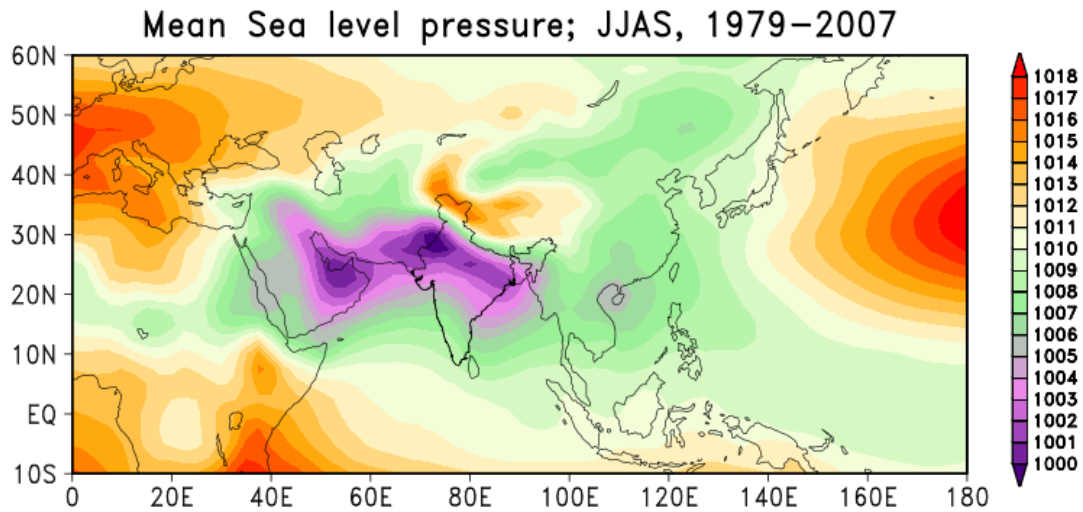


Figure 1.1: Spatial distribution of climatological mean sea level pressure (mb) during the June through September months (JJAS), over the 1979-2007 period.

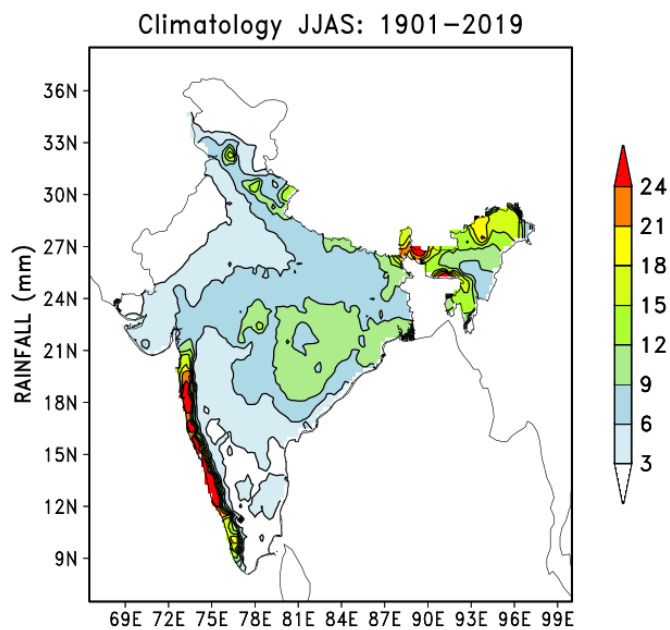


Figure 1.2: Climatology of rainfall (mm/day) over the Indian region for the JJAS season during the period 1901-2019.

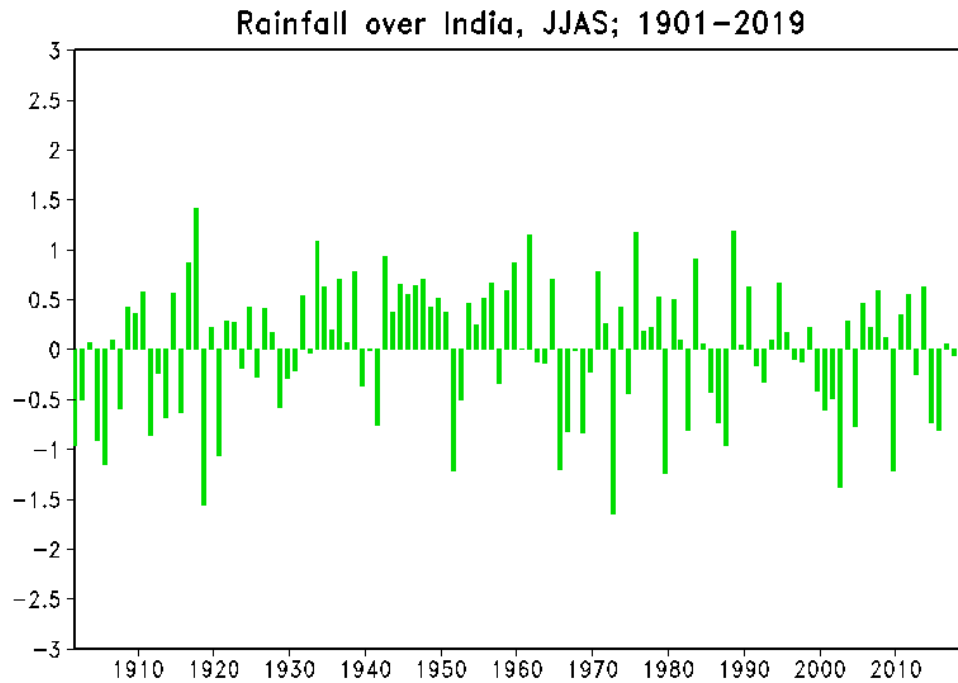


Figure 1.3: Interannual variability of the Indian summer rainfall over the Indian region during 1901-2019, provided as example of its temporal variations.

Recent investigations (Gadgil 2008), to be sure, have pointed out that while the summer onset in India is fundamentally affected by the differential heating of land and ocean, but its evolution is rather more significantly related to the complex high-frequency atmospheric disturbances, extra-monsoonal perturbations, and by the teleconnections between various climate drivers and monsoonal region. Studies suggest that the ISM rainfall varies at interannual, intraseasonal, decadal and multi-centennial time scales. From the statistical sense, a dominant mode of the Indian summer monsoon is interannual variability (Fig. 1.3), driven by the well-known El Niño-Southern Oscillation (ENSO), etc. The variability is forced by both external and internal drivers (e.g., Webster et al., 1998; Pant and Rupa Kumar, 1997). In addition to the interannual variability, the

ISM also varies on intraseasonal time scales (e.g., Chatterjee and Goswami, 2004). For example, there are cycles of active and break episodes. “During the active monsoon spell, the Indian region receives adequate rainfall over the majority of the Indian region; however, during the break, the rains are reduced or no none across most of the Indian region, primarily the monsoonal core region (central parts) of India” (Ramamurthy 1969, Goswami et al. 1998 and Lawrence & Webster 2001). This is also a major mode of variability of the ISM.

Therefore, in this thesis, my primary focus lies on the intraseasonal variability in the ISM rainfall, specifically to document the hitherto unexplored potential mechanisms behind the break monsoon and the consequent revival of the Indian summer monsoon. In the next sub-section, we review the available literature on the observational and theoretical aspects of break phase initiation and its development, and those on the subsequent revival of active conditions of ISM rainfall.

1.2 Intraseasonal Variability of rainfall during the Indian summer monsoon

During summer monsoon months, there occur a few low-rainfall periods/spells or void situations over most of the Indian region. Blanford (1886) named these “intervals of drought” during July & August, the peak summer monsoon months, as the break periods. These intervals with low rainfall situations over a major portion of the Indian region are now routinely called “breaks”. Breaks are typically announced as when the MT shifts northward from its mean position for at least 3 days. Consequently, the Himalayan foothill region receives high rainfall during this period, while the rainfall along the climatological MT position decreases drastically. For example, Fig. 1.4 depicts the mean sea level pressure during the break (01-09 August 2000) and prebreak period (23-31 July 2000), as well as the corresponding rainfall distribution over India. This break period was generated using the Rajeeven et al. (2010) criterion. In Fig. 1.4, we see the MT

weakening or shifting towards the Himalayan foothills during the break period, with massive rainfall across the Himalayan foothills. This causes a decrease in rainfall in the central parts of India, as well as the majority of the country. The monsoon revives after the break periods, often entering into the ‘active’ phase, which is, in a sense, exacerbation of the mean conditions.

Several researchers (e.g., Goswami et al. 2018) suggest that the total rainfall during any summer monsoon season critically depends on the number and intensity of the breaks. Therefore, understanding the intraseasonal variability of the ISM and diagnosing the mechanisms is important to enrich the seasonal prediction of the ISM, which has immense economic value. Thus, it is clear that the ISM rainfall is irregular in its occurrence and intensity. However, there is a lot of temporal and spatial variability in the ISM.

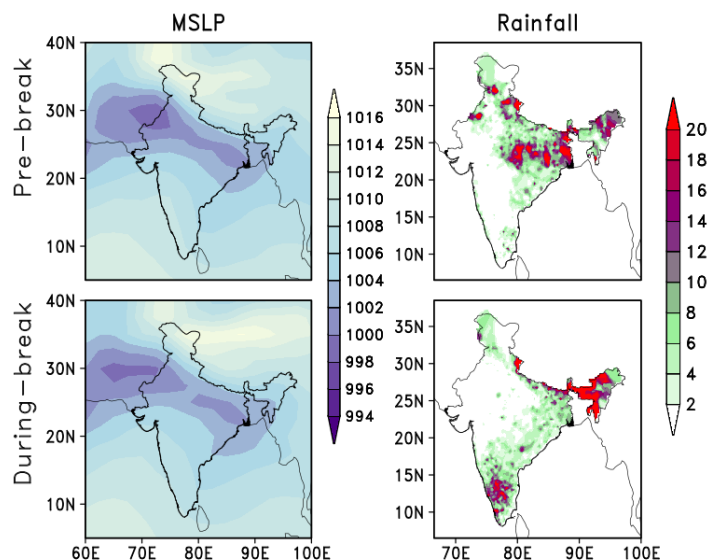


Fig. 1.4: Spatial distribution of mean sea level pressure and rainfall averaged over the duration of a break event (01-09 August 2000) and corresponding prebreak conditions (23-31 Jul 2000). The break event has been identified Based on the Rajeevan et al. (2010) criterion.

530 As per Koteswarm (1950), who studies 19 break events, the 'break phase' begins two days before
531 or later than the formation of a 'low' at 700 hPa in low latitudes. The movement of such lows
532 weakens the North-south pressure gradient over the Peninsula. According to Koteswaram (1950),
533 the monsoon frequently revives after a “break” by the formation of a synoptic low or a depression
534 in the North Bay of Bengal (BoB) or a shift towards the north of the BoB from lower latitudes,
535 resulting in the gradual revival of the MT to its normal position.

536 Using IMD observations, Ramamurthy (1969) classified active and 'break' periods based on MT
537 fluctuations in the apex monsoon months of July and August from 1888 to 1967. Ramamurthy
538 (1969) adopted the criteria of the movement of the MT moving northward from its mean position,
539 i.e., close to the foot hills of Himalayas for more than two days as a 'break', and a southward shift
540 from its mean position as an active event. He also claims that August is slightly more prone to
541 longer 'break' periods. Later, Ramaswamy (1971) observed a trough (active) and ridge (break)
542 pattern of westerly waves in the lower troposphere at 500 hPa near the Himalayas in a case study
543 during the 1965 summer monsoon using satellite data. Raghavan (1973) suggests that the break and
544 active phase of ISM are associated with the fluctuation of MT, and with the tropical synoptic low-
545 pressure systems developing in the Bay of Bengal region. Then the normal/active monsoon
546 conditions of rainfall are re-established when the monsoon trough returns and intensifies over
547 India. When a synoptic low-pressure system moves in the northerly direction towards the
548 Himalayas, then the monsoon trough of low pressure also moves from the plains to Himalayan
549 foothills. This weakens rainfall over the plains of northern India brings about the break-monsoon.

550 Ramanadham et al. (1973), studying a single case study 1967 summer monsoon, suggested that
551 the break monsoon situation sets in over the Indian region in association with the eastward
552 movement of the high amplitude 500 hPa middle-latitude westerly trough, that extends into

western parts of Pakistan and northern India. At the same time, Ramanadham et al. (1973) observed the subtropical anticyclonic ridge over Arabian region extends into central and peninsular India. At this particular time, the monsoon trough moves northward from its normal location over the Gangetic plains to the foothills of Himalayan region triggering break monsoon conditions.

The Monsoon Experiment was conducted during 1979 as a regional sub-program of the First GARP (Global Atmospheric Research Programme) Global Experiment (FGGE). This program has sourced valuable first time data related to the ISM and led to further investigation in different aspects. Studies from the project suggest that during the boreal summer monsoon, the intraseasonal variability of Asian Summer Monsoon (ASM) is dominated by the quasi-periodic intraseasonal oscillations (ISO) on time scales longer than 10 days but shorter than a season (Krishnamurti and Bhalme 1976; Krishnamurti and Ardanuy 1980; Yasunari 1980 and Goswami & Ajaya Mohan 2001). Apart from the important 30-60 day oscillations, some of the experiments were carried out to explore the 10-20 day oscillation of the Indian monsoon as the intraseasonal variability is extremely important for rainfall predictions and their socioeconomic applications (Murakami 1979; Das 1979 and Goswami 1997).

Krishnamurti and Ardanuy (1980) suggest that the break spells of ISM are accompanied by westward propagating trough-ridge systems of surface pressure with a periodicity of about 10–20 days, which have a steady quasi-linear propagation of phase. From a power spectral density analysis, they revealed that zonal wavenumbers 3–6 made a significant contribution to this quasi-biweekly mode. Krishnamurti and Ardanuy (1980) also suggest that the extrapolation of the phase information of the steady westward propagating quasi-biweekly mode predicts monsoon breaks moderately well. Sikka & Gadgil (1980) showed that there are two favourable and important locations for maximum cloud zone (MCZ) during JJAS for 1973–1977 period over the Indian

576 region, i.e., over the longitudes 70° - 90° E and in the monsoon zone north of 15° at 700 hPa, and
577 the secondary MCZ in the equatorial region, i.e., between $0-10^{\circ}$ N. In monsoon months JJAS, a
578 break phase develops just before the temporary disappearance of the MCZ from the mean summer
579 monsoon location, i.e., over the latitudes 15° - 28° N, and the active monsoon is established again
580 by the northward shift of the secondary MCZ over the monsoon zone.

581 Results from Pant (1983) suggest that Tibetan Plateau cooling (warming) leads to break (active)
582 periods over the Indian region. He showed that a ‘break’ period is associated with a zonally
583 oriented trough between 10° - 15° N at 700 hPa over south Indian region and the adjoining seas;
584 there is also a sharp decline in the temperature over the Tibetan Plateau region at 500 hPa level.
585 Subsequently, as the plateau warms, the trough shifts northward over the North Indian region and
586 leads to normal/active monsoon around 20° N. As per Pant, the Hadley cell has its upward limb in
587 the trough at $10^{\circ}-15^{\circ}$ N over south India and descent further north during the break monsoon
588 conditions. Whereas during the active phase of the monsoon, the Hadley cell is associated with a
589 rising limb at $\sim 20^{\circ}$ N, and southward descent is prominent.

590 Yasunari (1986) demonstrated the inter-correlations between ISM active/break cycles and westerly
591 circulation changes in the middle and high latitude, which are relevant to the low frequency 30–
592 50 day oscillations. The oscillation of the monsoon system of this 30-50 day period is closely
593 related with the east-west oscillation of the geopotential height with the node over Tibet. The lag
594 correlations between the northern hemisphere monsoon trough and the 500 hPa geopotential
595 height, suggested that this east-west oscillation is part of the response of the middle and high
596 latitude westerly waves to the northward moving monsoon heat source. Yasunari (1986) suggest
597 that the transition from the active to break phase of ISM, may be regarded as the transformation in

598 the thermodynamical process from the "diabatic limit" to the "advective limit" (Webster, 1981)
599 and vice versa.

600 Using atmospheric models, in a case study of the 1994 monsoon, Rodwell (1997) proposed a
601 mechanism that includes the infusion of dry, high negative potential vorticity air from the Southern
602 Hemisphere midlatitudes into the low-level monsoon inflow. Rodwell (1997) even suggests that
603 altering moisture fluxes in South Asia and the perturbation activity of Southern Hemisphere
604 midlatitudes associated with the cross-equatorial flow of monsoon can trigger monsoon breaks.
605 Furthermore, De et al. (1998) discovered a few more additional features of the monsoon basic
606 flow, which in turn leads to the monsoon break conditions during the 1968–1997 period, such as,
607 a) The movement of cyclonic circulation or a trough over low latitudes, b) existence of a trough in
608 mid-tropospheric westerlies over northern India, c) weak pressure gradient over the west coast of
609 India, and d) strong westerly winds over the northern India.

610 Krishnan et al. (2000) defined break. The study, using both observations and modelling
611 experiments, suggests that a quick northwest movement of anomalous Rossby waves from the Bay
612 of Bengal into northwest and central India initiates break conditions over Indian region.
613 Importantly, the intensification of “convective stable anomalies” over the Bay of Bengal triggers
614 these anomalous Rossby waves. Goswami and Ajaymohan (2001) using dynamical climate
615 models, that both the intraseasonal and interannual variability is governed by the mode of spatial
616 variability. As a result, if the frequency of occurrence of active (break) phases of Intraseasonal
617 oscillations (ISO) is larger than that of the opposite phases in a season, the seasonal mean monsoon
618 could be strong (weak). In further analysis, they have shown that the strong (weak) monsoon years
619 are indeed characterized by a higher frequency of occurrence of active (break) conditions and vice-
620 versa. Gadgil and Joseph (2003) obtained a significant negative correlation between all India

621 summer monsoon rainfall and the number of rain break days between 1901-1989 (i.e., $r = -0.56$).
622 They also obtained a significant positive relationship between active days and monsoon rainfall
623 between 1901-1989 ($r = 0.47$). Gadgil and Joseph's (2003) rainfall break composite is very similar
624 to Ramamurthy's (1969) rainfall break composite, as they have identified positive rainfall
625 anomalies over the Himalayan foothills and southeastern peninsular region.

626 AS per Joseph and Sijikumar (2004), when the strong cross-equatorial low-level Jetstream (LLJ)
627 oriented southeastwards with its core at 850 hPa changes its direction towards east between Sri
628 Lanka and the equatorial region, it triggers a break monsoon condition. During active monsoon
629 conditions, the LLJ axis passes from the central Arabian Sea eastwards through peninsular India
630 and it provides moisture for increased convection in the Bay of Bengal and the formation of a
631 monsoon depression. Ramesh Kumar & Dessai (2004) analyzed break monsoon cases during the
632 1901-2002 period. They find that during the breaks, the overall daily Indian rainfall is < 9 mm/day
633 and this condition lasts at least three days during July and August. The majority of the breaks lasted
634 for 3–4 days duration (49%) in the July and August months. Furthermore, the correlation between
635 the JJAS rainfall and break (active) days is -0.80 (0.38).

636 In a case study by Rao et al (2004), the authors suggest that the upper troposphere relative humidity
637 reduces between 600 hPa and 200 hPa by $\sim 30\%$ at least ~ 3 days before the break period over the
638 ITCZ along with a weakening of the low-level convergence by $\sim 50\%$. These lead to the drying of
639 the middle troposphere and trigger break monsoon conditions over the Indian region. Wang et al,
640 (2005) suggest that ISO of the Indian monsoon is associated with western equatorial Indian ocean
641 warming. As the beginning of the rain, conditions in the western equatorial Indian Ocean is
642 preceded by local surface wind convergence and central equatorial Indian Ocean warming. a “self-

643 induction mechanism” of eastern Indian ocean warming appears to be operating in the maintenance
644 of the active-break cycle of monsoon.

645 Mandke et al (2007) suggested a criterion for identifying active and break days based on the rainfall
646 anomalies over the Indian core region (73° – 82° E, 18° – 28° N). The periods were acknowledged as
647 active (break) events, when the standardized rainfall anomaly over the Indian core/central region
648 exceeds (below) 0.7 (-0.7) for at least three uninterrupted days between 15 June - 15 September.
649 Using observations of daily gridded rainfall datasets from 3700 stations, Krishnamurthy and
650 Shukla (2000) shown the above normal rain over central India and below normal rain across
651 Himalayan foothills (North India) and southern peninsular India, and that this pattern is transposed
652 during the break phase. The authors also suggest that the largescale droughts are associated with
653 huge negative rainfall anomalies across the central Indian region. Krishnamurthy and Shukla also
654 pointed out that the success in long-range forecasting will essentially depend on precise
655 quantitative estimates of the external factors affecting ISM due to high nonlinearity in the
656 intraseasonal prediction of ISM.

657 Krishnamurthy and Shukla (2007) found two dominant intraseasonal oscillations with periods of
658 45 and 20 days in the seasonal rainfall of ISM, using multichannel singular spectrum analysis of
659 the daily rainfall anomalies. Interestingly, they suggest that the active and break monsoon events
660 contribute very little to the seasonal mean rainfall of 45-day and 20-day oscillations. They
661 discovered that the lows and depressions formed during these active and break periods are
662 consistent, and that the number of depressions formed during the active phase is approximately
663 seven times that of the break phase. Rajeevan et al. (2010) suggest the active and break events are
664 the normalized anomaly of the rainfall over a critical area, called the monsoon core zone (18° – 28°

N & 65°-85° E), which exceeds 1.0 or is less than -1.0 respectively, provided that this criterion is satisfied for at least three consecutive days.

The Madden Julian Oscillation (MJO) is a dominant mode of tropical intraseasonal variability in the tropics. While it is relatively during the boreal summer, studies suggest that it is associated with northward propagation of cloud-bands over the Indian monsoon region. The work by Joseph et al. (2009) indicates that very long break periods can be deemed as a divergent Rossby response due to a dry phase of equatorial convection which can be due to a prevailing MJO signal. Pai et al., (2011) claim that 83% of the break cases are coincident with Phases 7, 8, 1, and 2 of the MJO. They also suggest an association between the active phase of the ISM with certain phases of the MJO. Madhu and Bhatla (2020) suggest a stronger association of the MJ with breaks than with active monsoon conditions over the Indian region. However, despite the aforementioned apparent association between the active-break cycle and MJO, it is possible that a particular phase of MJO in the equatorial Indian Ocean is actually triggered by the intraseasonal phase of the ISM itself, or that both of these are part of an encompassing phenomenon of the same periodicity.

Umakanth et al. (2014) suggested a new criterion for breaks during the ISM, primarily based on the threshold of area-averaged rainfall over each grid considered. They also suggest a strong association between breaks and total seasonal summer monsoon rainfall in India. Umakanth et al. (2019) suggest that the development of a mid tropospheric cyclonic circulation anomaly stretching over the sub-tropical/mid-latitude Asian continent may trigger break conditions over the Indian region.

Interestingly, they propose a strong association between El Niño and below-average monsoon rainfall. Pai et al. (2014)'s analysis of long term trends in lows, also suggests, in agreement with Raghavan (1969) the importance of the lows and depressions for revival after a break phase.

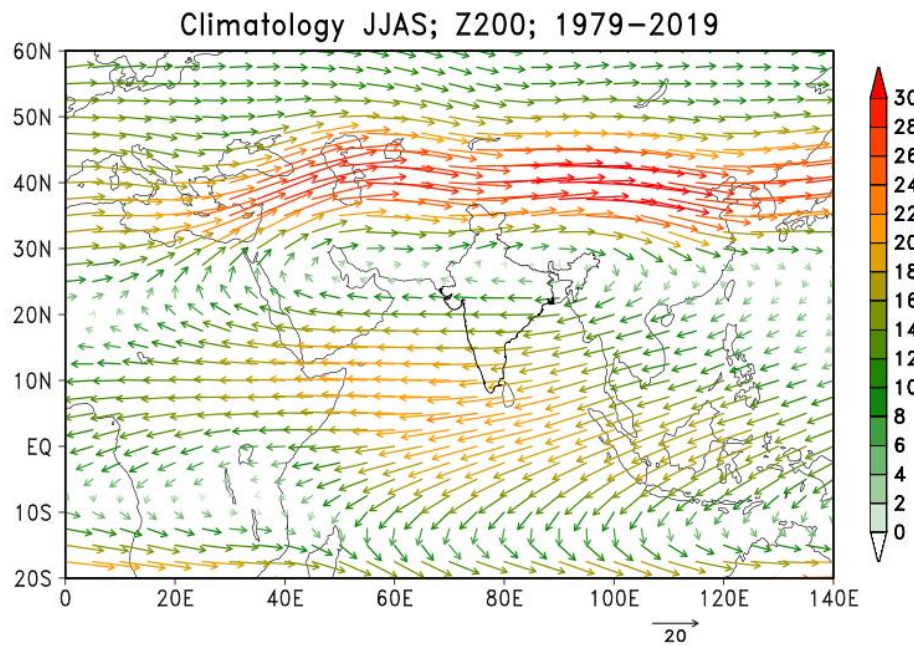
The spatial and temporal variability of ISM is unique throughout the world, owing to all of the complex features of the monsoon basic flow. Apart from the theories of the break phase discussed above, the Indian northern boundary is also associated with midlatitude interactions during ISM. In this context, the following subsection discussed the interactions between midlatitudes and ISM, prominently the role of upper tropospheric Jet streams and Blocking high.

1.3 Interactions between Indian summer monsoon and Midlatitudes

1.3.1 The impact of Jet streams on the Indian summer monsoon

Essentially, Jet streams are formed by the thermal contrast between the equator & the polar regions as a result of differential solar insolation and role of Coriolis force acting on the moving masses. These are characteristically continuous over long distances, but a discontinuity is also common. The pathway of the jet is normally meandering, and these meanders propagate eastward at slower speeds than the actual wind within the flow. These large meanders of waves within the Jetstream are known as long or ultralong Rossby waves or planetary waves. In the Northern hemisphere, the intraseasonal variability of ISM is associated with the upper-level Subtropical Westerly Jet (SWJ) and the Tropical Easterly Jet (TEJ). The SWJ core sustains at 200 hPa at 42° N, while the TEJ core is found at 150 hPa at 9° N during summer months i.e., June, July, August, and September (JJAS). For example, Fig. 1.5 shows the climatological wind vectors at 200 hPa for the season JJAS during the period 1979-2019. From Fig. 1.5, I observe a large anticyclonic circulation over the Indian region during climatological JJAS season, which is normally called

707 South Asian High (SAH). The SAH intensity depends on the distribution of heterogeneous diabatic
708 heating over the Tibetan Plateau region and the summer monsoon regional area (Wu and Liu 2003;
709 Wu et al. 2007). The SAH is another typical synoptic system over the Asian continent during
710 boreal summer (Mason and Anderson 1963), which indeed forms near the tropopause over the
711 Indo-Chinese peninsular region before the BoB monsoon onset (He et al. 2006; Liu et al. 2009,
712 2012). Figure 1.5 depicts the climatology of SWJ ranging from 30° to 40° N, as well as TEJ ranging
713 from 0° to 20° N. Large meanders of Rossby waves, however, extend beyond 50° N, accompanied
714 by a spatial shift of the jet core in day-to-day SWJ and TEJ beyond 20° N, due to diurnal variation,
715 seasonal shift, climate change, and so on. The detailed discussion of SAH and its influence over
716 the Indian region during the summer monsoon is detailed in the next subchapter.



717
718 **Figure 1.5:** Climatological mean JJAS wind vector and magnitude (m/s) at a height of 200 hPa
719 for the 1979 - 2019 period.

According to Rao et al. (1970) the changes in SWJ and TEJ flow result in four kinds of manifestations: a) intensifying or developing lower tropospheric lows or troughs, b) enhancing rainfall in pre-existing systems, c) re-curvature of depressions and lows and d) the onset of break conditions. Furthermore, even when the break is traced to westerly troughs, rainfall is increased by mid-latitude troughs before the break occurs over Indian region. During breaks, the subtropical ridge becomes more prominent and swings southwards. As a result, whether or not a break occurs after the passage of a westerly trough may be determined by the intensity of the ridge in the rear rather than the passage of the trough itself. Even after a break burst of rain along the Himalayas moving eastwards appear to be associated with the passage of subsequent troughs in westerlies, which are perhaps unable to break up the warm high that had formed earlier. A warm high forming and strengthening the seasonal subtropical ridge may explain 'breaks' that last for one to two weeks at times. The passage of a trough cannot account for the subsequent lengthy break. However, it is consistent with the behaviour of warm highs.

The STJ is a narrow band of fast-moving air that flows from west to east. Wind magnitude in a westerly jet stream typically varies between 40~60 m/s. Malurkar (1950) demonstrated that during ISM, the southward extension of an extratropical westerly trough into northern India results in a break condition. Later, Ramaswamy (1962) discovered that during a summer 'break' monsoon over India, large amplitude troughs in the middle latitude westerlies protrude into the Indo-Pakistan area based on break phase case studies during ISM. These eastward moving Rossby waves get retarded and elongated meridionally during movement across Tibetan Plateau and weakening Tibetan high at 500 hPa. Ramaswamy (1962) also showed that the ridge in the rear of the trough causes the extension of the anticyclone from the Tibetan plateau over to the northwest and central India. The associated upperlevel divergence causes heavy rainfall over and near the Himalayas. In

743 a few case studies, Ramaswamy (1962) also observed that the large amplitude trough remains
744 quasi-stationary over the northern India same as the 'break' lasts for a week or more. Thus, during
745 break conditions, I witness the remarkable spectacle of two jets of completely different types - the
746 Easterly and sub-tropical Westerly jets - passing within a short latitudinal distance of each other
747 and dynamically interacting with each other.

748 Bedi et al. (1981) noticed that a strong monsoon is associated with strong westerlies and weak
749 monsoon is associated with weak westerlies and westerly systems of middle-latitude are more
750 active during a weak monsoon. Sikka and Grossman (1981) found that a large amplitude trough in
751 the regime of subtropical westerlies was present during the break period of monsoon during the
752 year 1979. Further, very weak spells of the ISM are associated with the signature of what is called
753 as a low-index Rossby regime in the atmospheric general circulation over northern as well as
754 southern hemispheres (Ramaswamy and Pareek,1978).

755 The TEJ is a critical feature of the ISM circulation (Koteswaram 1958; Alaka 1958; Krishnamurti
756 and Bhalme, 1976) with the strongest winds of $\sim 0-50$ m/s positioned just to the west of the
757 southern tip of India and adjoining Arabian Sea (Reiter, 1961). The first detailed study on the TEJ
758 was carried out by Koteswaram et al. (1958), who showed that the TEJ is sourced due to the large-
759 scale arrangement of landmasses, the ocean, and the elevated heat source of the Tibetan Plateau.
760 Chen and Yen (1991) noticed that the interannual variability of TEJ follows the interannual
761 variability of stationary eddies over the summer monsoon region. Earlier studies also explained
762 that the interannual variation of TEJ can be related to the interannual variation of summer rainfall
763 (Chen and Yen 1991; Chen and Van Loon 1987; Kobayashi 1974). According to these studies,
764 TEJ is found to be weaker (stronger) during the deficient (excess) summer monsoon rainfall over
765 India.

The TEJ has a distinct intraseasonal oscillation south of its core, according to Chen and Van Loon (1987). The systematic shifting of locations caused by convection, low-level monsoon flow, and TEJ during different phases of the monsoon suggests that they are all possibly related (Sathiyamoorthy et al., 2007). During the weak (strong) monsoon season of 1987 (1988), the middle and upper tropospheric meridional temperature gradient between the Tibetan High and the Indian Ocean region decreased (increased) (Pattanaik and Satyan 2000). There have been few studies on the relationship between the long-term trend of TEJ and the Indian summer monsoon due to a lack of persistent long-term data over the upper troposphere.

Sathiyamoorthy, et. al. (2007) has shown that the 30-60 day ISO of the TEJ reported in individual case studies occurs during the 1979–1990 period of study. The Axis of the TEJ, at the east of $\sim 70^\circ$ E, and at 200 hPa is found along the near-equatorial latitudes during monsoon onset/revivals and the axis of TEJ propagates northward as the monsoon basic inflow advances across India. Also, they have found the axis is along $\sim 5^\circ$ N and $\sim 15^\circ$ N during active and break monsoon situations, respectively.

The next sub-chapter discusses extratropical systems, such as Blocking high and its interaction with ISM, particularly from the context of intraseasonal variability.

1.3.2 Relevance of Midlatitude Blocking High with Indian summer monsoon

The extratropical systems are associated with (a) troughs and (b) ridges of the SWJ. It may be presumed that the ridges can sometimes develop into a Blocking and the northern Indian region is just at the periphery of this system. Atmospheric blocking (hereafter as blocking) is a large-scale quasi-stationary low-frequency circulation pattern occurring in midlatitudes with a lifetime of ~ 10 to 20 days (Berggren et al. 1949; Rex 1950; Shukla and Mo 1983).

788 Blocking is defined as a continuous occurrence of high geopotential height for a relatively
789 persistent time in a confined region or a quasi-stationary form (Hartmann & Ghan 1980). Early
790 investigators (Berggren et al., 1949; Elliott and Smith, 1949 and Rex, 1950a, b) ascertained the
791 synoptic behaviour and blocking climatology. Berggren et al. (1949) demonstrated that the
792 formation of a blocking ridge is associated with the deepening of some extratropical wave cyclones
793 over central Europe. This quasi-stationary blocking 'wave' disrupts the zonal flow of fast-moving
794 frontal waves approaching from North America, resulting in extreme weather events. According
795 to Rex (1950a), blocking occurs when the basic westerly current at upper level splits into two
796 branches, each of which transports a significant amount of mass. The double-jet system must cover
797 at least 45° of longitude. Across the current split, a sharp transition from zonal type flow upstream
798 to meridional flow downstream must be observed, and this pattern must persist for at least ten
799 days.

800 Rex (1950b) noted that blocking most frequently occurs over the North eastern portions of the
801 Atlantic Ocean at 10° W and Pacific Ocean at 150° W (their Fig. 1), and normally this blocking
802 persists for 12 – 16 days and relatively stable in position. According to Rex (1950b), the
803 climatological mean position of the block is just downstream from the normal position of the major
804 mid-latitude jet streams, with noticeable warming occurring in the northern part of the blocked
805 zone and cooling occurring in the southern parts. During recent years it has also been shown that
806 planetary-scale waves influence atmospheric blocking (Lejenas and Madden, 1989). Using the
807 barotropic channel model, Egger (1978) showed that blocking could result from the nonlinear
808 interactions between the forced stationary waves and the slowly moving free waves. Using the
809 equation of motion, Hasanean and Hafez (2003) explained how the main westerly air current splits
810 into two branches leading to the formation of a blocking high. Later, Shutts (1983) conducted a

811 few numerical experiments to confirm a hypothesis that barotropic eddies reinforced blocking flow
812 patterns mostly through Reynolds stresses. This propagation of barotropic eddies into the split
813 Jetstream amplifies vorticity fields, resulting in blocking enhancement.

814 Tanaka (1998) defines atmospheric blocking as an abnormally persistent, quasi-stationary
815 anticyclone that blocks the mid-latitude jet stream and the onset of blocking is generated by Rossby
816 wave breaking. A blocking anticyclone forms at latitudes higher than where the normal sub-
817 tropical high forms and is frequently accompanied by a cutoff low in low latitudes. However, in
818 the Southern Hemisphere (SH), blocking highs tend to have a shorter lifetime than in the Northern
819 Hemisphere (NH) and form at somewhat lower latitudes in general (Van Loon, 1956; Taljaard,
820 1972).

821 Another characteristic feature widely recognized is that even when a blocking high does break
822 down or move away, there is a strong tendency for another high to form and intensify in the same
823 location (Taljaard, 1972; Wright, 1974). Blocking phenomenon in a barotropic atmosphere can
824 result due to resonant enhancement of Rossby lee waves forced by two stationary sources of
825 potential vorticity (Kalney and Merkine 1981). Resonant interactions among planetary-scale
826 waves may be an important physical mechanism in generating certain atmospheric blocking
827 (Colliucci et al., 1981). A block may be initiated by a deep cyclonic storm development, which
828 locks the flow pattern into a blocking equilibrium, or vice-versa when it drives the flow out of a
829 blocking equilibrium (Charney et al., 1981). All the above discussion gives a detailed idea of
830 blocking high genesis and its implications with the surrounding areas.

831 The blocking during the summer months is critical to the monsoonal regions, affecting the
832 monsoon basic flow directly or indirectly. This blocking is also known as SAH over the Asian

region. Wei et al. (2015) demonstrated a diagnostic analysis on the interannual time scale of SAH and its movement southeast-northwest, which is closely related to the Indian and East Asian summer monsoon rainfall. The SAH's reallocating towards southeast (northwest) is closely related to weak (more) Indian summer monsoon rainfall and high (less) rainfall in the Yangtze River valley of the East Asian summer monsoon region. There have only been a few studies that look at the relationship between ISM blocking and intraseasonal variability. They are as follows:

Unninayar and Murakami (1978) found that during weak monsoon situation anticyclonic circulation (at 200 hPa) over Himalaya bifurcates into two cells due to penetration of the midlatitude trough into sub-tropics, and at lowerlevel 700 hPa, they found that the easterlies around 20° N and 30° N over India are replaced by westerlies and north-westerlies intensified in the Arabian Sea. Keshvamurty et al. (1980) noticed a shift in the quasi-stationary large-scale features, i.e., Tibetan anticyclone and other monsoon disturbances of extratropical regions at 300 hPa and 200 hPa of tropospheric circulation in active-break monsoon situations. They also suggest that such shifts of quasi-stationary flow features and the associated changes in meridional and vertical shears may have a significant effect on summer monsoon circulation with the interactions of midlatitude westerly regime.

Raman and Rao (1981) suggested that the prolonged breaks causing severe summer monsoon droughts over the Indian sub-continent are due to upper tropospheric blocking high over East India (90° to 120°E) associated with baroclinic waves. Thus, this blocking high is considered as the initiator of the monsoon break. Recent studies have shown, the extreme cold spells during winter (Buehler et al. 2011), European temperature extremes during spring (Brunner et al. 2017), and heatwaves during summer (Della-Marta et al. 2007; Schaller et al. 2018) are often related to the formation and maintenance of mid-high latitude blocking, this has been an important research topic

in past decades (for example, Yeh 1949; Shutts 1983; Colucci 1985; Haines and Marshall 1987; Holopainen and Fortelius 1987; Luo 2000, 2005; Luo et al. 2014, 2019; Zhang and Luo 2020; Nakamura and Huang 2017, 2018; Aikawa et al. 2019; Paradise et al. 2019).

All these studies indicate that the Indian monsoon is more than a local phenomenon and interacts significantly with midlatitude circulation. We see a potential impact of upper-level jet streams and blocking high on the monsoonal circulation over the Indian region based on the literature reviewed above. In a similar vein, in the following sub-chapter 1.3, we will discuss whether the intraseasonal variability of North and South American monsoons has few similarities with ISM.

1.3.3 Quasi-Resonant Amplification

Another important aspect of the atmospheric upper level midlatitude phenomenon is Quasi-Resonant Amplification (QRA). In general, large-scale midlatitude atmospheric circulation is associated with free and forced Rossby waves (Charney and Drazin, 1961; Hoskins and Karoly 1981; Dickinson 1970; Held et al., 2002 and Branstator 2002). Petoukhov et al. (2013) hypothesized the effect of QRA of planetary waves and recent occurring weather extremes. A brief explanation of QRA hypothesis as follows

Assume, k and m are the zonal wave numbers of free synoptic waves and forced/quasi-stationary planetary-scale Rossby waves. The upperlevel atmospheric circulation is characterized as per Petoukhov et al. (2013)

- free synoptic-scale Rossby waves with zonal wave numbers $k \geq 6$ propagate mainly in the longitudinal direction with a phase speed $c \approx 0$, and

876 ▪ forced/quasi-stationary planetary-scale Rossby waves with zonal wave numbers m ,
877 with a phase speed $c \approx 0$, and frequency $\omega \approx 0$, are response of quasi-stationary
878 spatially inhomogeneous diabatic sources/sinks and orography.

879 The midlatitude free synoptic-scale waves with zonal wavenumbers $k \approx 6 - 8$ are normally weak,
880 with the meridional velocity less than $1.5 - 2$ m/s (Eliassen and Machenhauer (1965); Fraedrich
881 and Böttger (1978); Whitham (1960)).

882 Charney and DeVore (1979) shown that there is a “low-index” flow with a relatively weaker zonal
883 component and a strong wave component which is intertwined close to linear resonance; the other
884 is a “high-index” flow with a weak wave component and a relatively stronger zonal component
885 which is much distant from linear response of resonance using a barotropic channel model with
886 topographical forcing to study the planetary-scale circulations. It is suggested that the phenomenon
887 of blocking is a metastable equilibrium state of the low-index near-resonant character.

888 Later, Petoukhov et al. (2013) proposed that the extreme summer events occur due to the certain
889 persistent high-amplitude wave structures evolved in the field of the large-scale midlatitude
890 atmospheric meridional velocity to which the quasi-stationary component of free synoptic waves
891 with $k \approx 6 - 8$ made an exceptionally large contribution. These structures may arise from changes
892 in the midlatitude zonal mean state. When the aforementioned changes result in latitudinal trapping
893 within the midlatitude waveguides of quasi-stationary free synoptic waves with zonal
894 wavenumbers $k \approx 6 - 8$, the typically weak midlatitude response of wave numbers m 6 - 8 to quasi-
895 stationary thermal and orographic sources/sinks may be strongly amplified via quasi-resonance.

896 Hence the QRA plays an important role in extreme weather events. In recent years, the Northern
897 Hemisphere has experienced several regional weather extremes during summer, such as the

heatwave in the United States in 2011, the Russian heatwave and the Indus River flood in Pakistan in 2010, and the European heatwave in 2003. In this regard, I began to investigate the genesis of quasi-stationary anticyclones (blocking highs) as a result of the QRA, followed by the break phase development during the summer monsoon.

Kornhuber et al. (2016) showed that QRA explains roughly a third of all high-amplitude events with wave numbers 6 to 8 and that amplitudes of quasi-stationary waves of that wave number range are significantly higher when resonance conditions are met. They also showed that planetary waves amplified by QRA exhibit specific preferred phases due to the characteristic stationary orographic forcing of the Northern Hemisphere midlatitudinal orographic profile. Their analysis shows that the central United States and western Europe are specifically susceptible to persistent heat waves during QRA episodes (Kornhuber et al., 2017). Later Mann et al. (2018) have shown that the zonally averaged surface temperature field can be used to define a signature for the occurrence of QRA by examining the Coupled Model Intercomparison Project Phase 5 climate model projections. Mann et al. (2018) also suggest that QRA events are likely to increase by 50% this century under business-as-usual carbon emissions, but there is significant variation among climate models. The following subchapter discusses the importance of NAM & SAM.

1.4 Intraseasonal variability of North and South American monsoon regions

The classical monsoon definition was solely based on the annual reversal of prevailing surface winds (Ramage 1971). Monsoons are not the same throughout the tropics because they are influenced by the specific locations of continents, oceans, and regional wind and rain patterns. Figure 1.6, for example, depicts the climatology mean sea level pressure across global tropical regions. The black dashed line box in Fig. 1.6 (a) represents the North American, North African,

Indian, and East Asian monsoonal regions during the JJAS season, and in Fig. 1.6 (b) represents the South American, South African, and Australian monsoonal region during the DJF season. In this section, we discuss the monsoons of both Americas in comparison to ISM. Even though the geographical characteristics of specific monsoon regions significantly differ due to the vast continental region and much larger ocean basins that are adjacent to the monsoonal regions. In Fig. 1.6, we see a high on both sides of the American continent in the mean sea level pressure field, which is not seen in the Indian, East Asian, or African summer monsoons. In this regard, the environments of the monsoon systems are not comparable, but nonetheless have. Of course, they also have a few general similarities.

North American Monsoon (NAM) and South American Monsoon (SAM) circulation systems have a few similarities to other monsoons around the globe, which develop in response to thermal contrast between the continent and adjacent oceanic regions and provide a major component of continental warm season precipitation regime. Similar to other tropical monsoons, the evolution of the NAM and SAM system can be characterized in terms of onset, mature/peak, and subsidence phases. Importantly, A similar characteristic of intraseasonal variability in different American regions reveals a considerable commonality in the events occurring in ISM, despite the monsoons' distinct climatic conditions. Several examinations support the idea that intraseasonal rainfall within the NAM and SAM regions shows oscillatory behavior, though the specification of the preferred periodicities remains uncertain, the monsoon season consists of irregular burst and break periods and that the length (and number) of these periods is an important factor concluding the overall quality of the monsoon season. The goal of this research is to better understand the similarities between ISM, NAM, and SAM, in the genesis of a break phase during the summer monsoon.

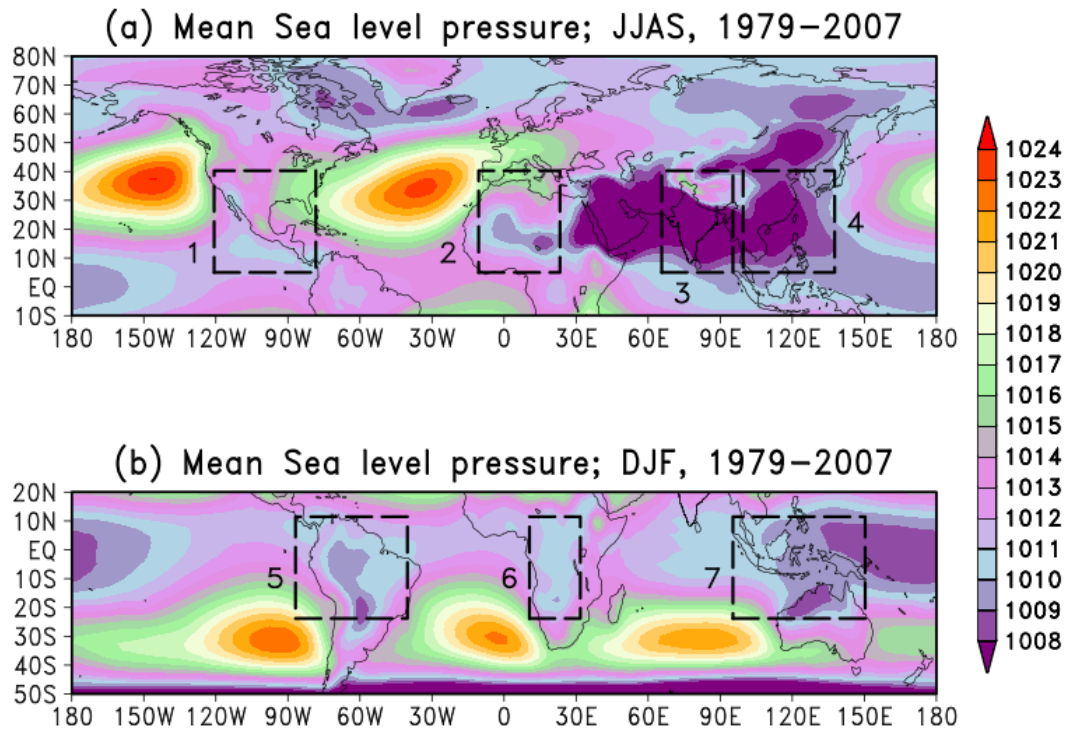


Fig. 1.6: (a) The black dashed line box denotes 1. North American, 2. African (northern western parts of the continent) 3. Indian and 4. East Asian monsoonal regions in the northern hemisphere and the shading represents the climatology mean sea level pressure of the JJAS season for the 1979-2007 period. (b) The black dashed line box denotes the 5. South American, 6. African (middle and southern African countries), and 7. Australian monsoonal regions in the southern hemisphere, and the shading represents the DJF season climatology mean sea level pressure for the 1979-2007 period.

1.4.1 Intraseasonal variability of North American summer monsoon

Importantly, the NAM exhibits substantial interannual variability, which is also connected with intraseasonal variability. For example, a good (poor) monsoon is often associated with an

early (late) start (Douglas and Englehart 1996; Higgins et al. 1999). The intraseasonal variability (IV) of the NAM is also of great importance, as New Mexico and the Southwestern states of the United States of America (USA) still depend on the summer rainfall and as well as human basic needs.

Douglas et al. (1993) showed that the onset of the Mexican monsoon, which is part of the NAM is characterized by heavy rainfall over southern Mexico, which quickly spreads northward along the western slopes of the Sierra Madre Occidental (hereafter SMO) and into Arizona and New Mexico by early July. Higgins et al. (1996) and Wallace (1975) have shown that there are increases in the amplitude of the diurnal precipitation cycles during the summer monsoon. Mo & Berbery (2004) shown the inverse relationship between the low-level jet (LLJ) over Great Plains and NAM, i.e., a strong jet with diurnal variations during summer between the mountains and the Great Plains at 925 hPa (e.g., Bonner 1968; Higgins et al. 1997). The NAM is accompanied by a decrease in midlatitude synoptic-scale transient activity over the adjacent United States of America (USA) and northern Mexico as the extratropical storm track weakens and migrates poleward close to the Canadian border by late June (Whittaker and Horn 1981; Parker et al. 1989).

In the case of NAM, the IV is associated with various factors such as a weak monsoon circulation, blocking high, intrusion of LLJ's, and moisture subsidence in the monsoon basic flow. The NAMS fully develops during the mature phase of boreal summer, and can be related to the seasonal evolution of the precipitation regime. Surges of maritime tropical air northward over the Gulf of California are linked to active and break periods of the monsoon rains over the deserts of Arizona and California (Hales 1972). Among others, Bryson and Lowry (1955), Carleton (1986), and Cavazos et al. (2002) shows the northward shifts in the subtropical ridge axis incline to promote active monsoon convection (bursts) with southward shifts inducing drying (break) periods.

Mullen et al. (1998) suggest a dominant 12–18 day oscillation in the precipitation over South eastern Arizona, which was later ascertained by Cavazos (2002), who, through a detailed analysis, also identified a secondary peak at 40–60 days in the region. Mo (2000) identified a 22-day mode in the summer rainfall over New Mexico and Arizona. Higgins and Shi (2001) define the NAM rainfall as the rainfall over the 5°–35°N, 125°–80°W region. They claim that a major portion of the intraseasonal variability in this area is due to a 30–60 day signal connected to the Madden–Julian oscillation. Similarly, used wavelet analysis to study monsoon rainfall in Arizona and New Mexico. While a wavelet analysis on the monsoonal rainfall in 1990 over southeast Arizona, which was above average, isolated strong periodicities in the 10–20 and 20–40 day ranges, but no such dominant periodic variations were observed for the monsoonal rains in 1993 over the region (Hall-McKim et al., 2002)

1.4.2 Intraseasonal variability of South American summer monsoon

The South American monsoon system (SAM) is distinguished by the seasonal change in precipitation and moisture caused by alterations in the trade winds, and associated processes such as surface pressure, convection, thermodynamic instability, etc. The SAM summer seasonal cycle is essentially due to a variance in heating between South American land mass and the Atlantic Ocean, analogous to the Indian monsoon scenario. In the pre-monsoon season, the insolation amplifies the diabatic heating over the central-western tropical South America, and the ensuing thermal gradient between the land and ocean facilitate the setup of the monsoonal circulations, enhances the moisture advection across the equator by trade winds towards the continent centre. Just as the ISM, the diabatic processes trigger convective instabilities. A combination of large scale and small scale processes make sure that the monsoon persists throughout the wet season (Fu et al. 1999; Fu and Li 2004, Fisch et al. 2004).

1000 The South Atlantic Convergence Zone (SACZ) is a critical feature of the SAM during the wet
1001 season. It is distinguished by a quasi-permanent cloud band, oriented north-west to south-east,
1002 and associated with low level wind and moisture convergence (Kodama 1992; Carvalho et al.
1003 2002; 2004). It has an impact on densely populated areas throughout tropical and subtropical South
1004 America. Heavy rain begins in late August in north western South America and moves south
1005 eastward until it reaches the Brazilian highlands. The wet season in the Amazon core peaks during
1006 austral summer (*December–February*).

1007 Other than interannual, decadal time scales, SAM exhibits persistent “active” and “break”
1008 periods/events in the intensity of rainfall, commonly known as Intraseasonal oscillations. Deep
1009 convection over the Amazon begins to weaken in early March, and the dry season lasts for the
1010 majority of the austral winter (Horel et al. 1989; Jones 1990). The SAM is a feature that has become
1011 increasingly popular to describe the strong summertime convective activity, intense precipitation,
1012 and large-scale atmospheric circulation features. The study of Zhou and Lau (1998) reviewed the
1013 early definitions of monsoon systems, originally formulated to account for reversals in the large-
1014 scale circulation that are driven by differential heating between landmasses and oceans. In
1015 addition, the analysis by Zhou and Lau (1998), which was based on monthly averaged data, indeed
1016 demonstrated that the summer season in South America contains the main ingredients to be
1017 characterized as a monsoon system. Zhou and Lau (1998) also suggest that the easterly winds
1018 prevail in the tropical Atlantic and eastern South America throughout the entire year, and seasonal
1019 reversal of surface winds is not immediately apparent. However, when the annual mean is removed
1020 from summer (January) and winter (July) composites of surface winds, the characteristic reversal
1021 in anomalous low-level circulation becomes evident (see Zhou and Lau 1998, their Fig. 9 and
1022 Grimm et al., 2021). In the context of monsoon systems, an important characteristic observed in

the Asian–Australian region relates to the persistent “active” and inactive (or “break”) periods in the intensity of rainfall amounts. In this respect, tropical intraseasonal oscillations have been found to strongly modulate these “active” and “break” periods of the Asian–Australian monsoon system (Vernekar et al. 1993). Thus, the Madden–Julian oscillation (MJO; Madden and Julian 1994) is key in the occurrence of active and break phases of the Asian–Australian monsoon. Paegle et al., (2020) suggest that IV of rainfall over South America is examined using singular spectrum analysis. The dipole convection pattern with centres of action over the SACZ and the subtropical plains is modulated by modes of different timescales. Both oscillatory modes with periods of 36–40 days (mode 40) and 22–28 days (mode 22) influence convection over the SACZ with the faster mode (mode 22) leading the variability over the subtropical plains.

In the recent Large-Scale Biosphere Atmosphere (LBA) experiment of the intensive field campaign in January–March 1999, some limited observational evidence of variations in the structure of convection and different large-scale circulation regimes was gathered. This experiment took place in the Brazilian state of Rondônia over the southwest part of the Amazon as part of the Wet Season Atmospheric Mesoscale Campaign (WETAMC) and LBA Tropical Rainfall Measuring Mission (TRMM) validation experiment (Silva Dias et al. 2000). Periods of easterly and westerly 850-hPa winds in several Rondônia locations appear to be associated with different physical and structural characteristics of mesoscale convective systems (Cifelli et al. 2002; Petersen et al. 2001; Carvalho et al. 2002).

The MJO is a key tropical intraseasonal process, which controls and varies convection and winds over the tropical and sub-tropical regions of South America (Carvalho et al. 2004; Jones and Carvalho 2002; Liebmann et al. 2004a & b). Circulation Changes due to intraseasonal variations, notwithstanding whether they are due to the MJO or otherwise, influence the moisture transport in

the monsoon region (Herdies et al. 2002; Carvalho et al. 2010b). Over central Brazil, active phases of the SAM are seen to co-occur with westerly wind anomalies, and break phases with easterly wind anomalies (Jones and Carvalho 2002). Such changes in circulation are noted in association with changes in the vertical distribution of moist static stability, which changes the 24-hour cycle of convection, and thereby, the characteristics of mesoscale convective systems (Petersen et al. 2002; Cifelli et al. 2002; Rickenbach et al. 2002; Carvalho et al. 2002b). The synoptic scale activity is associated with the active and break phases of the SAM regime, as well as variations in circulation and thermodynamic properties. Propagating frontal systems facilitate convection in the SACZ during active phase over central and eastern South America, (Garreaud 2000; Rickenbach 2002; Vera et al. 2006; Marengo et al. 2010; Cunningham and Cavalcanti 2006).

With all of these similarities in the three summer monsoons, I discuss the research objectives and scope in the following sub-chapter.

1.5 Objectives and Scope of the study

1.5.1 Objectives

From the review of literature carried out in the previous sections, it is clear that there are certain gaps in understanding the midlatitude interactions with the tropics during monsoons in India as well as Americas from the context of monsoonal intraseasonal variations. and in. Based on these gaps, I have framed four objectives. These are as follows:

1. To Delineate dynamic interactions between mid-latitudes and ISM from the context of the initiation and development of a break phase over the Indian region, through analysis of multiple cases.

2. Examine if the mid-latitude-tropical interactions are also seen for the North and South American summer monsoons as well.
3. Diagnosing the dynamical mechanisms behind the revival of ISM after a typical break phase through analysis of multiple cases
4. To ascertain dynamical mechanism for revival of active conditions of ISM after a break phase using WRF model.

1.5.2 Scope of the present study

The next section describes the datasets used in this study and the methodology adopted in this study is discussed. In chapters 3 & 4, a detailed analysis establishes a similar evolution of the breaks in the ISM, NAM, and SAM, and proposes a uniform mechanism based on the QRA and blocking that causes these breaks. In chapter 5, I propose a mechanism for the revival of summer monsoon rainfall after a break phase over the Indian region. In chapter 6, using the WRF regional model, we diagnose the mechanism proposed for the revival of active conditions after a break phase during ISM. In the final chapter 7, I discuss the conclusion of this research and the future scope of the study.

Chapter 2

Data & Methodology

Outline of the chapter

In this chapter, I discuss the source of the datasets used for the analysis. I also discuss the methodology followed to analyse the genesis of the Blocking high, the sustenance of the break phase in all three monsoonal regions and the mechanism for the revival of active conditions after a break phase during Indian summer monsoon. In addition, I briefly discuss about the Weather Research & Forecasting model and its importance.

1109 **2.1 Data used**

1110 **2.1.1 Observational data**

1111 I use gridded daily rainfall data (Pai et al., 2014) of $0.25^\circ \times 0.25^\circ$ resolution, derived from the
1112 observations of the India Meteorological Department (IMD) to observe the rainfall over the Indian
1113 during break situations for the 1979-2007 period. IMD (Cyclone eAtlas- IMD) data is also used to
1114 identify the dates of formation of Lows at the Head of the Bay of Bengal during the above period.

1115 In addition, I have also used Daily gridded precipitation data derived from the Global Precipitation
1116 Climatology Centre (GPCC) Full Data Reanalysis V.7 version (Schamm et al., 2016) to see the
1117 spatial distribution of rainfall over North American and South American regions during summer
1118 monsoon. This product is developed using weather stations data on a regular grid with a spatial
1119 resolution of 1.0° latitude $\times$ 1.0° longitude. This dataset is based on data provided by NCEP-PSL.

1120 **2.1.2 Reanalysis data**

1121 I use reanalysed daily data sets of globally gridded Zonal wind (U), Meridional wind (V) and
1122 Geopotential (Φ) from the NCEP-DOE Reanalysis II (Kanamitsu et al., 2002) of 2.5° latitude $\times$
1123 2.5° longitude resolution for the period 1979-2007. NCEP-II reanalysis datasets are developed
1124 using available large data from Physical Sciences Laboratory, satellite data from 1979 to the
1125 present and uses an updated forecast model, updated data assimilation system, improved diagnostic
1126 outputs.

1127

1128

1129 **Table 2.1:** A list of source/weblinks of datasets used in the present study.

Sr. No.	Dataset	Source	Weblink
1	Rainfall	IMD	http://www.imd.gov.in/
2	Rainfall	GPCC	https://www.esrl.noaa.gov/psd/data/gridded/
3	Zonal wind, Meridional wind, Sea level pressure, Geopotential height	NCEP-II	https://www.esrl.noaa.gov/
4	Lows	RMC, IMD	http://14.139.191.203/Login.aspx ; http://www.imdchennai.gov.in/

1130

1131 **2.2 Dates of break periods, and study region**

1132 Rajeevan and co-authors (2010) identified active and break events of ISM, and are defined as
 1133 periods during the peak monsoon months of July and August in which the normalized anomaly of
 1134 the rainfall over the core monsoon region (CMR) (18° N-28° N & 65° E- 88° E), exceeds 1 or is

1135 less than -1.0 , respectively, provided the criterion is satisfied for at least three consecutive days.
 1136 They have shown the CMR is almost the same as that of the average all-India rainfall, and also
 1137 shown that the interannual variation of the all-India summer monsoon rainfall is highly correlated
 1138 with a correlation coefficient of 0.91 with that of the summer monsoon rainfall over the CMR and
 1139 suggesting that it is a critical region for the interannual variation.

1140 Following this identification method by Rajeevan et al. (2010), 41 break events have been
 1141 identified through the 1979-2007 period. Various datasets over these breaks are used to generate
 1142 the weather and climate statistics for the break phases during the summer monsoon months over
 1143 India as tabulated below (Table 2.2).

1144 **Table 2.2:** A list of break days based on Rajeevan et al., 2010 for the period 1979-2007 (J-July &
 1145 A-August)

Sr. No.	Break Monsoon days (< 3mm/day)	Year
1	2–6J, 14–29A	1979
2	17–20J, 13–15A	1980
3	24–27A	1981
4	1–8J	1982
5	23–25A	1983
6	27–29J	1984
7	23–25A	1985

8	22–31A	1986
9	23–25J, 30J–4A, 8–13A, 16–18A	1987
10	14–17A	1988
11	18–20J, 30J–3A	1989
12	-	1990
13	-	1991
14	4–11J	1992
15	20–23J, 7–13A, 22–28A	1993
16	-	1994
17	3–7J, 11–16A	1995
18	10–12A	1996
19	11–15J, 9–14A	1997
20	20–26J, 16–21A	1998
21	1–5J, 12–16A, 22–25A	1999
22	1–9A	2000
23	31J–2A, 26–30A	2001

25	-	2003
26	10–13J, 19–21J, 26–31A	2004
27	7–14A, 24–31A	2005
28	-	2006
29	18–22J, 15–17A	2007

1146

1147 Mo (2000) has shown the 7-day running mean precipitation (mm/day) averaged over Arizona and
 1148 New Mexico, i.e., 107.5°–112.5°W, 32°–36°N based on daily precipitation analysis by Higgins et
 1149 al. (1996). Based on this study, we have chosen (from their Fig.6) three cases of North American
 1150 break monsoon (Table 2).

1151 **Table 2.3:** List of North American break monsoon days

Sr. No.	Break Monsoon days (< 3mm/day)	Year
1	17-25 July	1981
2	25 July -02 August	1995
3	16-24 July	2000

1152

1153 Gan et al. (2004) have shown the daily precipitation (mm/day) of west-central Brazil (primary
 1154 region of South American monsoon), i.e., 10°–20° S, 60°–50° W, during month of January for the

1155 1979-1996 period. Following Gan et al. (2004), six break spells are chosen (from their Fig. 19)
 1156 using the daily rainfall (Table 3) with less than 3mm rainfall of the west-central Brazil region.

1157 **Table 2.4:** List of South American break monsoon days

Sr. No.	Break Monsoon days (< 3mm/day)	Year
1	22-24 Jan	1980
2	05-10 Jan	1981
3	16-19 Jan	1988
4	21-26 Jan	1990
5	02-05 Jan	1993
6	25-27 Jan	1996

1158

1159 **2.3 Methodology applied**

1160 Atmospheric waves transport moisture, energy and momentum from their source. Atmospheric
 1161 waves extending large spatial extent are called planetary waves, and are a type of inertial waves
 1162 naturally occurring in rotating fluids. These waves are associated with pressure systems and the
 1163 Jet streams. Planetary waves play a major role in bringing the atmosphere to balance by
 1164 transferring heat, moisture and momentum from equator towards the poles and vice-versa. For
 1165 example, it is well known (e.g., Palmen 1951) that planetary waves in the subtropical westerly Jet
 1166 streams play a primary role in the maintenance of weather in mid-latitudes, and therefore seasonal
 1167 climate as well.

1168 Hide (1953), based on his laboratory experiments noted periodic changes in the shape of
1169 propagating of planetary atmospheric waves. He named this kind of flow with regular periodic
1170 change in the shape as a vacillating flow. Lorenz (1963) classified these waves into four categories
1171 based on their progression and shape as a) a steady symmetric regime, b) a steady-wave regime,
1172 c) a vacillating wave regime and d) an irregular wave regime. For example, Hadley circulation can
1173 be seen as a steady symmetric regime; and a steady-wave regime & a vacillating wave regime
1174 together form a Rossby regime. A Rossby regime is a group of waves such as steady waves as
1175 well as steady waves with periodical change in the shape during progression. For example, SWJ
1176 is a typical example of the vacillating flow in a barotropic atmosphere. The westerly waves
1177 propagate from west to east and are associated with trough and ridges due to various factors such
1178 as, for example, earth's rotation, thermal gradient between equator and poles, etc.

1179 Later, Charney and Devore (1979) found three equilibrium states for orographically-forced Rossby
1180 waves in a simple barotropic channel model, of which two were stable: a high index (resonant)
1181 type state of blocking with strong waves, and a low index (non-resonant) zonal state. Most of the
1182 subsequent observational studies identified quasi-stationary resonant waves in winter. In this
1183 context, the subtropical westerly Jetstream undergo meridional changes along with the zonal
1184 changes, such as low-index flow (LIF) and high-index flow (HIF) (Charney and DeVore 1979).
1185 During LIF, the jet is weakened and tends to develop elongated troughs and ridges. In the case of
1186 HIF, however, the jet is stronger than normal and narrower. As the jet weakens, the associated
1187 troughs and ridges have considerable effect on the nearby regions. During southwest monsoon of
1188 India, middle latitude westerlies exercise considerable influence on the monsoon weather over
1189 northern India, essentially during breaks (for relevant studies, see subsection 1.3.2) In this context,

1190 it will be interesting to examine the relevance of vacillating flows in facilitating break monsoon
1191 conditions.

1192 **2.3.1 Dynamical analysis carried out**

1193 **2.3.1.1 Eliassen–Palm fluxes**

1194 From the view point of understanding the midlatitude interactions with the tropics at the upper
1195 level, it's important to observe the changes in the planetary baroclinic waves. To study the
1196 maintenance of blocking high episodes, computation of the Eliassen–Palm (1961) (EP) fluxes will
1197 be very useful as they represent the effect of transient and stationary eddy fluxes on the zonal-
1198 mean circulation. The properties and application of EP fluxes have been discussed by several other
1199 authors as well (Andrews and McIntyre, 1976a; 1978a; Edmon et al., 1980; Tung, 1986) in detail
1200 in previous studies.

1201 We summarize them here in the Quasi-Geostrophic ambit for simplicity. Briefly, the EP flux is a
1202 useful diagnostic of the propagation of waves (both magnitude and direction) in the meridional
1203 plane. In the case of small amplitude waves, it represents the flux of wave activity, and its
1204 divergence is zero for steady conservative waves. EP fluxes are also an important diagnostic for
1205 wave propagation, so we try to figure out the Rossby wave propagation and its importance in the
1206 formation of blocking high and its maintenance. Meridionally propagating eddies transport
1207 momentum towards their generation region (Held, 1975; 2000).

1208 The EP flux in pressure coordinates on the sphere is given, following Andrews et al., (1983), by

$$1209 \quad F_{(\phi, P)} = r_o \cos(\phi) [-(\overline{u'v'}), f(\overline{v'\phi})/\theta_p] \quad \text{-----} (2.1)$$

1210 Where ϕ is the latitude, P is the pressure, θ is the potential temperature, r_0 is the radius of the earth.
 1211 Bars ($\overline{\quad}$) and primes ($\prime$) denote zonal means and deviations respectively, and u , v are zonal and
 1212 meridional velocities.

1213 **2.3.1.2 Energetics computations**

1214 Lorenz (1955) and Oort et al., (1962) defined the importance and consequences of the growth of
 1215 perturbations through the expense of the mean available potential energy and the mean kinetic
 1216 energy in a baroclinic and barotropic atmosphere respectively. They also have given mathematical
 1217 expressions for the spatial and temporal variations of the energetics. In our case, during break
 1218 phase of ISM, quantifying the energy exchange between the SWJ and TEJ (e.g., Rao 1971) is
 1219 important.

1220 The rate of conversion of mean kinetic energy to eddy kinetic energy is expressed as

1221

$$1222 \quad C(\overline{K}, K') = \int U \frac{\partial}{\partial y} \overline{v' u'} dm \text{-----} (2.2)$$

1223 Where m is the mass, $\overline{K}$ is the Mean kinetic energy (Joule/sec), K' is the Eddy kinetic energy, U
 1224 is the zonal wind (m/s) and V is the meridional wind (m/s). The u' and v' have been obtained as
 1225 the daily anomalies from the zonal mean of the U averaged over 20° E and 120° E, and v' , similarly
 1226 from the zonal mean of V . Equation 2 means that if there is divergence (convergence) of eddy
 1227 momentum transport in region of westerlies, $\overline{K}$ gets converted into K' (K' gets converted into $\overline{K}$),
 1228 that is, the disturbance is barotropically unstable (stable).

1229 The eddy momentum flux $\overline{u'v'}$ controls the structure of the mean zonal surface wind and of
 1230 meridional cells. This flux is significant for global circulation because it fulfils the global
 1231 momentum balance equation. The equatorial region's zonal momentum is transported to the
 1232 midlatitudes, where a flux convergence zone can be found. This is due to the global wind pattern,
 1233 which consists of easterlies in the equatorial region and westerlies in the midlatitudes. Whereas
 1234 the atmospheric waves are responsible for the transfer of zonal momentum.

1235 Hence, the structure of the eddy momentum flux is fundamental to the mean state of Earth's
 1236 atmosphere.

1237 Assuming the stream function ψ as a function of latitude and longitude,

1238 Given,

1239 $\psi = R_{\psi}(y) \sin[Kx + \delta(y)]$

1240 leads to

1241 $\frac{\partial \psi}{\partial x} = v = KR \cos(Kx + \delta)$

1242 Therefore, $u = -\frac{\partial \psi}{\partial y} = -\left[\frac{\partial R}{\partial y} \sin(Kx + \delta) + \frac{\partial \delta}{\partial y} R \cos(Kx + \delta)\right]$

1243 and $vu = -KR \frac{\partial R}{\partial y} \sin(Kx + \delta) \cos(Kx + \delta) - \frac{\partial \delta}{\partial y} KR^2 \cos^2(Kx + \delta)$

1244 In the above, $\overline{()}$ represents the zonal mean. From the above equation, the mean zonal momentum
 1245 can be expressed as

1246
$$\overline{(vu)} = \frac{\overline{-v^2}}{K} \frac{\partial \delta}{\partial y} \text{-----}(2.3)$$

1247 1. If $\frac{\partial \delta}{\partial y}$ is negative, the wave tilts Northeast to Southwest (NE-SW).

1248 2. if $\frac{\partial \delta}{\partial y}$ is positive, the wave tilts Northwest to South east (NW-SE), in the first case (1) the eddy
 1249 momentum transport is northward and while in the second case (2) the transport of eddy
 1250 momentum is towards southward.

1251 **2.3.1.3 Condition for unstable waves**

1252 My research problem is closely connected to barotropic stability theory, and previously a few
 1253 authors have discussed different mathematical methods of which three important methods are
 1254 detailed below.

- 1255 • The eigen-value problem of linearized equation: Kuo (1951), by using linearized vorticity
 1256 equation, defined the distinct profile conditions of zonal mean flow that would decide
 1257 whether a disturbance would grow or dampen. Kuo also discussed the implications of such
 1258 growing or weakening perturbations to the zonal mean flow.
- 1259 • Initial value problem of non-linear equation: Platzman (1952), Kuo (1953) and Syono &
 1260 Aihara (1957) treat this problem as an initial value problem. This is done by prescribing
 1261 certain initial conditions for the distributions of zonal mean flow and perturbations. Then,
 1262 the higher order time derivatives of their kinetic energies are computed.

1263 • A method with statistical assumption: While Lorenz (1953) also acknowledges that this is
1264 an initial value problem, he confines to prescribing only a few initial statistics of the flow,
1265 rather than the whole profile of the flow pattern.

1266 In steps of Kuo (1953), Syono & Aihara (1957) and Rao (1968; unpublished thesis) found an
1267 analytical solution of the growing waves in a barotropic atmosphere. Rao 1971 introduced the
1268 concept of neutral wavelength 'L', which isolates the stable shorter waves and unstable longer
1269 waves in to two distinguished groups. It is to be noted that the final output of the solution is not
1270 associated with the contribution of earth's rotation. The detailed derivation is found in appendix I.
1271 From this theory, it is proposed that the expression for the time change of perturbation kinetic
1272 energy at the initial time is given by Rao (1968) as

$$1273 \quad \left(\frac{\partial^2 K_r}{\partial t^2} \right)_o = 0 \quad \text{when } L=2D/\sqrt{3} \quad (2.4)$$

$$1274 \quad > 0 \quad \text{when } L>2D/\sqrt{3}$$

$$1275 \quad < 0 \quad \text{when } L<2D/\sqrt{3}$$

1276 K_r is perturbation kinetic energy and the left hand side of the above equation is a measure of
1277 barotropic instability. Indeed, the associated waves longer than L (wavelength) become unstable
1278 and below L are stable (Starr & White, 1954; Aihara, 1959). Whereas D/2 is the zonal width
1279 between subtropical westerly jet and tropical easterly jet.

1280

1281

1282

1283 **2.3.1.4 Condition for barotropic instability**

1284 In our analysis, we use the criterion by Kuo (1949), which states that, for barotropic instability to
1285 happen at a location, the meridional gradient of the absolute vorticity has to be either maximum or
1286 minimum.

$$1287 \quad \frac{d}{dy} \left(\frac{-d\bar{U}}{dy} + f \right) = 0 \quad \frac{d\zeta}{dy} = 0 \quad (2.5)$$

1288 Where $\bar{U}$ is the mean zonal wind, f is the Coriolis force and ζ is the absolute vorticity. We
1289 use the criterion shown in equation (3) to explain the mechanism behind the formation of the after-
1290 break synoptic disturbances over the Indian region and the Bay of Bengal, which reactivate the
1291 Indian summer monsoon.

1292 **2.3.2 Statistical methods**

1293 **2.3.2.1 One sample Student's t test for composite analysis**

1294 In our composite analysis, the statistical significance of the composited signal is assessed by
1295 computation of the 't' statistic from a 2-tailed Student's t-test. This is a standard test for composite
1296 analysis. The 't' is calculated as

$$1297 \quad t = \frac{X - \mu}{\sigma / \sqrt{n}} \quad (2.6)$$

1298 Where X is the sample mean, μ is the population mean, σ is the standard deviation of the population
1299 mean and n is the number of cases.

1300

1301 **2.3.2.2 Student's t test for two unequal variances**

1302 Two-sample T-Test with unequal variance can be applied as the samples are normally distributed,
1303 and the standard deviation of both populations are unknown and assumed to be unequal with
1304 sufficiently large sample size (>30)

1305
$$t_{df} = \frac{(\bar{x}_1 - \bar{x}_2) - (\mu_1 - \mu_2)}{\sqrt{\left(\frac{s_1^2}{n_1} + \frac{s_2^2}{n_2}\right)}} \quad (2.7)$$

1306

1307
$$df = \frac{\left(\frac{s_1^2}{n_1} + \frac{s_2^2}{n_2}\right)^2}{\frac{\left(\frac{s_1^2}{n_1}\right)^2}{n_1 - 1} + \frac{\left(\frac{s_2^2}{n_2}\right)^2}{n_2 - 1}}$$

1308 x_1 = values of the first sample; $\bar{x}$ = mean of the values of the first sample; x_2 =values of the second
1309 sample; $\bar{y}$ =mean of the values of the second sample, and df=degrees of freedom

1310 **2.3.2.3 Pearson's correlation coefficient**

1311 Pearson's correlation is also referred as Pearson's r. It is a measure of linear correlation between
1312 two sets of data, such that the result always has a value between -1 and 1.

1313
$$r = \frac{\sum(x_i - \bar{x})(y_i - \bar{y})}{\sqrt{\sum(x_i - \bar{x})^2 \sum(y_i - \bar{y})^2}} \quad (2.8)$$

1314

1315 r = correlation coefficient; x_i = values of the x-variable in a sample; $\bar{x}$ = mean of the values of the
1316 x-variable; y_i =values of the y-variable in a sample, and $\bar{y}$ =mean of the values of the y-variable

1317 **2.4 A brief introduction to the Weather Research and Forecasting model**

1318 I prefer carry out few sensitivity experiments using a regional climate model in order to validate
1319 the theory proposed in the present study. Here in our case, we use the Weather Research and
1320 Forecasting (WRF) Model for its wide range of applicability in prediction & validation. WRF
1321 model is mesoscale numerical weather prediction system designed for both atmospheric research
1322 and operational forecasting applications. The model is useful for a variety of meteorological
1323 applications at scales ranging from tens of metres to thousands of kilometres. It features two
1324 dynamical cores, a data assimilation system, and the other one is a software architecture supporting
1325 parallel computation and system extensibility.

1326 WRF model is a collaborative output from the National Centre for Atmospheric Research (NCAR),
1327 National Centres for Environmental Prediction (NCEP) and the Earth System Research
1328 Laboratory, the U.S. Air Force, the Naval Research Laboratory, the University of Oklahoma, and
1329 the Federal Aviation Administration (FAA). WRF offers operational forecasting a flexible and
1330 computationally-efficient platform, while reflecting recent advances in physics, numeric, and data
1331 assimilation contributed by developers from the expansive research community.

1332 Importantly, all the sensitivity experiments conducted using WRF model and their discussion is
1333 carried out in the chapter 6.

1334

1335

Chapter 3

Breaks in Indian Summer Monsoon

Outline of the chapter

In this chapter, I carried out a few dynamical and statistical analyses, to understand the interactions between the upper level midlatitude circulation features and the tropical summer monsoon circulation of Indian region. Based on the results from the analyses, I propose a mechanism for the genesis of the Blocking high and how it facilitates a monsoon break phase over the core monsoonal region of the Indian region.

3.1 Introduction

The ISM experiences intraseasonal variability associated with higher and lower rainfall patterns. The lower spells of rainfall are called breaks in monsoons. These are characterized by relatively sparse rainfall over the core monsoon region (CMR) of India, which is bounded by 18° N to 28° N, and 65° E to 88° E. The region is referred to as the core monsoon region because the summer monsoon rainfall in this region is highly correlated with the area-averaged summer monsoon rainfall over the whole of India (correlation coefficient=0.91) (Rajeevan et al., 2010). The rainfall variations in CMR are very important for several aspects such as agricultural production and energy generation. Intraseasonal variations of rainfall are manifested as (a) monsoon breaks of low rainfall, (b) normal (c) or heavy rainfall in the CMR region, which are referred to as active periods. Hence the years with low rainfall in the CMR region associated with prolonged breaks are of vital importance. In this context, we propose to study the dynamics leading to break situations.

From a dynamical perspective, Raman and Rao (1981) suggested that the prolonged breaks causing severe summer monsoon droughts over the Indian sub-continent are due to upper tropospheric blocking high over East India (90° E to 120° E) associated with baroclinic waves. Thus, this blocking high is perceived to be the initiator of the monsoon break. However, Raman and Rao (1981)'s conclusions are based on studying a small number of cases. In this context, I evaluate the potential importance of large-scale circulation that leads to the blocking high and subsequent generation of break monsoon conditions. In this analysis, we find the fundamental characteristics that are suggested to lead to the blocking high, namely, the planetary wave resonance as envisaged by Charney and DeVore (1979).

In this context, the following subchapters 3.2 & 3.3 discusses the results of the conditions that lead to break phase and conditions for the genesis of blocking high using reanalysis data in the Indian monsoonal region. The key findings are summarized in the subchapter 3.4 by proposing a robust method to identify a break phase in the monsoon regions with the interactions of midlatitudes.

3.2 Breaks in the Indian Monsoon Season: Rainfall characteristics and associated synoptic situation

Fig. 3.1 illustrates 200 hPa composite geopotential height (Z_{200}) zonal anomalies for a) prebreak, b) break, and c) after-break periods for the 41 cases of Indian break monsoon (Table 2.1). The hatched areas in Fig. 3.1 indicate composite Z_{200} anomaly significant the 95% confidence using a two tailed student's t-test. The composites of the zonal anomaly of Z_{200} , shown in various panels of Fig. 3.1, indicate that the high pressure region is split into two vortices, as already noted earlier (Siu & Bowman, 2020). Siu & Bowman (2020), who refer to this *quasi-stationary high-pressure zone during the summer season* as the 'Asian monsoon anticyclone', claim that the split structure is due to a combination of various factors such as diabatic heating, its interactions with Rossby waves propagating along with the subtropical jet, and internal dynamics within the anticyclone. Notably, during the break phase, we see this high over India extending to lower levels at the 1000 hPa (figure not shown).

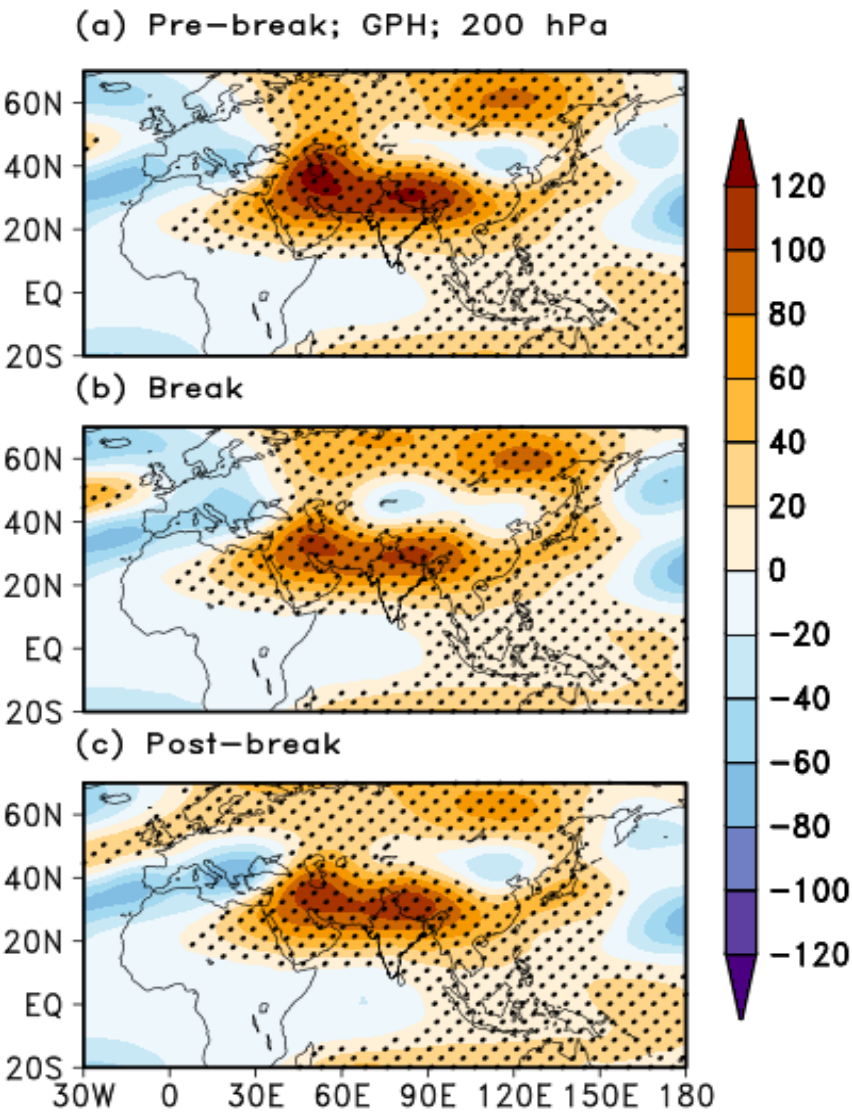
Differences in zonal anomaly of Z_{200} are not immediately perceivable among figures 3.1a, 3.1b and 3.1c. Therefore, to focus on the changes in the upper tropospheric high during the break monsoon situation, we present in Fig. 3.2a the difference obtained by subtracting the composite daily zonal anomalies of Z_{200} for the prebreaks from that over the breaks. We have also subtracted the composite zonal anomalies of Z_{200} during the after-breaks from that over the breaks (Fig. 3.2b). Similar differences have also been generated for the SLP (figures 3.2c & 3.2d). In figures 3.2a and

3.2b, we see negative values of the Z_{200} over northern and central India, statistically significant at 95% confidence levels from a 2-tailed Student's t-test.

Fig. 3.2a suggests, in a relative sense, a substantially weak high and so an increased convergence over the north Indian region during breaks compared to the prebreaks and also relative to the after-breaks (Fig. 3.2b). The positive values of the SLP in figures 3.2c & 3.2d suggest a relative filling up of the surface low, i.e., increase in the SLP. In other words, the various panels in Fig. 3.2 suggest a relative sinking of air from the upper level to lower level over north India during breaks compared to the prebreak and post break conditions. This is because of the differences that are indicative of a relative upper level 200 hPa convergence and a relative low level divergence during the breaks. This is also clear from a similar composite difference of the daily 1000 hPa geopotential height zonal anomaly between monsoon breaks and prebreaks (Fig. 3.3a) and that between monsoon breaks and after-breaks (Fig. 3.3b).

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Figure 3.1: Composite daily zonal anomalies of Geopotential height (m) over the monsoon season of 1979-2007 in the Indian region during (a) prebreak periods, (b) break periods, and (c) after-break periods. The hatching in the Figures 1 a-c indicates the regions where the composite Geopotential height is significantly different from zero at 95% confidence level. Statistical significance has been obtained using a two-tailed one sample Student's t test. Note that

significance test has not been applied to negative values. The shaded region vary between -120 m to 120 m, with an interval of 20 m.

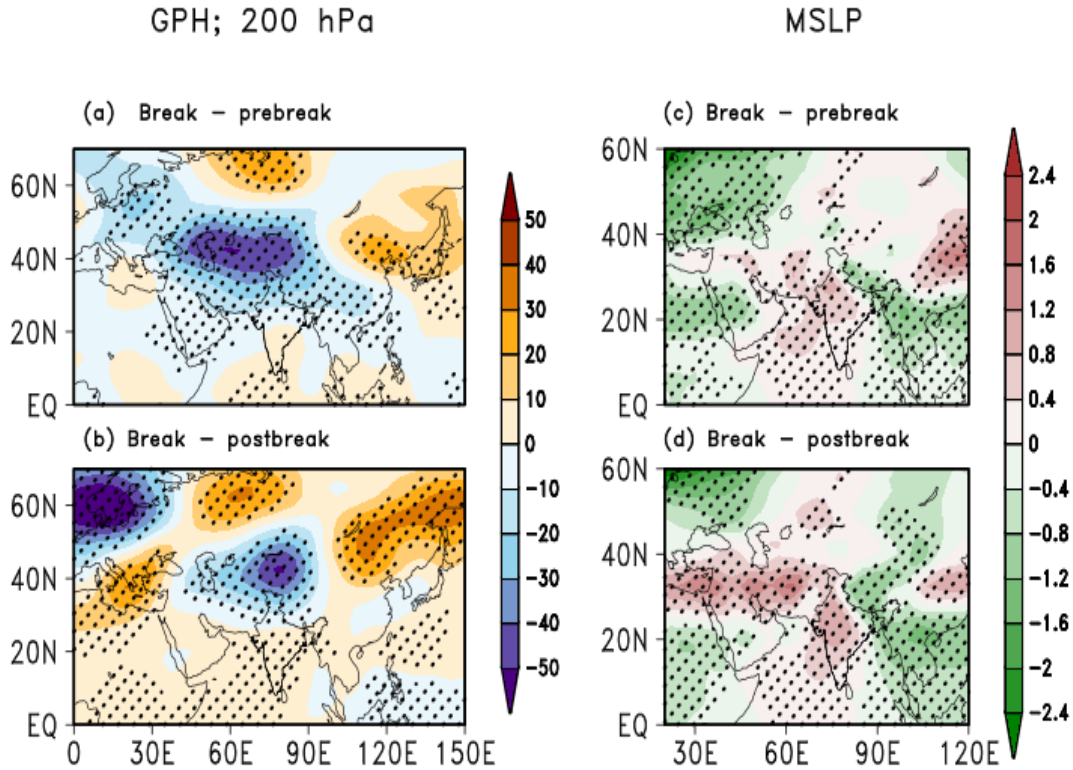
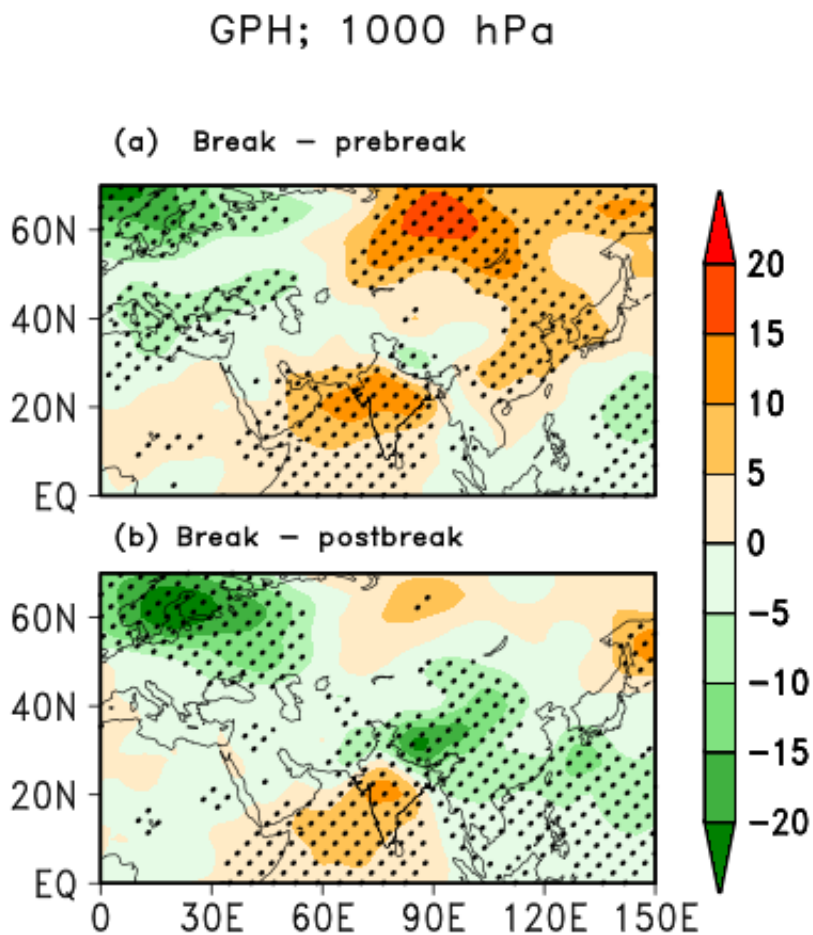


Figure 3.2: Difference between the composite daily zonal anomalies of Geopotential height (m) during the 1979-2007 period over the Indian region between (a) break and prebreak periods (b) break periods and after-break periods. Difference in composites of zonal anomalies of daily mean sea level pressure (hPa) over the Indian region during the 1979-2007 period between (c) during-break and prebreak periods, and (d) during-break and after-break periods. The shaded region in the panels (a) & (b) vary between -50 m to 50 m, with an interval of 10 m and those in panels (c) & (d) vary between -2.4 hPa to 2.4 hPa, with an interval of 0.4 hPa. The hatched regions indicate locations of significant differences, at 95% confidence level, in daily zonal anomalies of Geopotential height in panels (a) & (b), and that of daily mean sea level pressure in panels (c) &

1441 (d). Statistical significance has been obtained using a two-tailed two sample Student's t-test with
 1442 unequal variances.

1443



1444

1445 **Figure 3.3:** Difference between the composite daily zonal anomalies of Geopotential height at
 1446 1000 hPa (m) during the 1979-2007 period over the Indian region between (a) break and prebreak
 1447 periods (b) break periods and after-break periods. The shaded region in panels (a) & (b) vary
 1448 between -20 hPa to 20 hPa, with an interval of 5 hPa. The hatched regions indicate locations of
 1449 significant differences, at 95% confidence level.

In addition to Ramaswamy (1962), Krishnan et al. (2009) also note the intensification of the anomalous westerly troughs was accompanied by retardation in their eastward movement from west central Asia toward the Tibetan plateau for 20 years mean boreal summer season. Krishnan et al. (2009) also suggest that, using upper-level circulation charts during weak monsoon periods indicate zonally asymmetric variations of the Tibetan high under the influence of southward penetrating midlatitude troughs (lows) over west central Asia and the formation of stagnant blocking highs between 90° and 115° E over East Asia (Raman and Rao 1981). According to Raman and Rao (1981) at 200 hPa, a dominant ridge on the North, with a cut off high or blocking high on the southern side of the westerly jet, tends to associate with a break phase over the Indian region. Raman and Rao (1981) show the various stages of blocking high over Southwest Asian region (their Fig. 1), of which the first stage is with an amplified ridge at 200 hPa. This amplification is clearly akin to the increase of amplitude of wave number 7, (in their Fig. 1 we note an approximate distance between two ridges or wave length of about 50 degrees) and this amplification is a sign of resonating affect between the planetary Rossby waves of the westerly jet. Further, we also see this amplification associated with the ridge in the individual case study of ISM of *daily geopotential heights at 200 hPa & 1000 hPa* and rainfall distribution, and are shown in supplementary information (figures 3.4 a-c, 3.5 a-c & 3.6 a-c). The daily area-averaged normalised rainfall anomaly over the core monsoon region during this event was lower than ~3 for at least three of the 10 days over which the long break event occurred. Qualitatively, the other features, such as the anomalous shift of the monsoon trough northward have also been observed during this case (figure not shown).

This dominant ridge remains for about a few days up to two weeks. Also, these authors found a closed “warm” high extending the entire column with an increase of surface pressure by about 20

hPa. The warmth of the high indicates sinking motion. Raman and Rao (1981) also noted a trough at upper levels. From this context, it is a typical case with a few days of exacerbated break conditions. This typical case study provides a glimpse into the blocking high dynamics, which are critical in understanding the characteristics of sub-seasonal variability.

Fig. 3.7 depicts the vertical velocity (ω ; in Pa/s) of the Indian summer monsoon during the prebreak, break, and after-break periods, as well as the difference in it between the break and prebreak periods and that between the break and after-break periods. A raising motion can be seen in the composites of prebreak, break, and after-break in figures 3.7a, 3.7b, and 3.7c, but the magnitudes differ as shown in figures 3.7d & 3.7e. In figures 3.7d & 3.7e, we see a sinking motion (positive ω values extending vertically) at 20° N. The observed positive ω values confirm the sinking motion during breaks, as previously discussed in Fig. 3.2. The changes in the ω (figures 3.7 a-e) are significant at 95% confidence level from a two tailed students t-test.

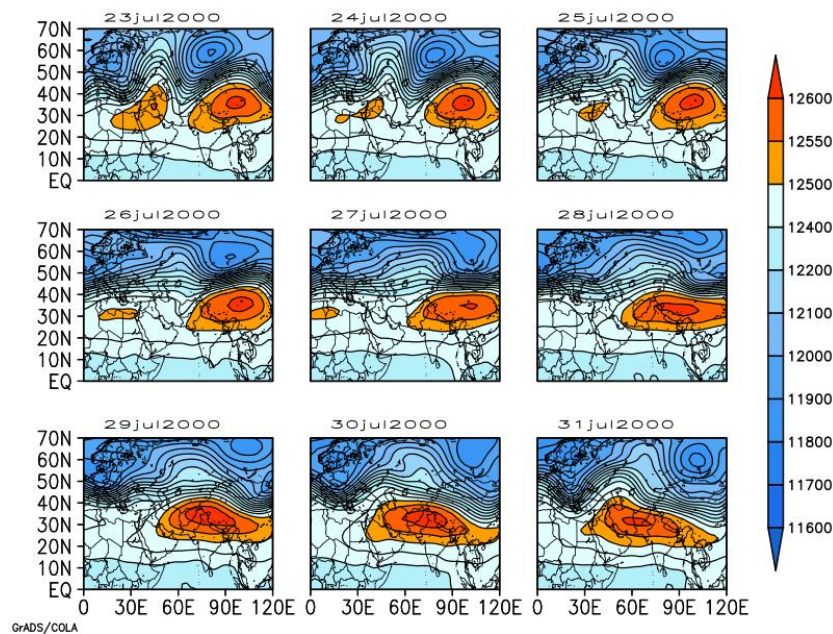


Figure 3.4a: Daily geopotential height distribution at 200 hPa level for a prebreak period (23-31 July 2000) over the Indian region.

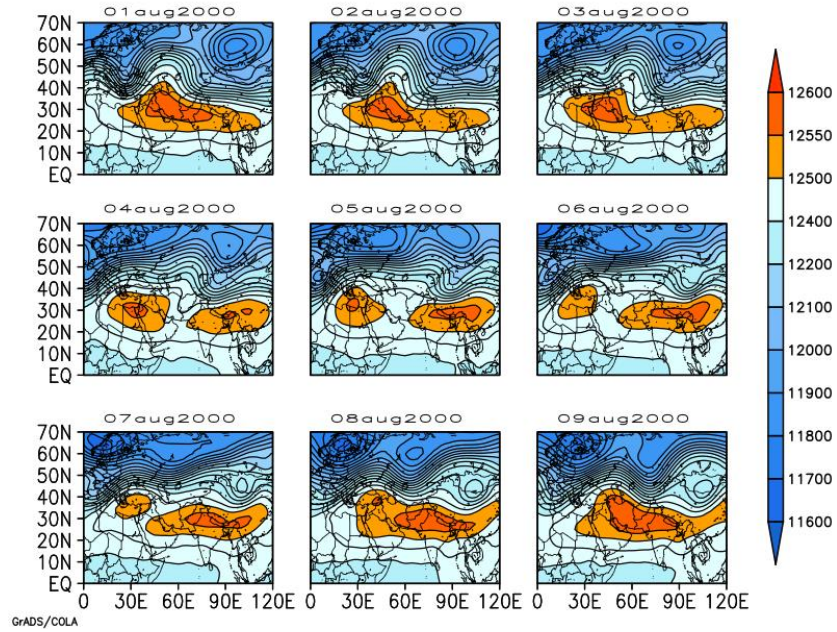


Figure 3.4b: Daily geopotential height distribution at 200 hPa level for a break period (01-09 August 2000) over the Indian region.

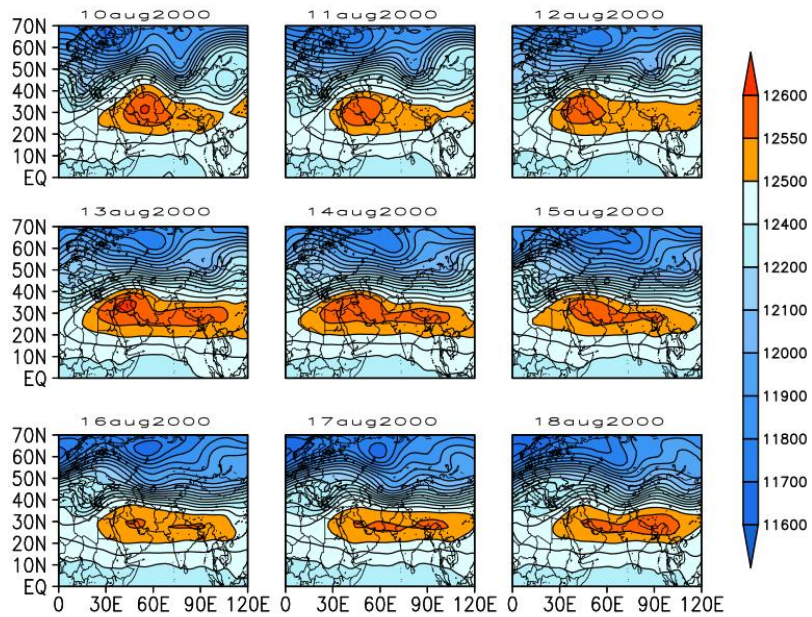


Figure 3.4c: Daily geopotential height distribution at 200 hPa level for an after-break period after (10-18 August 2000) over the Indian region.

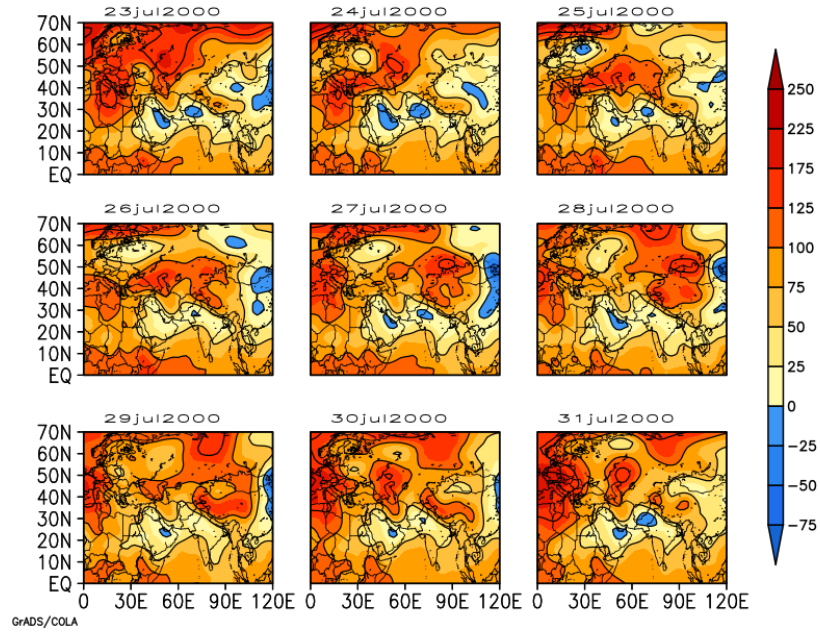


Figure 3.5a: Daily geopotential height distribution at 1000 hPa level for a prebreak period (23-31 July 2000) over the Indian region.

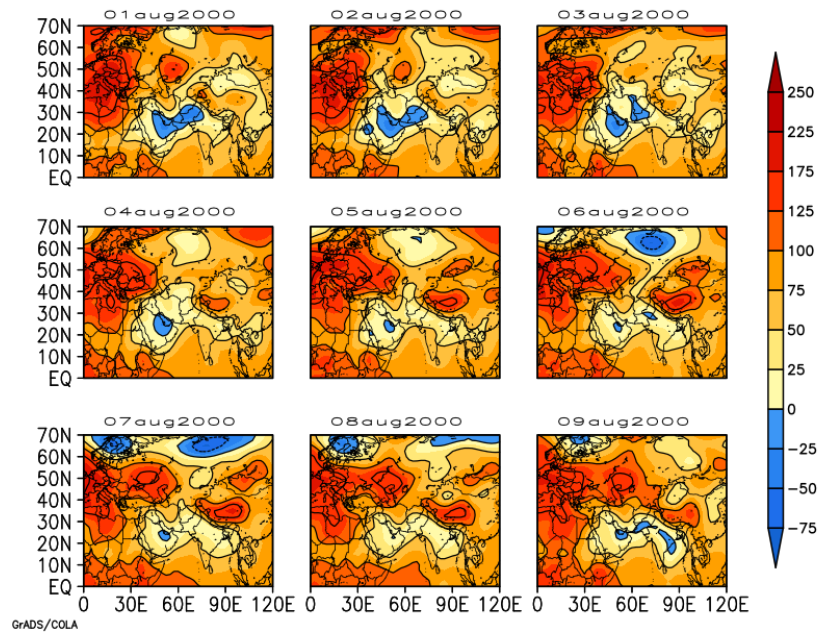


Figure 3.5b: Daily geopotential height distribution at 1000 hPa level for a during-break period (01-09 August 2000) over the Indian region.

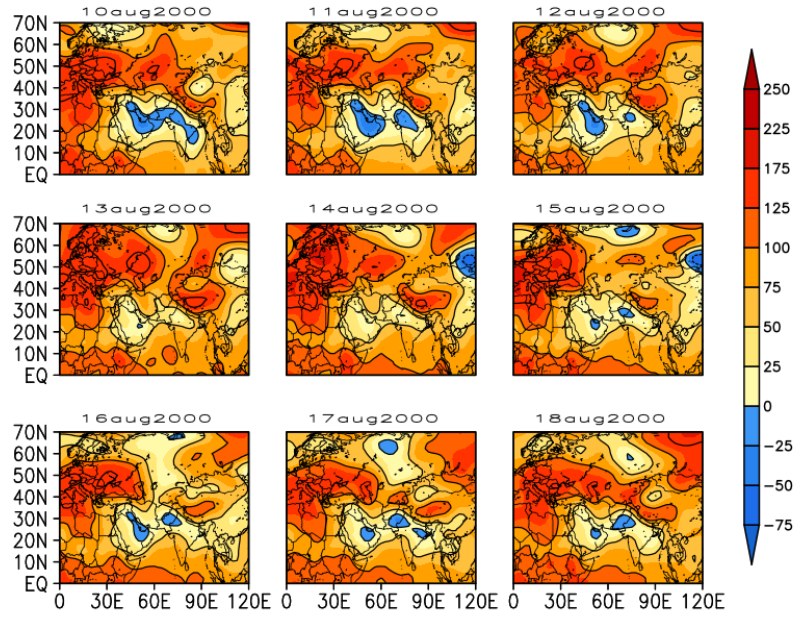


Figure 3.5c: Daily geopotential height distribution at 1000 hPa level for a after-break period (10-18 August 2000) over the Indian region.

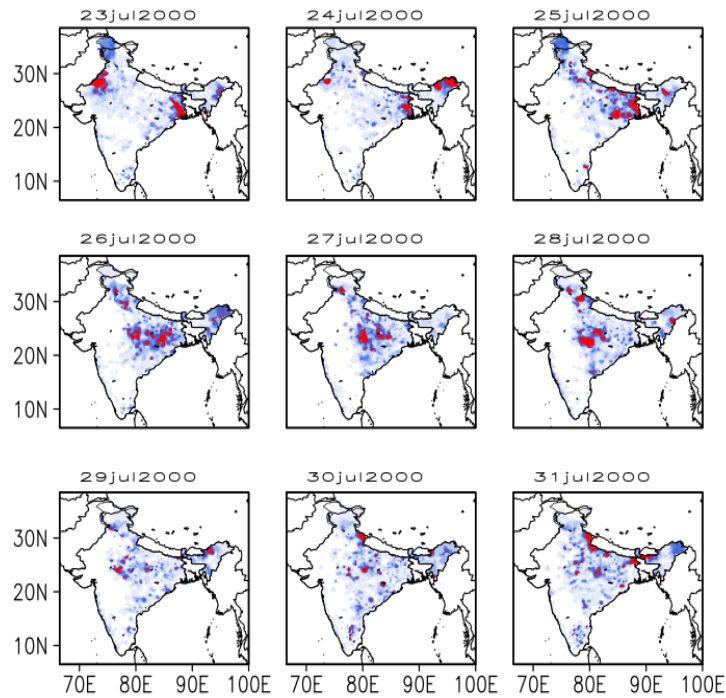


Figure 3.6a: Daily spatial rainfall distribution for a prebreak period (23-31 July 2000) over the Indian region.

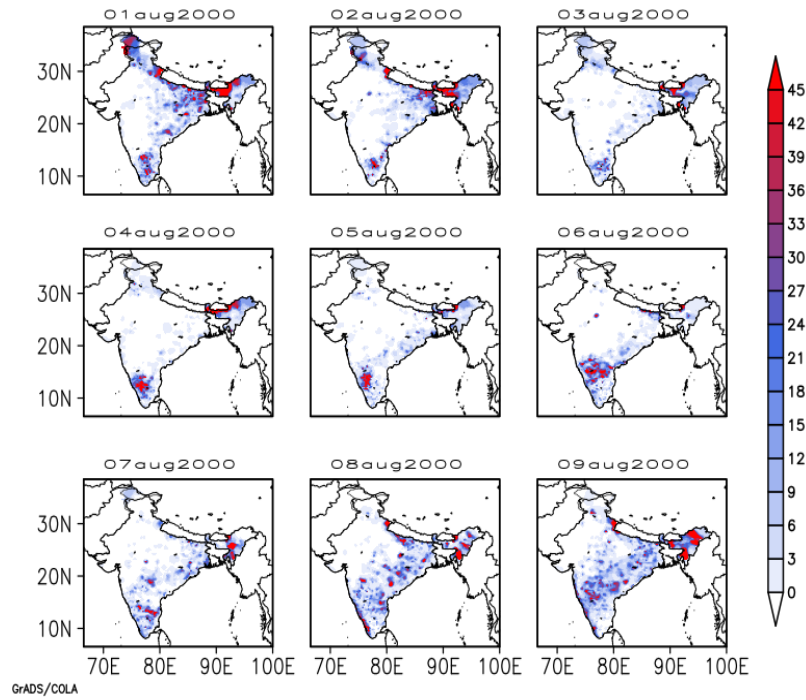


Figure 3.6b: Daily spatial rainfall distribution during (01-09 August 2000) a break period over the Indian region.

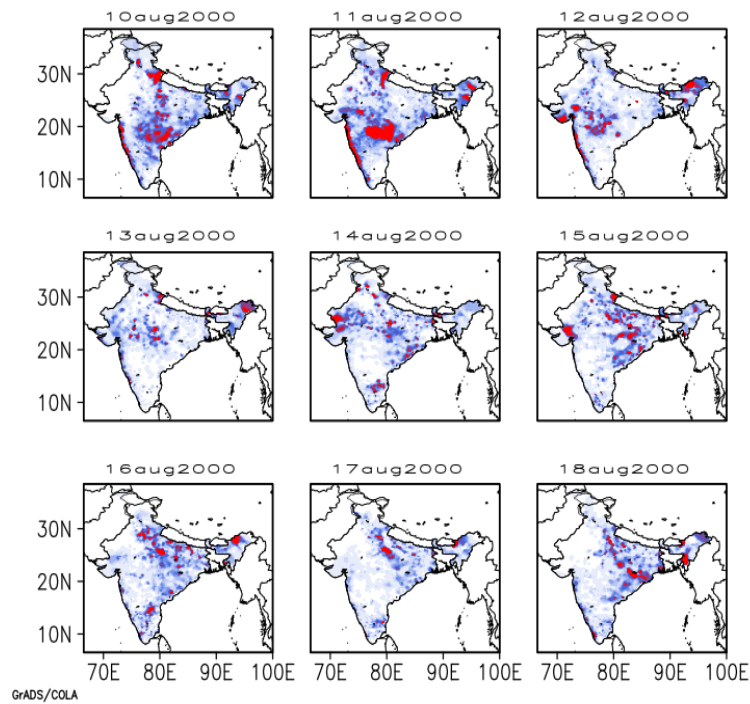


Figure 3.6c: Daily spatial rainfall distribution for the after-break period (10-18 August 2000) over the Indian region.

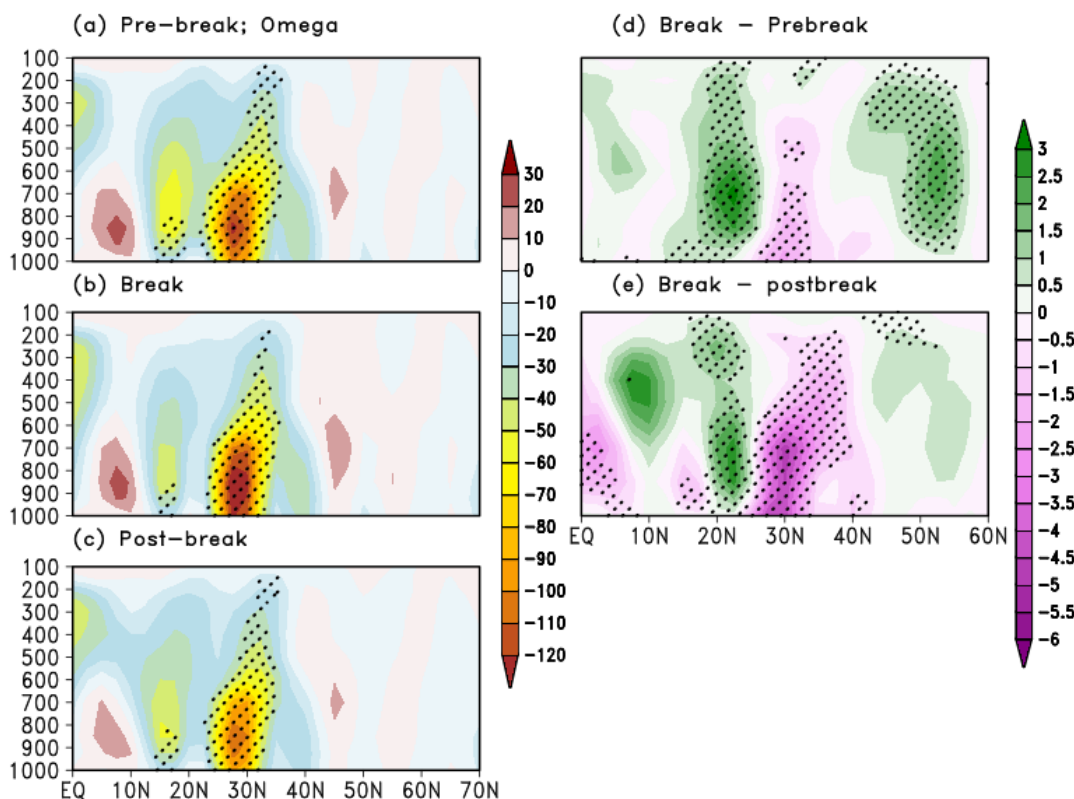


Figure 3.7: Composite vertical wind zonal anomalies ($\times 10^{-3}$) (ω) (Pa/s) over Indian region and longitudinally averaged along 60°E to 100°E during the summer monsoon for the 1979-2007 period, **(a)** prebreak periods **(b)** during-break periods, and **(c)** after-break periods, **(d)** difference between the composite break and composite prebreak periods, and **(e)** difference between the composite break and composite after-break periods. The shaded region in panels (a) (b) & (c) vary between -120 hPa to 30 hPa, with an interval of 10 hPa and those in panels (d) & (e) vary between -6 hPa to 3 hPa, with an interval of 0.5 hPa. The hatched regions indicate composite omega zonal anomalies are significant at 95% confidence from a two-tailed Student's t-test.

1522
1523 Nonetheless, enhanced subsidence should be associated with reduced rainfall., figures 3.8a, 3.8b
1524 & 3.8c show the corresponding composite daily rainfall over the Indian region for the prebreak,
1525 break, and after-break periods, respectively. The rainfall changes, shown in figures 3.8a, 3.8b &
1526 3.8c are significant at 95% confidence level from a two tailed students t-test. Figures 3.8d & 3.8e
1527 show the difference in rainfall between the composites of break & prebreak, and that between the
1528 break & after-break periods, respectively. Figures 3.8d & 3.8e demonstrates that the Indian region
1529 undergoes a decline in rainfall during the breaks as compared to the prebreak, and after-break
1530 composite periods. The low or no rainfall in the CMR of India during the breaks is evident due to
1531 the relative sinking motion (Figures 3.2 & 3.7) and the changes in the panels of figures 3.8d &
1532 3.8e are significant at 95% confidence level using two tailed students t-test.

1533 From the above observations, we notice two significant findings after conducting composite
1534 analyses (Figures 3.1, 3.2, 3.7 & 3.8) and individual study of each case (figures not shown)

- 1535 • The sinking due to upper level (200 hPa) convergence and SLP divergence at lower level.
- 1536 • A blocking high forming due to the increase of amplitude of the wave, provoking “sinking”
1537 and this condition is discussed further below.

1538 Also noted is that, in all the “break” monsoon cases, there is an increase of rainfall near “foot hills
1539 of the Himalayan region”. We note that at the lower levels (figures 3.2c & 3.2d), the monsoon
1540 trough shifts to the foot hills of the Himalayan region during the breaks (e.g., Raghavan, 1973).
1541 Apart from the raising motion in the monsoon trough in this location, we also note that there are
1542 “northerlies (or north westerlies)” near the foothills. These strong north westerlies at Himalayan
1543 foothills are along the monsoon trough, cannot cause high rainfall as explained below (fig. 2).
1544 However, we see a good rainfall across the foot hills of the Himalayan region, and a low or no

1545 rainfall situation at the core monsoon region or most of the Indian region with the sinking
1546 conditions during the break phase. The rising motion at the foothills as explained below

1547
$$w = u \frac{\partial h}{\partial x} + v \frac{\partial h}{\partial y},$$

1548 Here, h is the height of the mountain in the northwest (x) direction, and u is the daily zonal wind.

1549 Here in the first term, u is positive because of the westerlies *during the break*. The $\frac{\partial h}{\partial x}$ is positive

1550 because the trough is on the west side of the foot hills, and the height increases towards east. The

1551 other term is almost zero because along the foot hills for the Northerlies $\frac{\partial h}{\partial y}$ is very small or zero.

1552 Note that figures 3.2c & 3.2d show a strong gradient of SLP zonal anomalies at North of 20° N

1553 and East of 90° E, indicating approximately, a geostrophic westerly zonal wind $u_g = -\frac{1}{f} \frac{\partial p}{\partial y}$, i.e.,

1554 we note a high pressure zone on the left of the Himalayan and central Indian region, low pressure

1555 zone on Tibetan region i.e., right of the Himalayan region (figures 3.2c & 3.2d), and the Himalayan

1556 region height gradient is $\frac{\partial h}{\partial x}$ is strong along the monsoon trough region, indicating a pressure

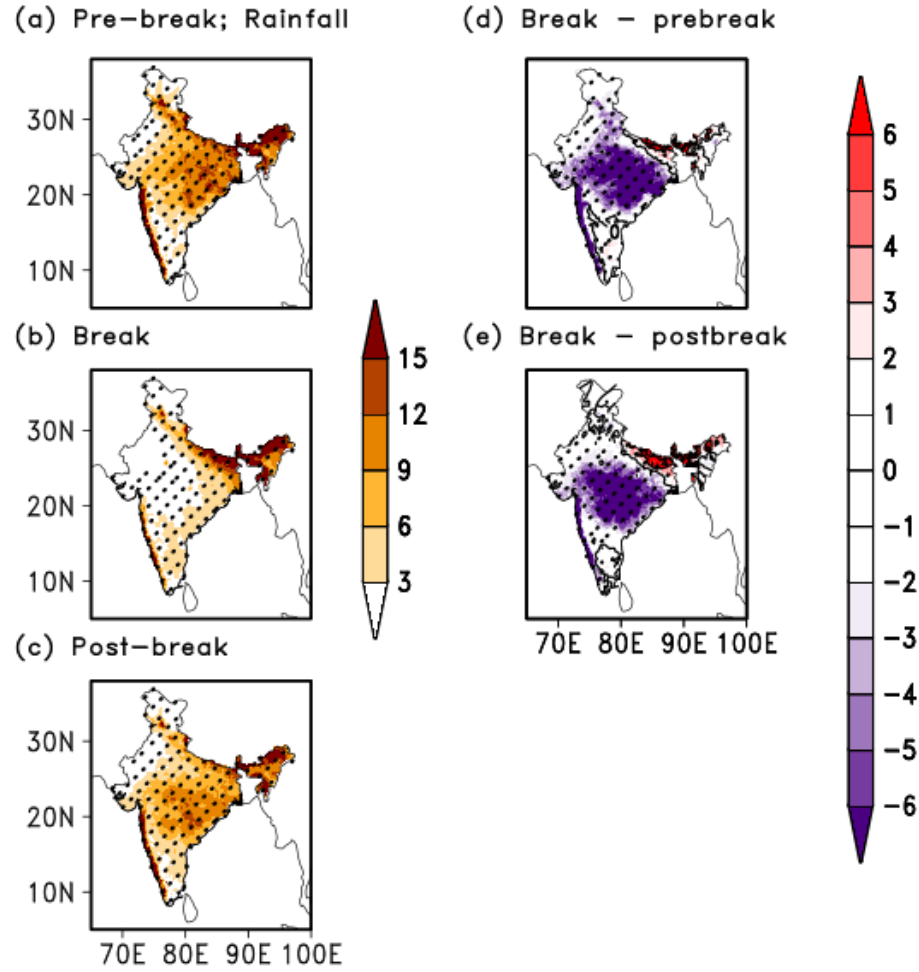
1557 decrease from south to north. Thus, a strong raising motion and high rainfall occurs. Also, as we

1558 suggested, in “all” cases of break monsoon a Blocking high forms. By considering anomalies of

1559 all the cases, we see in composite Fig. 3.2, a relative “upper level low”. This low, together with

1560 the implied convergence at 200 hPa and a high with divergence at the surface as indicated by the

1561 SLP (figures 3.2c & 3.2d) provokes the sinking ahead of the trough, as inferred above.



1562

1563 **Figure 3.8:** Composite Rainfall (mm) over Indian region during the Indian summer monsoon for
 1564 the 1979-2007 period, **(a)** prebreak periods **(b)** during-break periods, and **(c)** after-break periods,
 1565 **(d)** difference between the composite break and composite prebreak periods, and **(e)** difference
 1566 between the composite break and composite after-break periods. The shaded region in the panels
 1567 (a) (b) & (c) vary between 3 mm/day to 15 mm/day, with an interval of 3 mm/day. The shaded
 1568 region in panels (d) & (e) vary between 2 to 6 mm/day & -2 to -6 mm/day, with an interval of 1
 1569 mm/day. The hatched regions indicate composite rainfall are significant at 95% confidence from
 1570 a two-tailed Student's t-test.

3.3 Genesis and Maintenance of the Blocking High Episode

From the context of Figures 3.1 3.2, 3.6 & 3.7, the occurrence of monsoon breaks boils down to the question of how the blocking forms, develops, and subsides with time. Some theoretical aspects on blocking high formation, maintenance, and dissipation were proposed in the previous studies. The pioneering theory on the formation of blocking high by Charney and DeVore (1979), for example, envisages interactions of free Rossby waves of different wavelengths with those of stationary waves. They suggested that the resonance of stationary free (Rossby, 1939) planetary waves interacting with stationary Rossby waves forced by topography and/or differential heating generates large-amplitude planetary waves which are essential in the formation of a block.

In continuance, Fig. 3.9a depicts the global average of zonal wind composite at 200 hPa (hereafter as U_{200}), for prebreak, during-break and after-break periods. We see a peak of ~ 20 m/s, at 40°N in all three composites. This indicates a strong and narrow jet peak in the northern hemisphere, which is prevalent around mid-latitudes during the northern summer. Such strong and narrow jets form stable, zonally oriented waveguides, as per Kornhuber et al., 2017. Kornhuber et al., (2017) show that the preferred phase position of the zonal wind peak at a single location is conducive for QRA of waves with wavenumbers 7 and 8 (Kornhuber et al., 2017). This suggests that we can expect an amplification of such waves due to QRA. Sometimes, we also find bimodal peaks in zonal winds in the Northern hemisphere, from June to August (also includes Indian summer monsoon). As an example, we present, in Fig. 3.9b, the zonal wind at 200 hPa as a function of latitude during various stages of a strong break event, i.e., prebreak (23-31 July 2000), during-break (01-09 Aug 2000), and after-break (10-17 Aug 2000) period. During the prebreak phase, we note a maximum U_{200} of 23 m/s at 45°N and a secondary peak of about 7 m/s at 75°N . During the break phase, the peak U_{200} with a magnitude of 18 m/s is located at 50°N , and with a secondary peak at 80°N of

1594 value 11 m/s. These profiles are very similar to those given by Mann et al., (2018) (their Fig. 1b).
1595 The secondary peak at 80° N is highly favourable for the resonant amplification of the waves with
1596 wavenumbers 5-8 range. The profile during the break with a double peak structure of the zonal
1597 wind is also an efficient waveguide as noted by Manola et al., (2013). If the waveguide is
1598 circumpolar, then the wave energy is efficiently trapped and waves constructively interfere with
1599 the forced waves leading to resonance.

1600 To verify the possible application of Charney and DeVore (1979) theory, we show in Fig. 3.10, a
1601 composite of the amplitude spectra of the first ten harmonics of 200 hPa meridional wind during
1602 the prebreak conditions, at 45° N. We note that wave number 7 is prominent for the individual
1603 case, with an amplitude of ~ 14 m/s (Fig. 3.10). Indeed, the corresponding composite for the
1604 prebreak shows that the wavenumbers 5, 6 & 7 are dominant, with an amplitude of ~ 7 m/s (Fig.
1605 3.10). The increase in the amplitude of the wavenumber 7 as compared to its amplitude during
1606 prebreak is twice its prebreak value. Kornhuber et al (2017) suggest that most QRA events are
1607 found for wave number 7, which also is associated with the longest duration. This shows the
1608 important role of QRA for the formation of the Blocking high over India during the 41 break
1609 monsoon cases in agreement with several authors (Manola et al 2013; Kornhuber et al., 2017
1610 Mann et al., 2018).

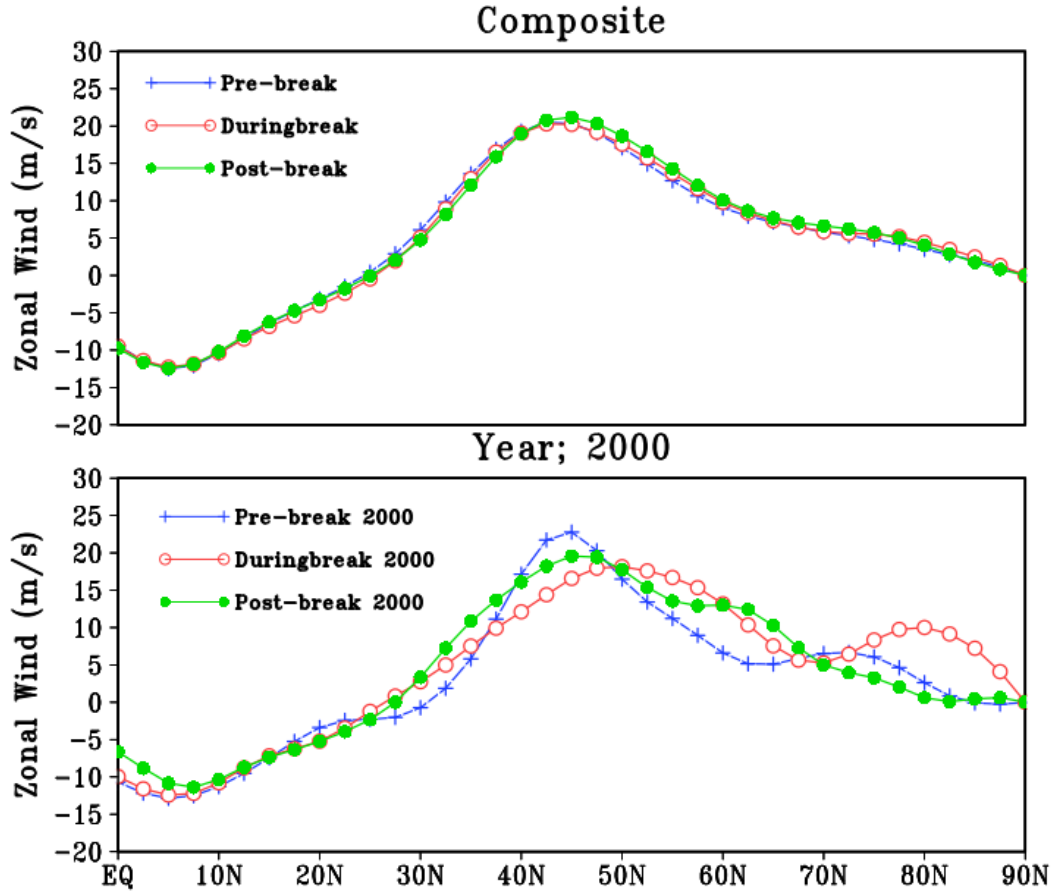


Fig. 3.9: Globally zonal averaged U wind (m/s); **(a) (upper panel):** for the composite prebreak, composite during-break & composite after-break periods of ISM (Table 1), cases available during the 1979-2007 period; **(b) (lower panel):** Similarly, for the case study, prebreak 23-31 July 2000, during-break: 01-09 Aug 2000 and after-break 10-17 Aug 2000 of ISM.

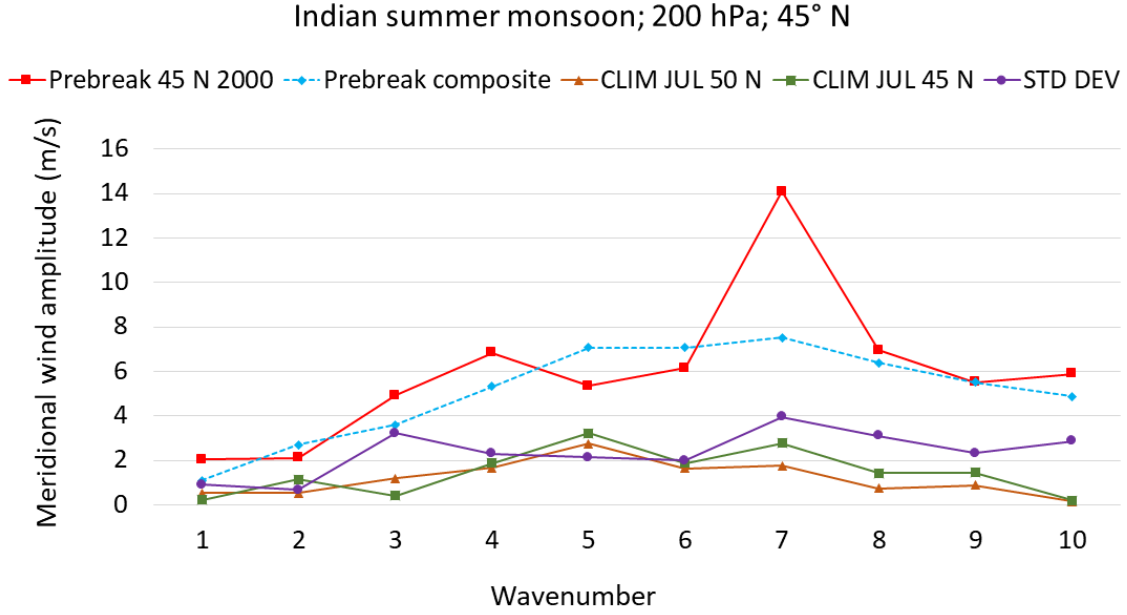


Fig. 3.10: Wavenumber spectrum using meridional wind (m/s) at 200 hPa level for; the case study prebreak: 23-31 July 2000, composite prebreak periods during 1979-2007 of the Indian summer monsoon, climatological mean of month July during the 1979-2018 period at latitude 45° N & 50° N, and standard deviation for all the ten harmonics of July at 45° N for the period 1979-2018.

As mentioned earlier, Edmon et al., (1980) derived the EP flux (vector $F_{(\phi,P)}$, from equation 1) from the quasi-geostrophic equation for the large-scale studies (Andrews et al., 1983). The EP flux is a direct measure of the total eddy forcing of the zonal mean state by eddies. Edmon et al., (1980) suggest that the $F_{(\phi,P)}$ (see equation 1) is also a measure of net wave propagation in the vertical and the latitudinal direction. Edmon et al., (1980) also mentioned that for quasi-geostrophic planetary waves, the sense of arrow representing $F_{(\phi,P)}$ gives the sense of the group velocity. Also, the contours of $F_{(\phi,P)}$ represent the role of zonal force on the mean state by the total effect of the quasi-geostrophic eddies.

1631 Fig. 3.11a shows the EP flux for the composite of 41 break cases during the period 1979 - 2007.
 1632 In this figure, we see a convergence maximum at 150 hPa level near 30°N , and even at 200 hPa,
 1633 close to 30°N . This convergence of EP fluxes implies a deceleration of zonal wind, thus keeping
 1634 the blocking high in place during the period of the breaks. Also, this high over the main monsoon
 1635 region provokes sinking and scanty or no rainfall or a break monsoon. Subsequently, Fig. 3.11b
 1636 depicts EP fluxes for the JJAS climatology during the 1979 – 2019 period, which shows much
 1637 more weakened convergence at 150 hPa & 200 hPa. Similarly, in the single case study (Fig. 3.11c),
 1638 the EP cross-section shows a very strong convergence at 200 hPa around 30°N during the 01-09
 1639 August 2000 period, particularly when the blocking high is observed over the Indian region.

1640 Further, during the amplification of the Rossby waves of the sub-tropical westerly jet due to QRA,
 1641 we observe the development of two remarkable features during the break phase i.e., (1) strong
 1642 zonal wind shear in the lower latitudes and (2) a split in the jet (core) structure is seen in the higher
 1643 latitudes, these are the typical characteristics of the blocking high, as noted by Trenberth (1986)
 1644 for the Southern Hemisphere (SH) scenario. In this context, Fig. 3.12 shows the cross-section of
 1645 mean zonal wind anomalies averaged over 60°E to 100°E for all prebreak, during-break and after-
 1646 break periods. Here, this feature is noted for the first time during the southwest monsoon break
 1647 case and is statistically significant at 95% confidence using a two tailed student's t-test. The strong
 1648 zonal wind shear is known to cause barotropic instability generating a monsoon depression (Rao,
 1649 1971; Govardhan et al., 2017). In the case of a single case study, Fig. 3.13 shows the cross-section
 1650 of daily mean zonal wind averaged over 60°E to 100°E from 23July – 18 August 2000. Initially
 1651 on 23July 2000, a westerly jet shifts south to 50°N and an easterly jet to 5°N . At the beginning of
 1652 the break phase, over CMR on 31 July 2000, the easterly jet intensifies and shifts north to 20°N ,
 1653 and the westerly jet shifts south to 35°N , generating a strong zonal wind shear in the upper

1654 troposphere in lower latitudes. Then, further to the north of 60° N, the increasing speed of the sub-
1655 tropical westerly jet is associated with a split jet structure on 06 August 2000.

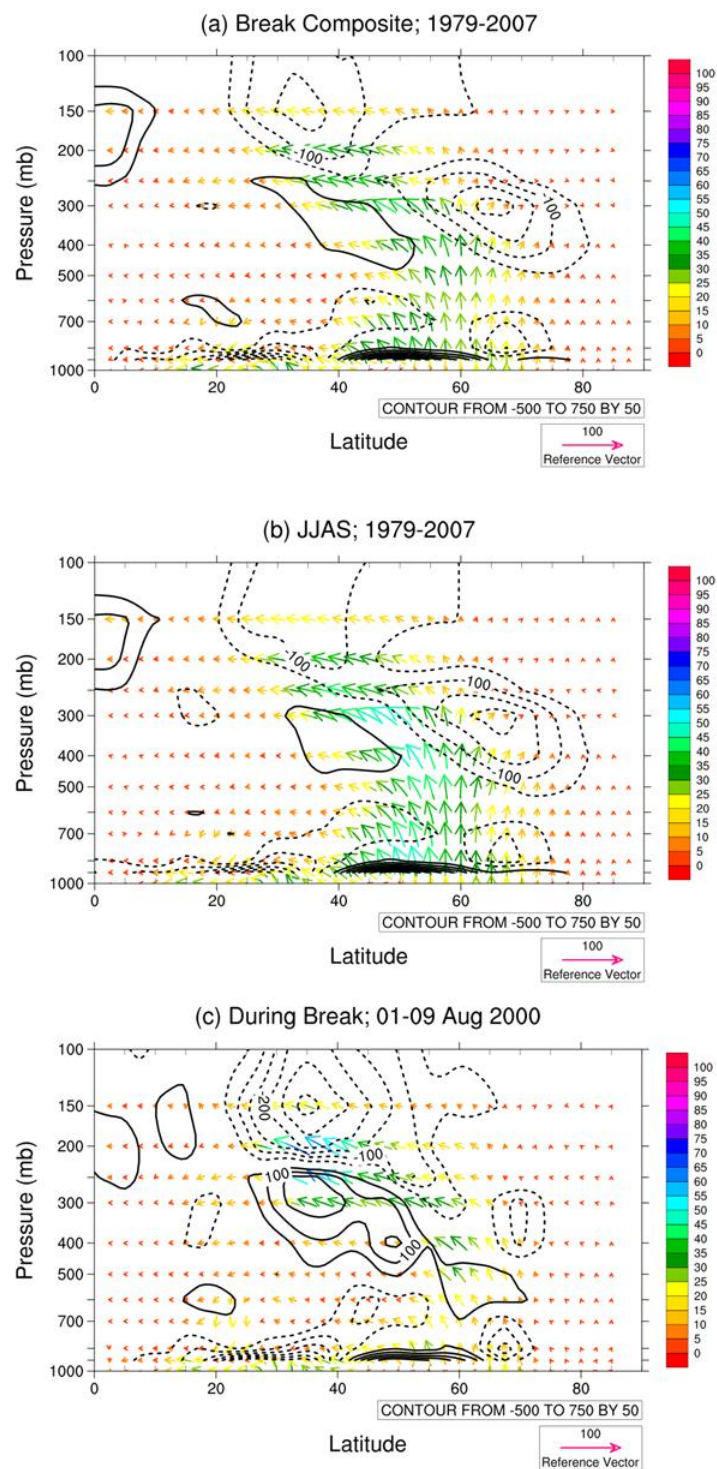


Fig. 3.11 (a): Eliassen-Palm fluxes cross-section of composite during-break periods of Indian summer monsoon; **(b):** Climatological Eliassen-Palm fluxes for the summer season, JJAS during the 1979 – 2007 period. **(c):** Eliassen-Palm fluxes for the case study during the break period 01-09 August 2000 of ISM. In all the figures, the contour interval is 50 m^3 . The dashed lines represent convergence and continuous lines represent divergence.

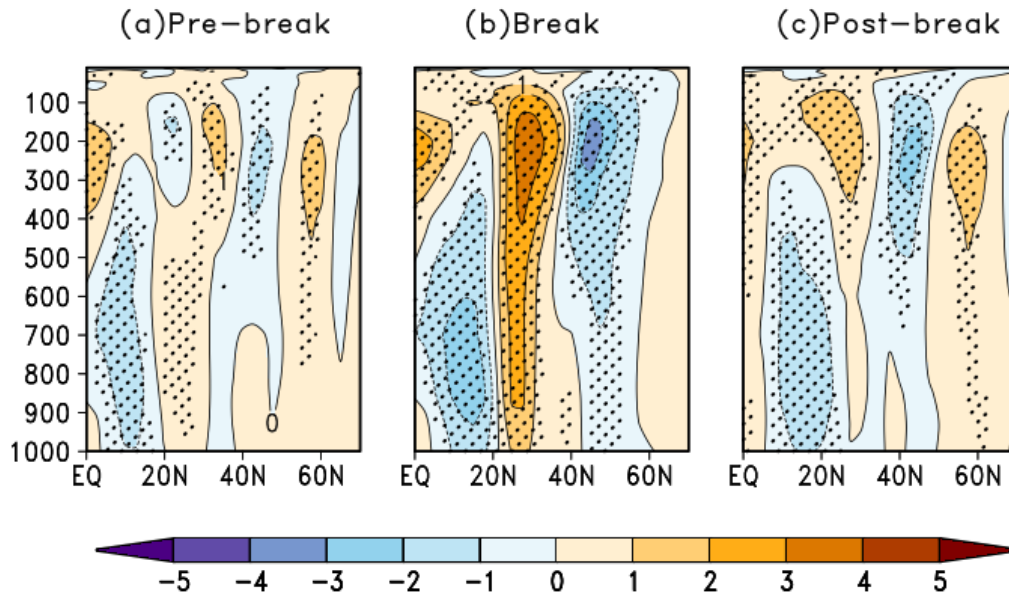
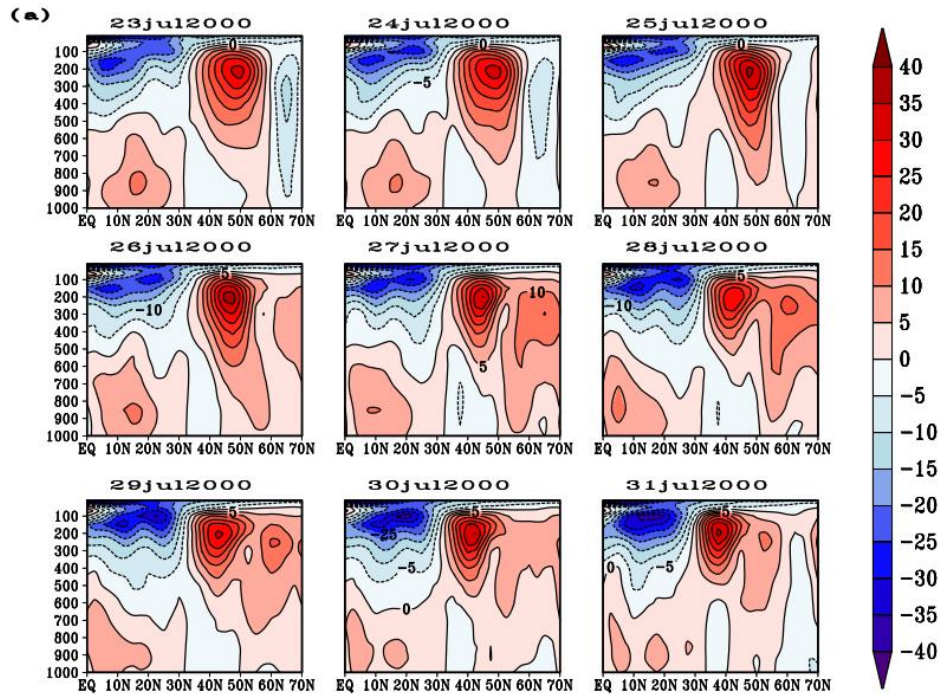
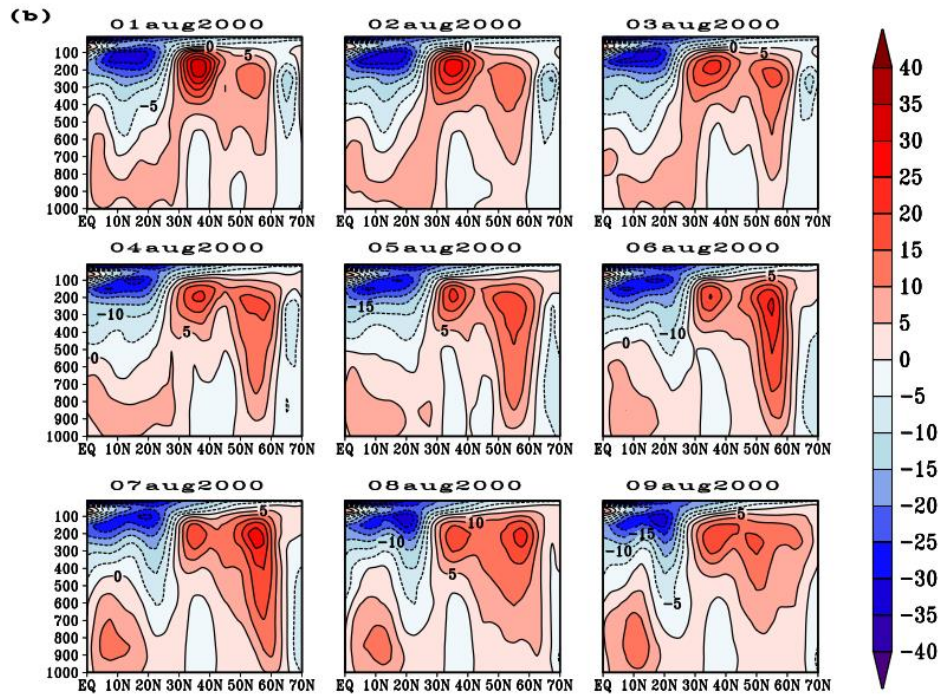


Figure 3.12: Vertical-latitude structure of zonal anomalies of daily U wind (m/s) which are longitudinally averaged between 65°E to 100°E , for **(a)** ccomposite prebreak, **(b)** composite during-break and **(c)** composite after-break spells of Indian summer monsoon. The hatched region is significant at 95% confidence using a two-tailed Student's t test. In all the figures, contouring vary between -5 to 5 (m/s) with an interval of 1 m/s



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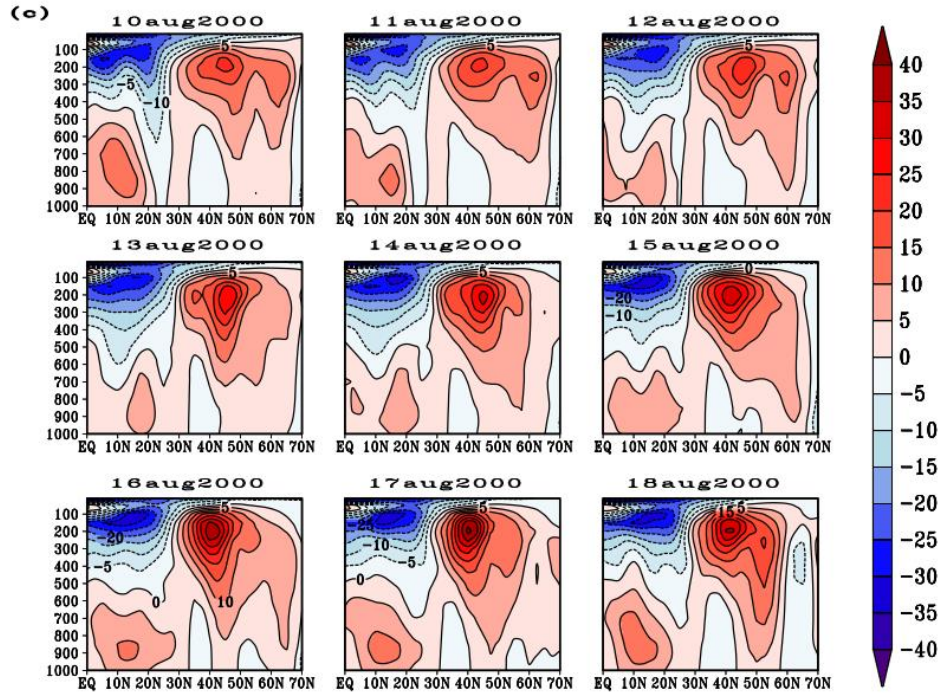


Figure 3.13: Structure of daily zonal wind anomalies which are zonally averaged between 60°E to 100°E, before a break **(a)** prebreak (23-31 July 2000); **(b)** during-break (01-09 August 2000) and **(c)** after-break (10-18 August 2000).

Fig. 3.14a shows the latitudinal variation of composite zonal wind, eddy momentum transport, and condition for the barotropic instability (henceforth as CBI) variation for prebreak, during-break and after-break periods. Similarly, the case study (Fig. 3.14b) illustrates the break period 01-09 August 2000 at 200 hPa. A subtropical westerly jet at 35° N and a high latitude jet at 55° N are noted. The eddy momentum transport increases from the equator to around 42° N, decreases to a minimum at 60° N, and then increases again. The necessary condition for barotropic instability (Kuo, 1949; 1951), $\beta - \frac{d^2u}{dy^2} = 0$ occurs at 20° N, 40° N, 60° N and near 80° N. A split jet can be seen at 45° N with mean zonal wind less than 10 m/s.

1683 A similar study by Trenberth (1986) (Fig. 2 in Trenberth's paper) shows the existence of CBI,
1684 indicating that eddies may grow locally at the expense of the kinetic energy of the zonal flow, on
1685 the poleward flanks of the two jets. A minor difference was observed between our results and
1686 Trenberth's (1986) analyses, that the zonal wind values in our study for the Northern Hemisphere
1687 (NH) summer are much lower than in his work for the SH. In our case (Fig. 3.14), there is a
1688 divergence of eddy momentum $\overline{u'v'}$ between 30° N and 45° N and the zonal wind (U) is positive
1689 from 30° N and 45° N. Thus, from equation 2 of methodology, we can infer an increase of eddy
1690 kinetic energy or barotropic instability. Similarly, on the poleward side of the mid-latitude jet, there
1691 is a divergence of momentum from 65° N to near about 80° N and U is positive. Again, this indicates
1692 the occurrence of barotropic instability. This instability is similar to that noted by Rao (1971) and
1693 Govardhan et al., (2017).

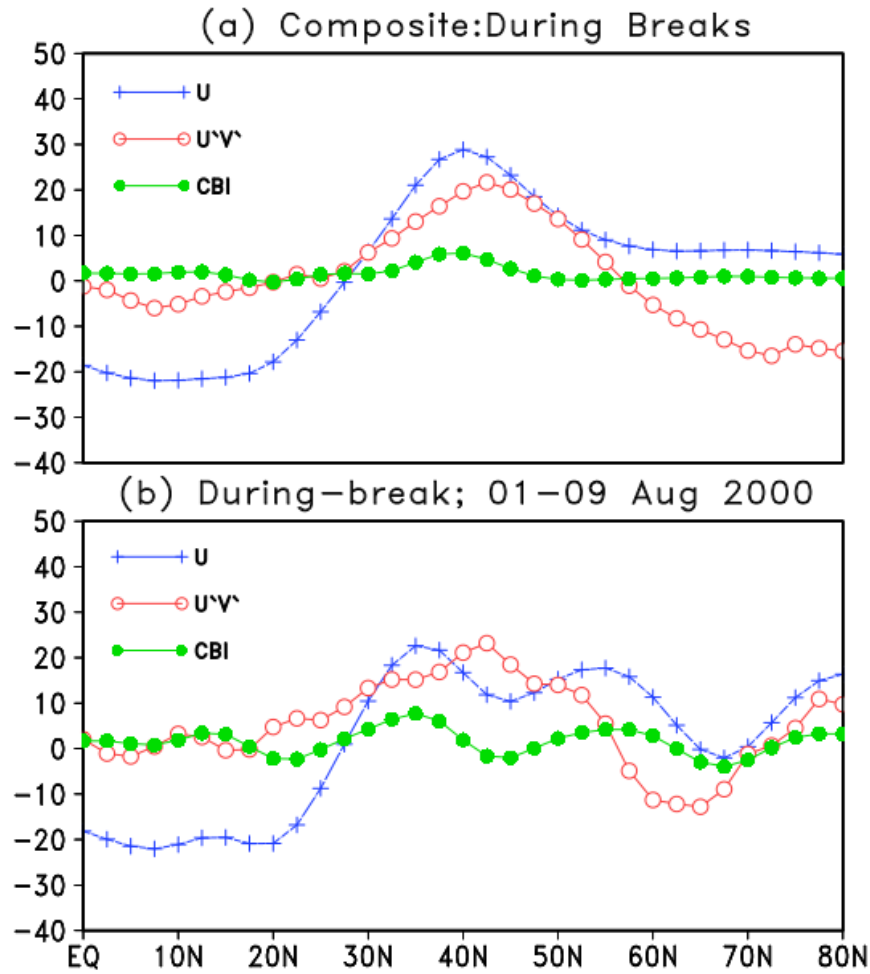


Fig. 3.14: (a) Latitudinal variation of zonal wind (m/s), eddy momentum (m^2/s^2) transport and condition for the barotropic instability (CBI), which are zonally averaged between $60^\circ E$ to $100^\circ E$; for composite during-break periods at 200 hPa during the 1979-2007 period of ISM. **(b)** Latitudinal variation of zonal wind (m/s), eddy momentum (m^2/s^2) transport and condition for the barotropic instability (CBI) for the case study during the break period 01-09 August 2000 at 200 hPa height of ISM.

3.4. Summary of the chapter

Using observed rainfall datasets and NCEP/NCAR II reanalysed datasets over the broad period 1979-2007, I study 41 known break cases in the Indian break summer monsoon. I find that the formation of a blocking high over sub-tropical west Asia a few days earlier is crucial for the manifestation of break monsoon conditions over the India region. Importantly, our study essentially shows that the formation of these blocking highs preceding the breaks is due to the quasi-resonant amplification of planetary waves.

For the occurrence of breaks, the aforementioned transient blocking high is weaker than the seasonal high pressure at upper levels over the Indian region. Because of the presence of this transient blocking, the high pressure weakens, causing a sinking motion relative to the prebreak or after-break period, resulting in a filled-up surface low and, eventually, a reduction in daily rainfall. The following chapter examines whether the sequence of events associated with Indian breaks, namely the formation of an upper level blocking high due to resonance of planetary Rossby waves at mid-latitudes, followed by the break and relative sinking over the core monsoon region, could also explain North and South American summer monsoon breaks.

Chapter 4

Breaks in North & South American summer monsoons

Outline of the chapter

In this chapter, I investigate the interactions between the upper level midlatitudes and tropical summer monsoon circulation in North and South American regions, which is similar to the break formation mechanism discussed in chapter 3.

In this context, I look into similarity between the Indian summer monsoon and the North and South American summer monsoon systems, specifically the mechanism for a break phase, with the goal of developing a unified theory.

4.1 Introduction

In continuation of chapter 4, this chapter is primarily concerned with identifying similar break dynamics between ISM and NAM, SAM systems. The mean evolution of the monsoon from Mexico to the USA, together known as NAM region, is characterized by the regular northward progression of heavy precipitation from southern Mexico by early June, which quickly spreads north into the southwest United States of America (USA) by early July (Higgins et al. 1999). Cavazos et al. (2002) made a systematic and important study about the NAM for the period 1980 - 93. They found “bursts” and “breaks” in a monsoon season. In a spectral analysis of daily summer rainfall, they also found a significant peak at 12-18 days period and also a secondary significant peak near 40 days period. A zonal wet mode (enhanced monsoon ridge) is the most typical mode that characterizes the mature phase of the monsoon in the southwest states of the USA. Wet (dry) summers in the southwest states of the USA are linked to strengthening and northward displacement (weakening and southward displacement) of the monsoon anticyclone (Carleton et al., 1990; Douglas et al., 1993; Higgins et al., 1998; Castro et al., 2001). But, Cavazos et al. (2002) found that eastward location of anticyclone is associated with wet conditions.

The onset of monsoon in Arizona is frequently abrupt, due to the sudden replacement of desert air mass with humid air mass (Moore et al., 1989). Very similar conditions happen over the South American Monsoon region and perhaps even over India, the monsoon onset is abrupt (Rao & Ergodan 1989). Higgins and Shi (2001) noted wet and dry periods of NAM, in western Mexico and South West USA and are linked to 30-60 day variations associated with MJO. This could be again seen similar to India's summer monsoon variations. In the case of the South American Monsoon System (SAM), during the austral summer season as well, there exist periods of enhanced (reduced) convection and higher (lower) rainfall in central and South Brazil (Bolivian

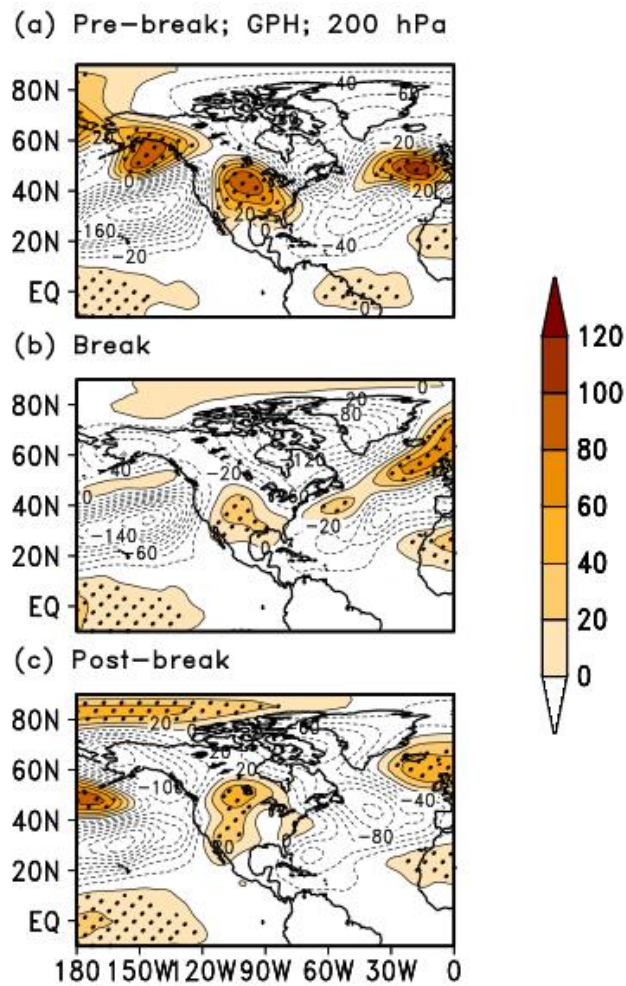
1763 Altiplano and northern South America). The direction of wind anomalies in the Rondonia state,
1764 Brazil, is used by Jones and Carvalho (2002) to classify a 3-4 day period known as the Westerly
1765 and Easterly low-level wind regimes. During the prevalence of westerly (easterly) anomalies,
1766 enhanced (reduced) convection and rainfall is observed over the Bolivian Altiplano and in the
1767 North of South America (their figure No. 9) (Jones and Carvalho, 2002), which also shows a clear
1768 blocking high, similar to the Indian summer monsoon case.

1769 Similar to the mechanism of break phase of ISM, i.e., the potential relevance of mid-latitude
1770 blocking high for the break monsoon, it is important to ascertain the robustness of this hypothesis
1771 and its implications for changing of meridional zonal wind stress based on examination of the
1772 sufficient number of break monsoon cases. Furthermore, it will also be pertinent to examine
1773 whether these mechanisms could also apply to the break monsoon condition in the North and South
1774 Americas in addition to the Indian Monsoon region. These are the primary goals of the present
1775 investigation. In this context, this chapter is structured as follows. The following subchapter 4.2 &
1776 4.3 discusses the results of the conditions that lead to break phase and conditions for the genesis
1777 of blocking high using reanalysis data in the North and South American monsoonal regions. The
1778 key findings are summarized in section 4.4.

1779 **4.2 Breaks in North American Monsoon**

1780 Vera et al., (2006) carried out a brief discussion of intra-seasonal variability for the SAM
1781 and NAM systems. Vera et al., (2006) noted, a high located east of the United States of America
1782 during the dry (or break) period; in contrast, during the wet period, the high is located to the west
1783 (their figures 10a & 10b). The eastern high during the break period is reminiscent of the blocking
1784 high over Indian and South American regions.

1785 Fig. 4.1 shows the composite daily zonal anomalies of Z_{200} during prebreak, break and after-break
 1786 periods. In all the panels of Fig. 4.1, we note a high at 40° N and 90° W over the United States of
 1787 America (USA). During the break period, the corresponding composite Z_{200} zonal anomalies (Fig.
 1788 4.1a) over the North American region shows a clear blocking high, which is very much similar to
 1789 the dry period high (Fig. 4.1) of Vera et al., (2006). Also, one can note a high pressure at 40° N in
 1790 southern parts of the United States of America (USA). In our case, Fig. 4.2 illustrates the dry/break
 1791 rainfall period for 16-24 July 2000 over the NAM region. Notably, we see the deficit of rainfall
 1792 over the Mexican country and southern states of the USA.



1793

Figure 4.1: Composite daily zonal anomalies of Geopotential height (m) over the monsoon season of 1979-2007 in the North American region during (a) prebreak periods (b) break periods, and (c) after-break periods. The hatching in the Figures 10 a-c indicates the regions where the composite Geopotential height is significantly different from zero at 95% confidence level. Statistical significance has been obtained using a two-tailed one sample Student's t test. In all the figures, positive values are shaded, and negative values are shown in contours. Note that significance test has not been applied to negative values. Contours vary between -120 m to 0 m, with an interval of 20 m. Same interval in shading, as evidenced by the greyscale shown, is used for positive values.

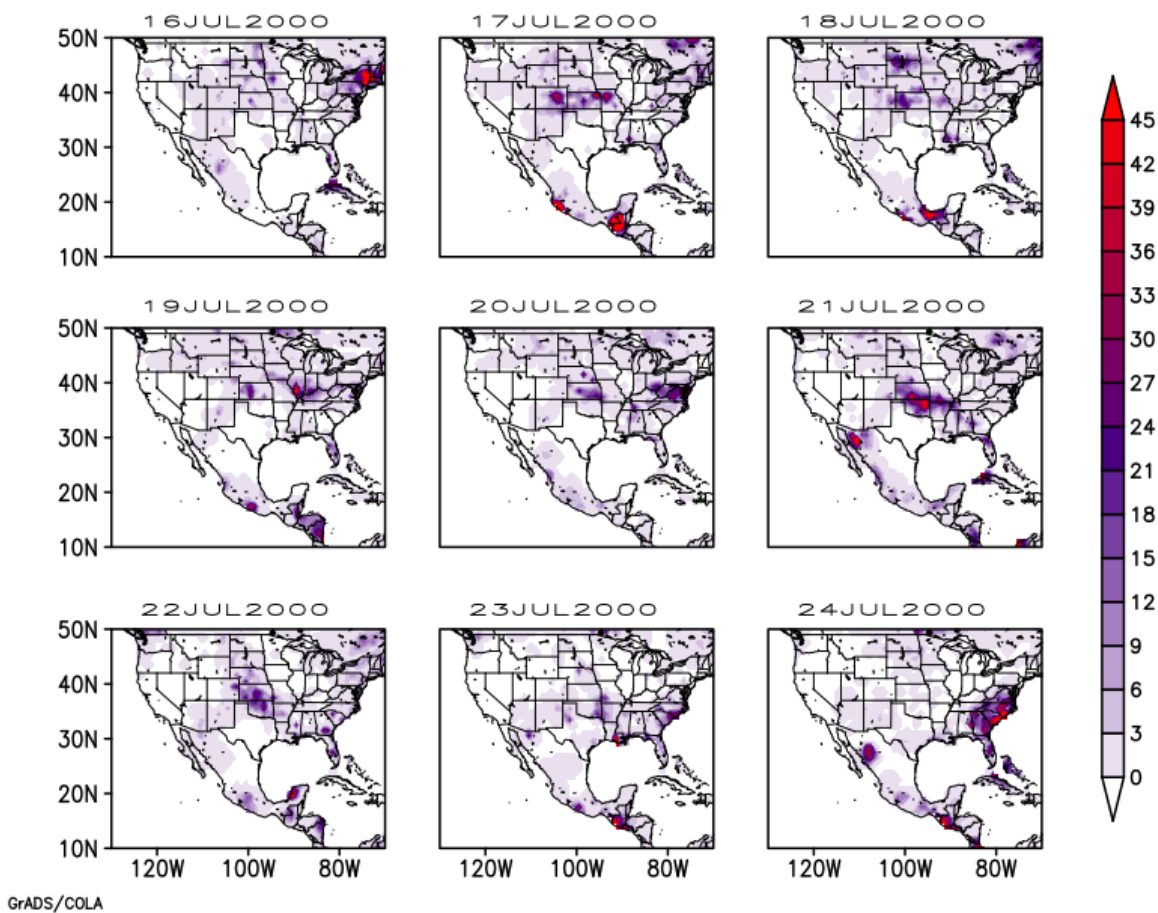


Figure 4.2: Daily rainfall during dry/break period, 16-24 July 2000 over the North American region.

1805

1806 Figures 4.3a & 4.3c show the differences in the composite Z_{200} between break& prebreak periods
1807 and that between SLP zonal anomalies. Figures 4.3b & 4.3d show corresponding differences
1808 between breaks & after-break periods. Figures 4.3a & 4.3c depict a clear low pressure region at
1809 200 hPa over the Southwest region of the USA during the breaks as compared to prebreaks. The
1810 SLP zonal anomalies show a corresponding relatively high pressure region to the eastern United
1811 States and a very weak low to the west. This situation, just as in the case of the Indian monsoon
1812 breaks, suggests a convergence at a high level and a divergence at a low level indicating sinking
1813 motion in accordance with the break conditions. In the after-break situation, figures 4.3b & 4.3d
1814 are somewhat similar situation prevails, but to the western side of USA. The similarity to the Indian
1815 summer monsoon in terms of the extent of significant signal isn't as obvious statistically because
1816 only a few cases were only available for this NAM composite analysis. Fig. 4.4 shows the ω (Pa/s)
1817 during the prebreak, break, and after-break periods, as well as the difference between the break
1818 and prebreak periods and the difference between the break and after-break periods for NAM
1819 system. A similar raising motion can be seen in the composite prebreak, break, and after-break
1820 periods in figures 4.4a, 4.4b, & 4.4c, but the magnitudes differ as shown in figures 4.4d & 4.4e. In
1821 Fig. 4.4d, we see a strong sinking motion (positive ω values) at 10° N and in Fig. 4.4e we note a
1822 very weak raising motion compared to the figures 4.4a, 4.4b & 4.4c. The changes in the ω in figures
1823 4.4a, 4.4b & 4.4c are significant at 95% confidence level, from a two tailed students t-test. Even
1824 though the composite ω in figures 4.4d & 4.4e is not as significant as expected, this weak raising
1825 motion can be considered as a good observation because we already note sinking motion in Fig.
1826 4.4d, which could be elucidated as a significant result leading to the NAM system's break phase,
1827 although only three break cases were found.

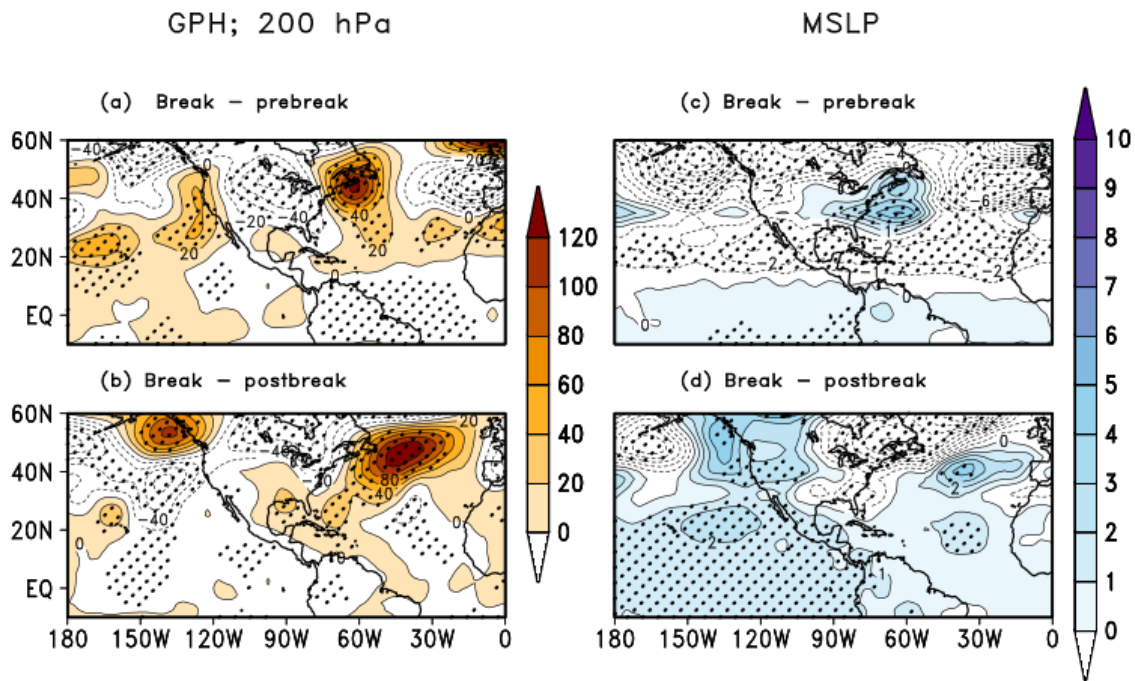
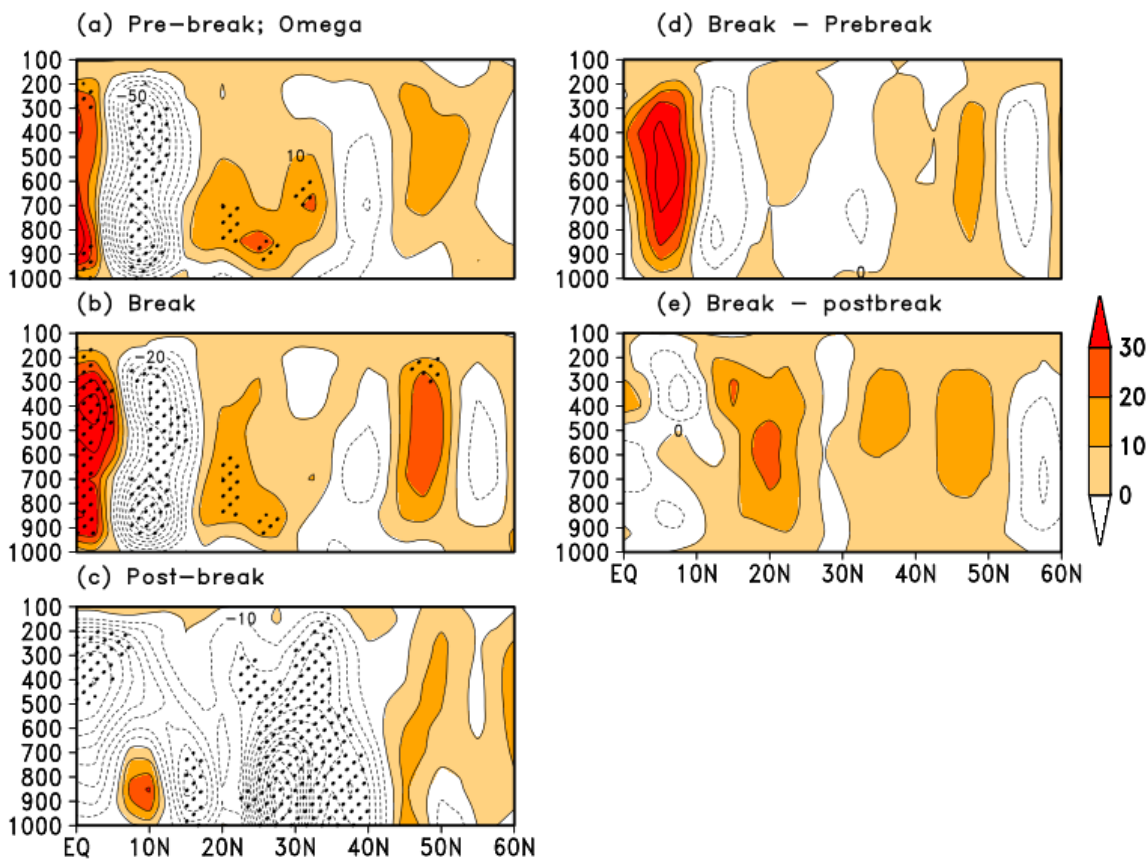


Figure 4.3: Difference between the composite daily zonal anomalies of Geopotential height (m) during the 1979-2007 period over the North American region between (a) break and prebreak periods (b) break periods and after-break periods. Difference in composites of zonal anomalies of daily mean sea level pressure (hPa) over the North American region during the 1979-2007 period between (c) during-break and prebreak periods, and (d) during-break and after-break periods. The hatched regions indicate locations of significant differences, at 95% confidence level, in daily zonal anomalies of Geopotential height in panels (a) & (b), and that of daily mean sea level pressure in panels (c) & (d). Statistical significance has been obtained using a two-tailed two sample Student's t-test with unequal variances. In all the figures, positive values are shaded, and negative values are shown in contours. The contours in panels (a) & (b) vary between -120 m to 0 m, with an interval of 20 m and those in panels (c) & (d) vary between -10 hPa to 0 hPa, with an interval of 1 hPa. Same intervals as used for contours are used for the shadings, and indicated in the grayscale shown.

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Figure 4.4: Composite vertical wind zonal anomalies ($\times 10^{-3}$) (ω) (Pa/s) over North American region and longitudinally averaged along $60^{\circ}\text{W} - 125^{\circ}\text{W}$ during the summer monsoon for the (a) prebreak periods (b) during-break periods, and (c) after-break periods, (d) difference between the composite break and composite prebreak periods, and (e) difference between the composite break and composite after-break periods. The contours in panels (a) (b) (c) (d) & (e) vary between -30 Pa/s to 0 Pa/s, with an interval of 10 Pa/s. The hatched regions indicate composite omega zonal anomalies are significant at 95% confidence from a two-tailed Student's t test.

1852

To explore the possible occurrence of resonance as suggested by Charney and Devore (1979) for the initiation of the blocking high in the North American monsoon breaks, we constructed the first ten simple harmonics for the meridional wind over the region (Fig. 4.5) for the prebreak at 45° N composite, July climatology at 45° N and 50° N, and an individual case for 1995 prebreak period. For the North American monsoon case, we note a dominant wave number 5 and 6 respectively for the 1995 case and the composite prebreak. Unlike, the wavenumber 7 for Indian summer monsoon case, we do not see a clear dominance of a single wave. This may probably be because of the small number of break cases we used for the NAM system. In any case, compared to the climatology and the standard deviation curves in Fig. 4.5, we do see that wave numbers 5, 6 and 7 are most dominant. This agrees with what is expected from the Charney and Devore perspective and supports the QRA theory for the break monsoon case of North America.

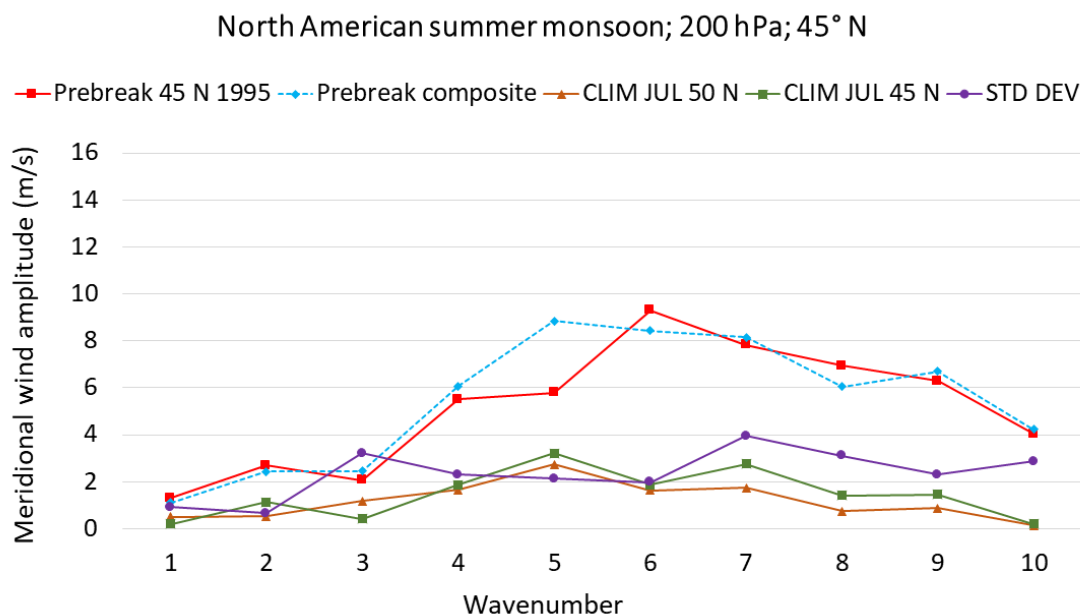


Figure 4.5: Wavenumber spectrum using meridional wind (m/s) at 200 hPa level; for composite prebreak periods of North American monsoon (Table 2.3); case study prebreak period, 10-15 Jul

1867 1995; the climatological mean of month July during 1979-2018 periods at 45° N and 50° N, and
1868 standard deviation for all the ten harmonics of July at 45° N for the period 1979-2007.

1869

1870 Fig. 4.6a shows the EP flux for the composite during-break periods of NAM (Table 2.3). In this
1871 figure, we see a convergence maximum at 150 hPa level near 30° N, and even at 200 hPa, close to
1872 30° N. This convergence of EP fluxes implies a deceleration of zonal wind, thus keeping the
1873 blocking high in place during the period of the breaks. Also, this high over the main monsoon
1874 region provokes sinking and scanty or no rainfall or a break monsoon. Subsequently, Fig. 4.6b
1875 depicts EP fluxes for a single case study, the EP cross-section shows a very strong convergence at
1876 200 hPa around 30° N during the 16-24 July 2000 period, particularly when the blocking high is
1877 observed over the North American region.

1878 Fig. 4.7a illustrates the composite of mean zonal wind and Fig. 4.7b shows the profiles of prebreak
1879 periods of each case. Fig. 4.8 shows the zonal wind profiles during the break phase for each case.
1880 Here, the composite zonal wind profiles (Fig. 4.7a) prebreak, during-break, and after-break periods
1881 show a minimal difference of about 23 m/s. In each case, during prebreak spells, we note zonal
1882 wind of 25 m/s and during the break phase (Fig. 4.8) we note 27 m/s. Manola et al., (2013) have
1883 shown that stronger and narrower jets tend to stronger zonal waves and smaller wavenumbers (4
1884 and 5). Also, we see a double-jet structure (Fig. 4.9) in the composite zonal wind anomalies and
1885 the steep sub-tropical jet is strongly linked to QRA (Kornhuber et al., 2017). The resonance
1886 amplification depends on the characteristics of mean zonal wind. In the case of SAM (ISM) the
1887 strength of the jet is higher and the stationary free waves are of higher wavelengths or lower
1888 wavenumbers (lower wavelengths and higher wavenumbers)

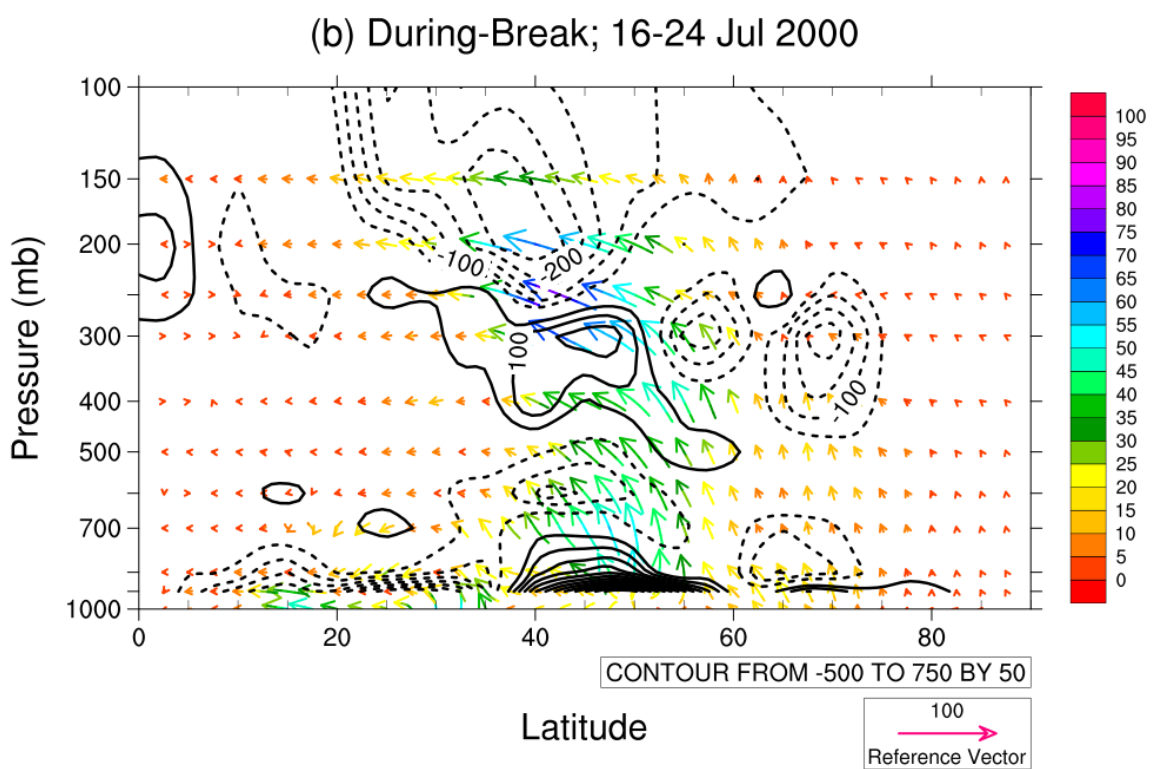
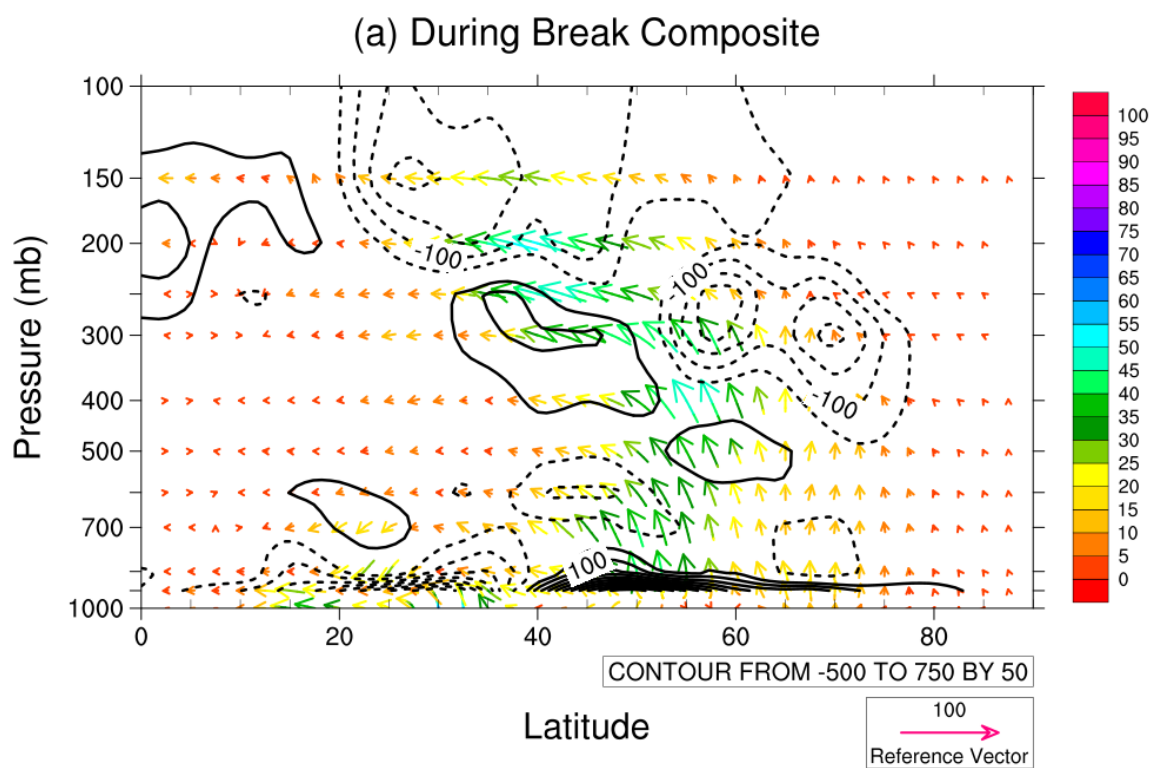


Figure 4.6: (a)Eliassen-Palm fluxes cross-section of composite during-break periods of NAM;
(b):Eliassen-Palm fluxes for the case study during the break period 16-24 July 2000 of NAM. In
all the figures, the contour interval is 50 m^3 . The dashed lines represent convergence and
continuous lines represent divergence.

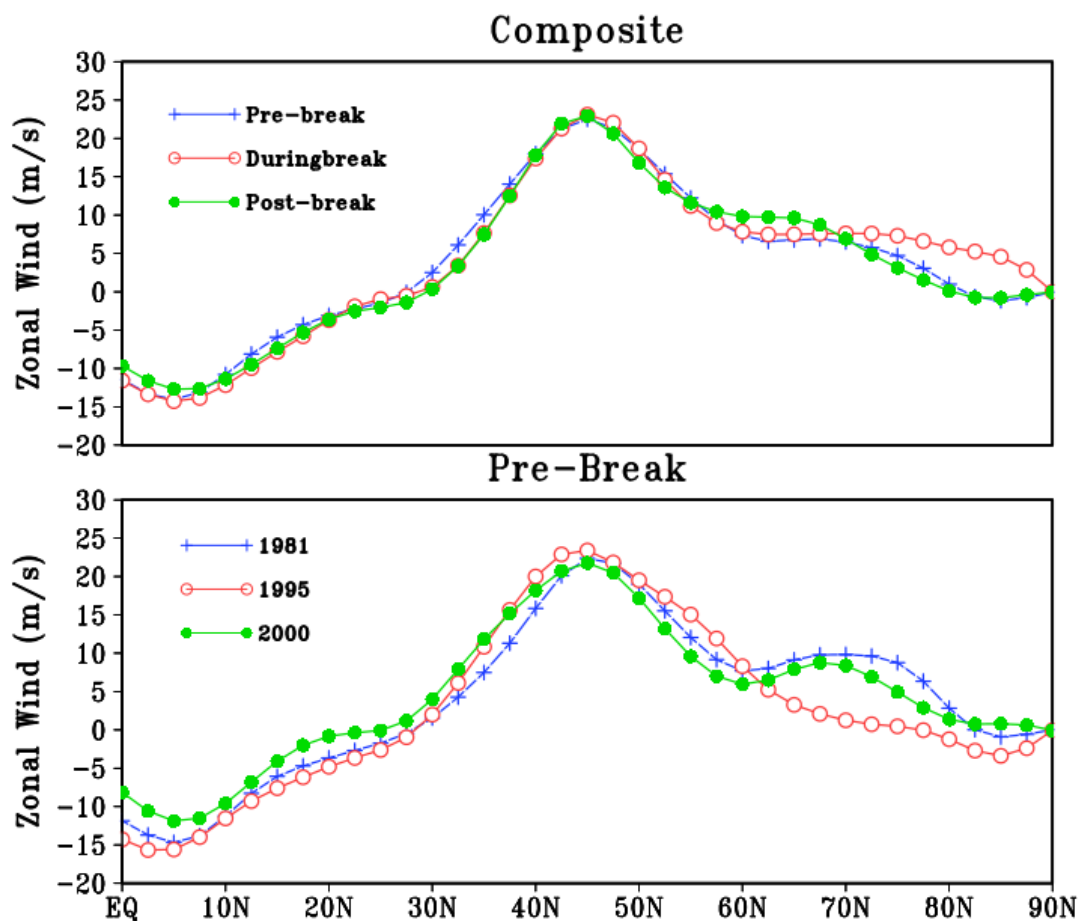


Figure 4.7(a):(Top) Composite globally zonal averaged U wind (m/s) of composite prebreak,
composite during-break and composite after-break periods of North American monsoon **(b):**
(Bottom) Globally zonalaveraged U wind of prebreak periods in three different cases: 10-16 July
1981, 20 - 24 July 1995 and 10-15 July 2000 of NAM.

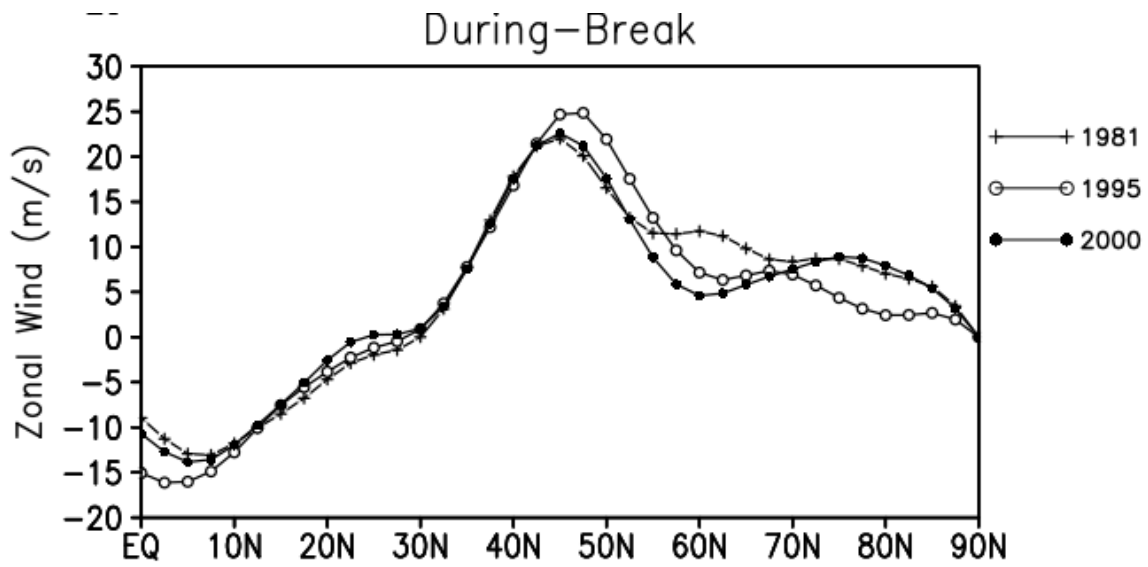


Figure 4.8: (c) Zonal wind profiles during the break phase for 1981, 1995 and 2000 events.

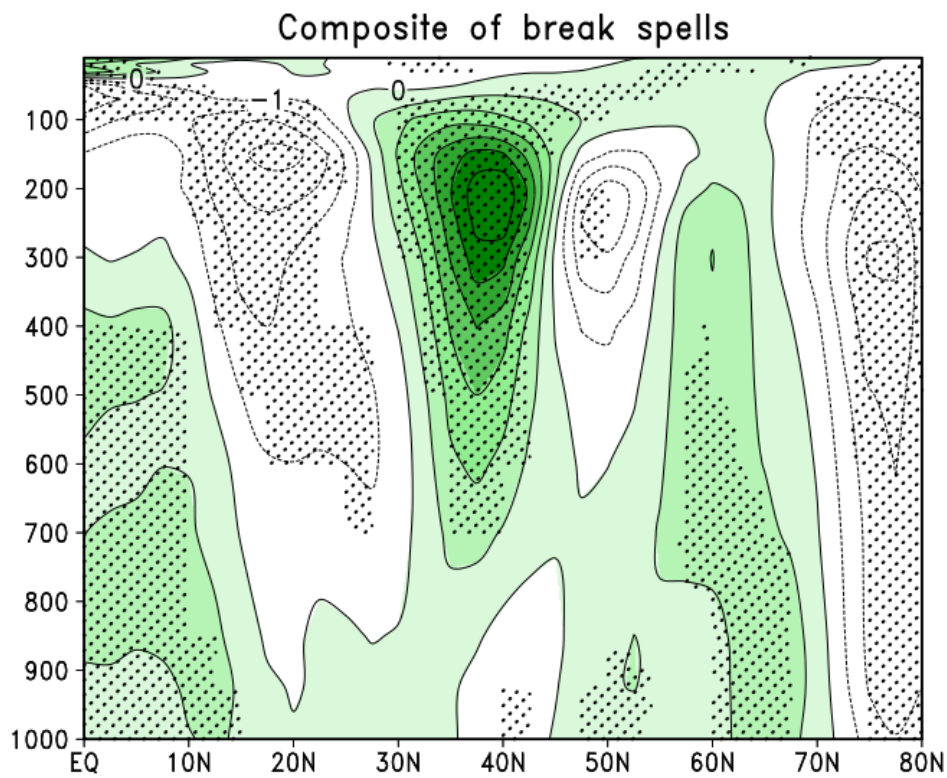


Figure 4.9: Composite structure of daily U wind zonal anomalies (m/s) for the break cases of NAM, 10-16 July 1981, 20 - 24 July 1995 and 10-15 July 2000; which are zonally averaged in between 60° W - 125° W. The hatched region indicates 95% confidence using a two-tailed Student's t test. In this figure positive values are shaded and negative values are shown in contours. Contouring between -5 to 0 (m/s) with an interval of 1 m/s

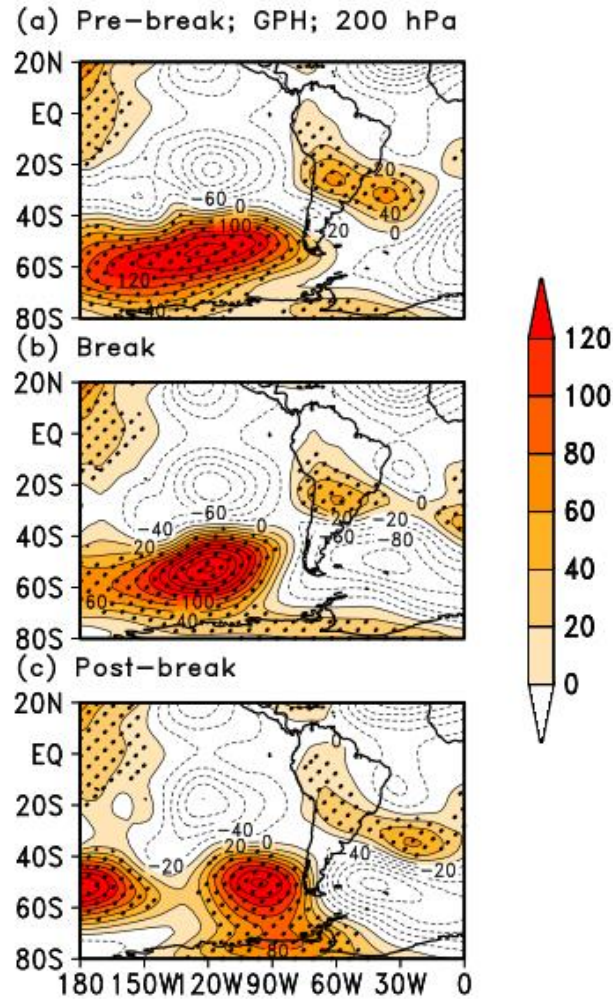
4.3 Breaks in South American Monsoon

The strong summer convective activity in South America is associated with large-scale atmospheric circulation. Intra-seasonal variability of SAM is related to convection on a wide range of spatial-temporal scales. Previously, the intra-seasonal variability of the 10-90 days scale is observed over eastern South America and particularly over northeast Brazil (Jones and Carvalho, 2002; Carvalho et al., 2002; Ciffelli et al., 2002; Grimm et al., 2005; Souza et al., 2005). Intra-seasonal variations of SAM may occur by the propagation of mid-latitude perturbations into the convection region, and as well as, with a subsequent increase in north-westerly cross-equatorial moisture transport over southern tropical South America (Carvalho et al., 2010) with the amplification of wave activity in the Northern hemisphere.

A study on SAM by Gan et al., (2004) found some similarities and differences with the Indian Monsoon using a 21-year (July 1979-June 2000) period of data. They noted a strong easterly jet maximum at 100 hPa during January, similar to the easterly jet over south India (Halley, 1686; Koteswaram, 1958). They found that only zonal wind anomalies show a perceptible reversal in direction, i.e., when the annual cycle is removed as noted by Zhou and Lau (1998). Further, they had observed wet and dry periods of SAM which are similar to the Asian Monsoon (Jones and

1926 Carvalho, 2002). Thus, the blocking high configuration in the upper troposphere, Fig 9 of Jones
1927 and Carvalho (2002) is similar as noted earlier for Indian monsoon breaks.

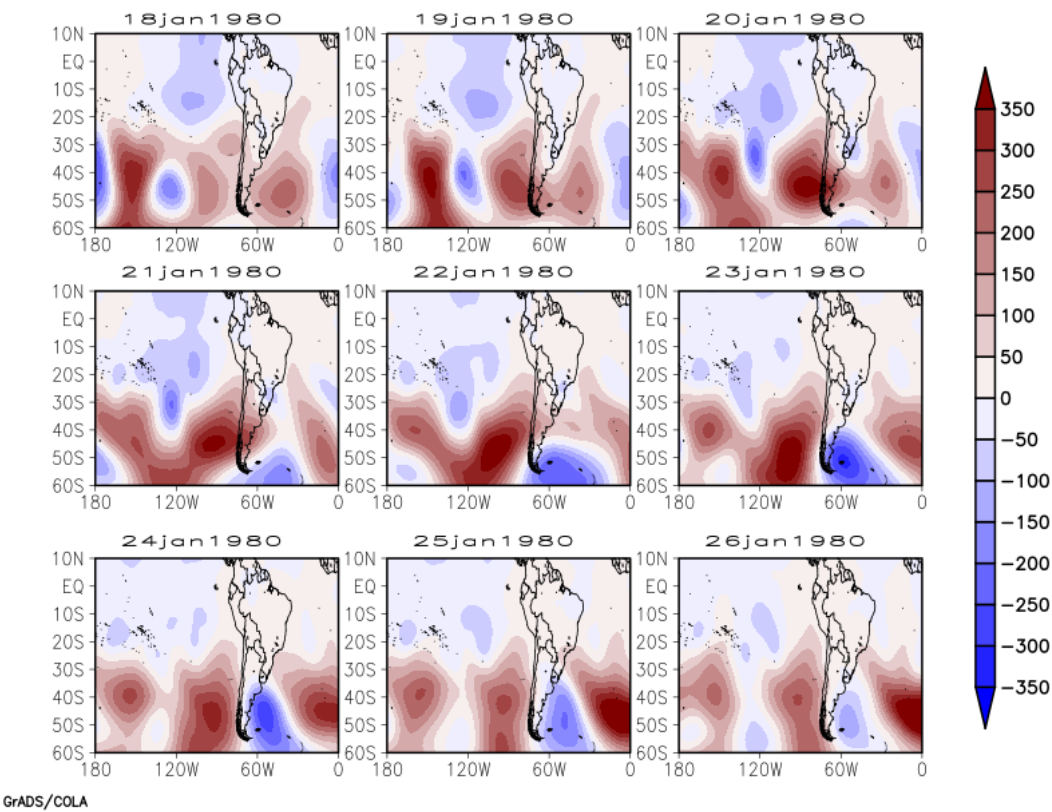
1928 Fig. 4.10 shows the composite zonal anomalies Z_{200} and all panels are statistically significant at
1929 95% confidence using a two tailed students t-test. We observe a high at 60° S & 120° W and another
1930 high extending between 40° S to 0° and at 60° E, i.e., central and north of South America. Figures
1931 4.11 a-b illustrates the Z_{200} anomaly pattern for the break period in January 1980 & 1981. A clear
1932 blocking high over central South America can be noted with two troughs on either side of the
1933 blocking high at 20° S. The formation of a blocking high preceding a break/drought situation in
1934 this case of SAM is analogous to that for the Indian summer monsoon situation (Fig. 4.11a). Thus,
1935 these features are a unifying characteristic for the break in the ISM (Fig. 3.1) and SAM. The
1936 troughs which are associated with blocking high are strongly oriented towards NW to SE tilt which
1937 implies southward (poleward) transport of eddy momentum. Fig. 4.11b depicts Geopotential
1938 Height for the break case during 06-11 January 1981. The characteristics in this figure are very
1939 similar to those in January 1980 case.



1940

1941 **Figure 4.10:** Composite daily zonal anomalies of Geopotential height (m) over the monsoon
 1942 season of 1979-2007 in the South American region during **(a)** prebreak periods **(b)** break periods,
 1943 and **(c)** after-break periods. The hatching in the Figures 17 a-c indicates the regions where the
 1944 composite Geopotential height is significantly different from zero at 95% confidence level.
 1945 Statistical significance has been obtained using a two-tailed one sample Student's t test. In all the
 1946 figures, positive values are shaded, and negative values are shown in contours. Note that
 1947 significance test has not been applied to negative values. Contours vary between -120 m to 0 m,

1948 with an interval of 20 m. Same interval in shading, as evidenced by the greyscale shown, is used
1949 for positive values.



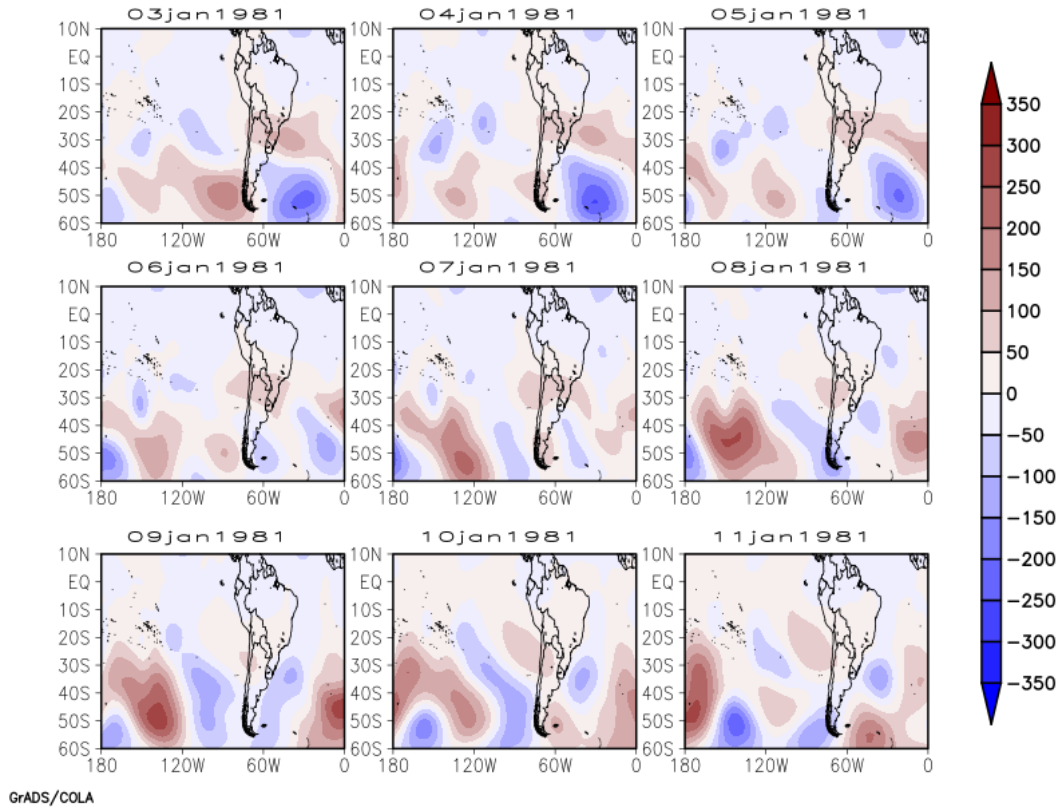
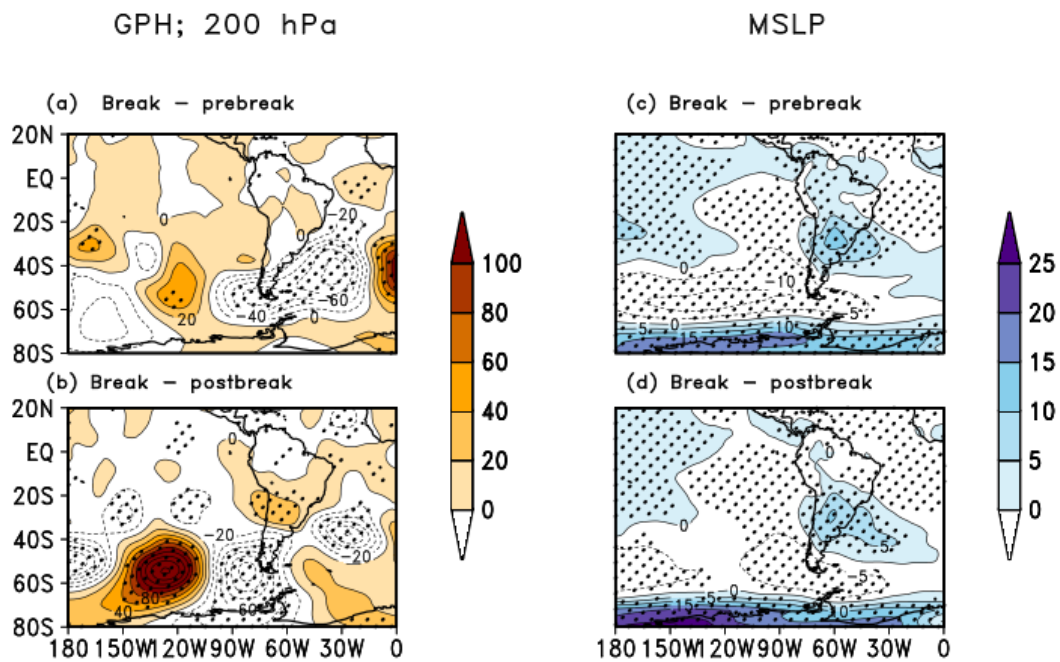


Figure 4.11: (a) Geopotential Height anomalies (m) at 200 hPa during the break spell, 21-26 January 1980 over the South American region. **(b)** Geopotential Height anomalies (m) at 200 hPa during the break spell, 06-11 January 1981 over the South American region.

Figures 4.12 a-d show the difference between the composites of break & prebreak periods, and breaks & after-break periods of Z_{200} and SLP zonal anomalies. In Fig. 4.12, we note a strong low pressure region at 200 hPa over southern South America and eastern South Atlantic. While in Fig. 4.12(c) we note a relatively high pressure region, this indicates again similar to Indian and North American break monsoons convergence at higher levels and divergence at lower levels (figures 4.12c & 4.12d; SLP figures) indicating a sinking motion. Thus, in all the break monsoon regions

1962 very similar conditions exist and the analysis using Z_{200} and SLP zonal anomalies, and strongly
 1963 supports sinking and reduction of rainfall causing break monsoons. All the changes in the panels
 1964 of the Fig. 4.12 are statistically significant at 95% confidence level using a two tailed students t-
 1965 test. Figures 4.13a, 4.13b & 4.13c shows the vertical velocity (ω) (Pa/s) of the South
 1966 American summer monsoon during the prebreak, break, and after-break periods, as well as the
 1967 figures 4.13d & 4.13e difference between the break and prebreak periods and the difference
 1968 between the break and after-break periods. A strong sinking motion is seen in the composite
 1969 prebreak& break between 10° S - 0° in figures 4.13a & 4.13b, and raising motion in after-break
 1970 periods between 10° S - 0° in fig. 19c, but the ω magnitudes differ as shown in Figures 4.13d
 1971 & 4.13e. In Figures 4.13d & 4.13e, we note a narrow region of raising motion at 10° S and strong
 1972 sinking between $0^\circ - 10^\circ$ N. The changes in the ω wind in figures 4.13 a-e are significant at
 1973 95% confidence level using two tailed students t-test.



1974

Figure 4.12: Difference between the composite daily zonal anomalies of Geopotential height (m) during the 1979-2007 period over the South American region between **(a)** break and prebreak periods **(b)** break periods and after-break periods. Difference in composites of zonal anomalies of daily mean sea level pressure (hPa) over the South American region during the 1979-2007 period between **(c)** during-break and prebreak periods, and **(d)** during-break and after-break periods. The hatched regions indicate locations of significant differences, at 95% confidence level, in daily zonal anomalies of Geopotential height in panels (a) & (b), and that of daily mean sea level pressure in panels (c) & (d). Statistical significance has been obtained using a two-tailed two sample Student's t-test with unequal variances. In all the figures, positive values are shaded, and negative values are shown in contours. The contours in panels (a) & (b) vary between -100 m to 0 m, with an interval of 20 m and those in panels (c) & (d) vary between -25 hPa to 0 hPa, with an interval of 5 hPa. Same intervals as used for contours are used for the shadings, and indicated in the grayscales shown.

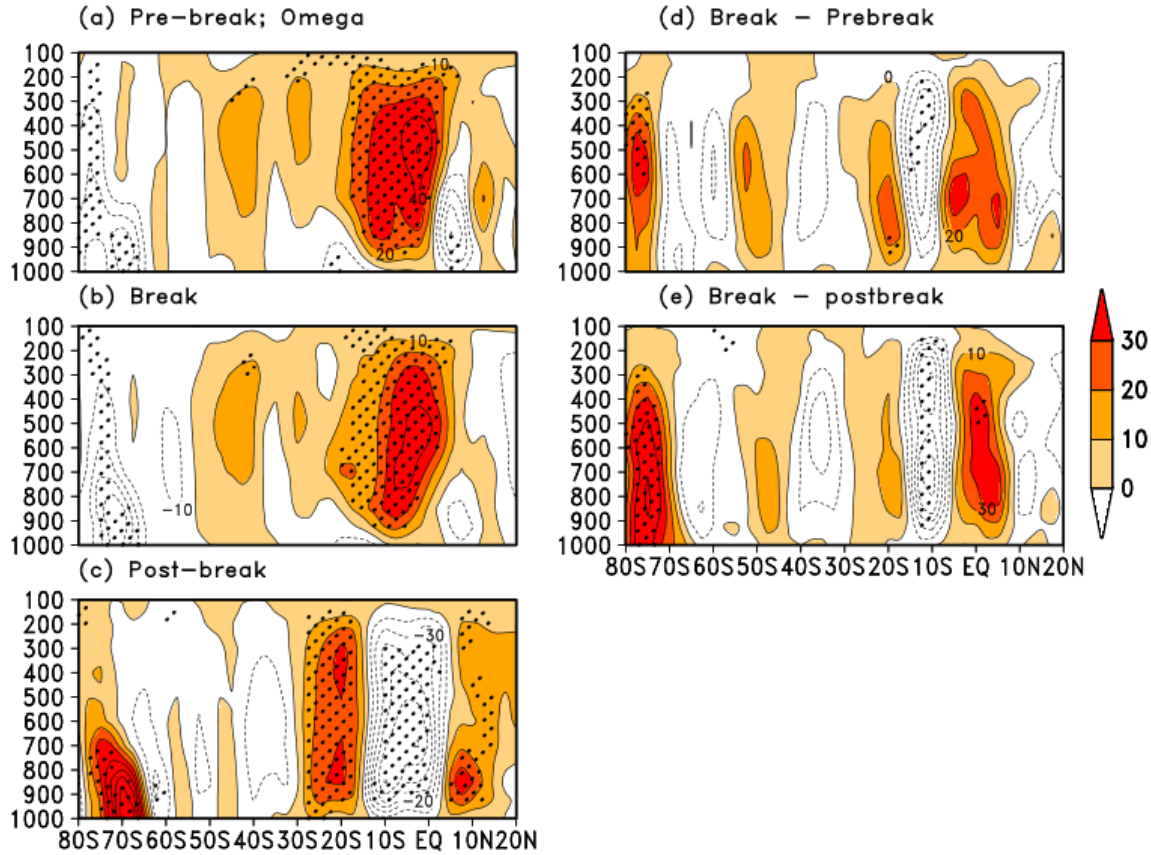


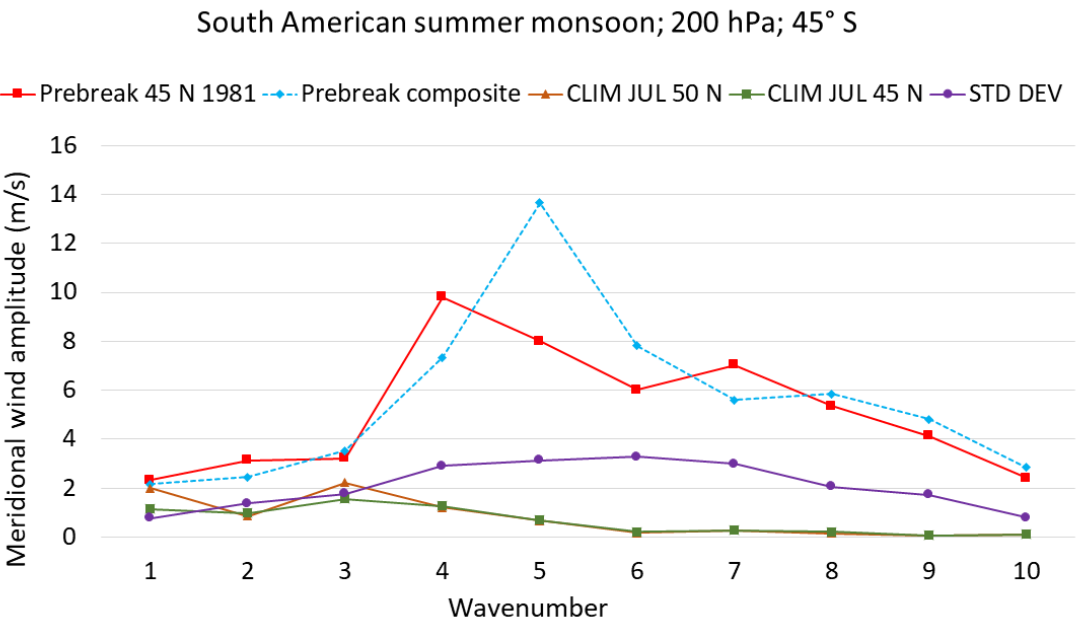
Figure 4.13: Composite vertical wind zonal anomalies ($\times 10^{-3}$) (ω) (Pa/s) longitudinally averaged along 60° W to 125° W over South American region during the summer monsoon for the 1979-2007 period, **(a)** prebreak periods **(b)** during-break periods, and **(c)** after-break periods, **(d)** difference between the composite break and composite prebreak periods, and **(e)** difference between the composite break and composite after-break periods. The contours in panels (d) & (e) vary between -30 Pa/s to 0 Pa/s, with an interval of 10 Pa/s. The hatched regions indicate composite ω zonal anomalies are significant at 95% confidence from a two-tailed Student's t-test.

Fig. 4.14 illustrates the wavenumber spectrum of meridional wind at 200 hPa height for composite prebreak periods, a case study during 1981 January at 45° S, and the climatology of January month for the 1979-2018 period at 45° S & 50° S, and standard deviation. For the South American monsoon case, we see the dominance of wave numbers 4 & 5 for the 1981 prebreak case and 4, 5 & 6 for the prebreak composite. The values of the amplitude for wave number 4 and 5 are much higher than for the climatology and also above standard deviation. In the case of austral summer Kornhuber et al., (2017) found waveguides approximately 70% for the analysed period for wavenumber 4 & 5. In our case, the dominance of wavenumber 4 & 5, thus is very similar to that noted by Kornhuber et al., (2017). This indicates a strong case for wave resonance as suggested by Charney and DeVore (1979) and quasi-resonant amplification for the SH case noted by Kornhuber et al., (2017).

The difference in the dominant wave numbers is due to the difference in the strength of the mean zonal wind. In the Indian break monsoon case, the maximum velocity of the zonal wind is around 22 m/s, as explained Kornhuber (2017, their Fig. 1) the QRA is expected for wave numbers ≥ 6 . While in the case of the SH the maximum velocity of composite zonal wind is around 28 m/s (Fig. 4.15a), as explained by Kornhuber et al., (2017) leads to the QRA mechanism for wave numbers ≥ 4 . Fig. 4.15b shows the zonal wind profiles for the prebreak period of each case, and Fig. 4.16 shows the during-break and after-break periods for all cases. Jones and Carvalho (2002) studied independent events of westerly and easterly wind samples, 113 and 104 respectively for SAM (their Fig.7). Westerly and easterly phases are associated with lower and higher rainfall over the SAM core monsoon region. Their fig. 9 shows a clear blocking high at 200 hPa level during the westerly regime and break monsoon case. Note that these characteristics are for a composite of 104 break events. Also, they mentioned in their figure, the resemblance to wave trains implying

2021 Rossby wave propagation akin to those envisaged by Charney and DeVore (1979), Kornhuber et
 2022 al., (2017), and others mentioned earlier. Thus, the break monsoon events in South America are
 2023 analogous to break events in Indian CMR.

2024



2025

2026 **Figure 4.14:** Wavenumber vs. Amplitude spectrum using meridional wind (m/s) at 200 hPa, for an
 2027 individual case 1981 at 45⁰ N, composite of prebreak periods of South American monsoon (Table
 2028 3), the climatological mean of month July during 1979-2018 periods at 45⁰ N and 50⁰ N, and
 2029 standard deviation for all the ten harmonics of July at 45⁰ N for the period 1979-2007.

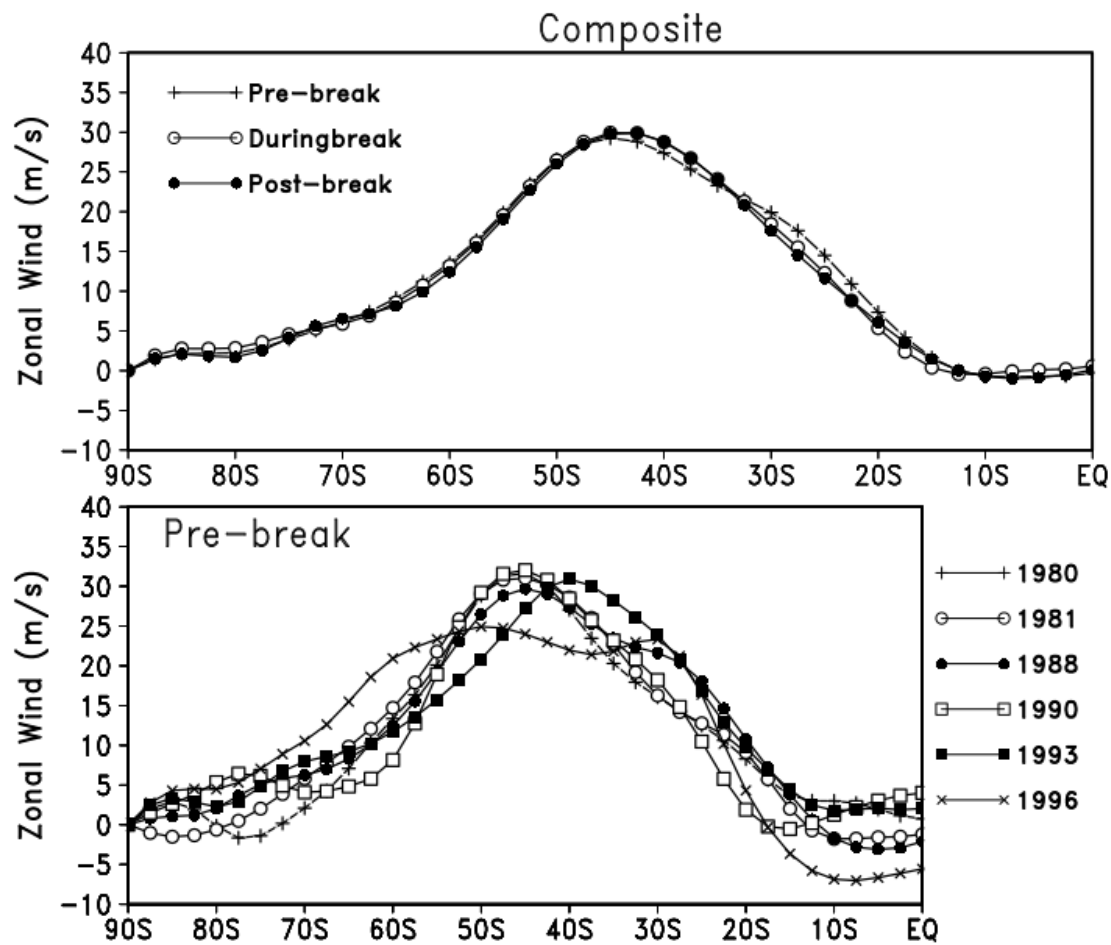


Figure 4.15(a):(Top) Globally zonal averaged U wind (m/s) of composite prebreak, composite during-break and composite after-break periods of SAM **(b): (Bottom)** Globally zonal averaged U wind (m/s) for prebreak periods of six different cases of SAM: 18-21 January 1980, 01-04 January 1981, 10-15 January 1988, 16-20 January 1990, 25 December-01 January 1993, 19-24 January 1996.

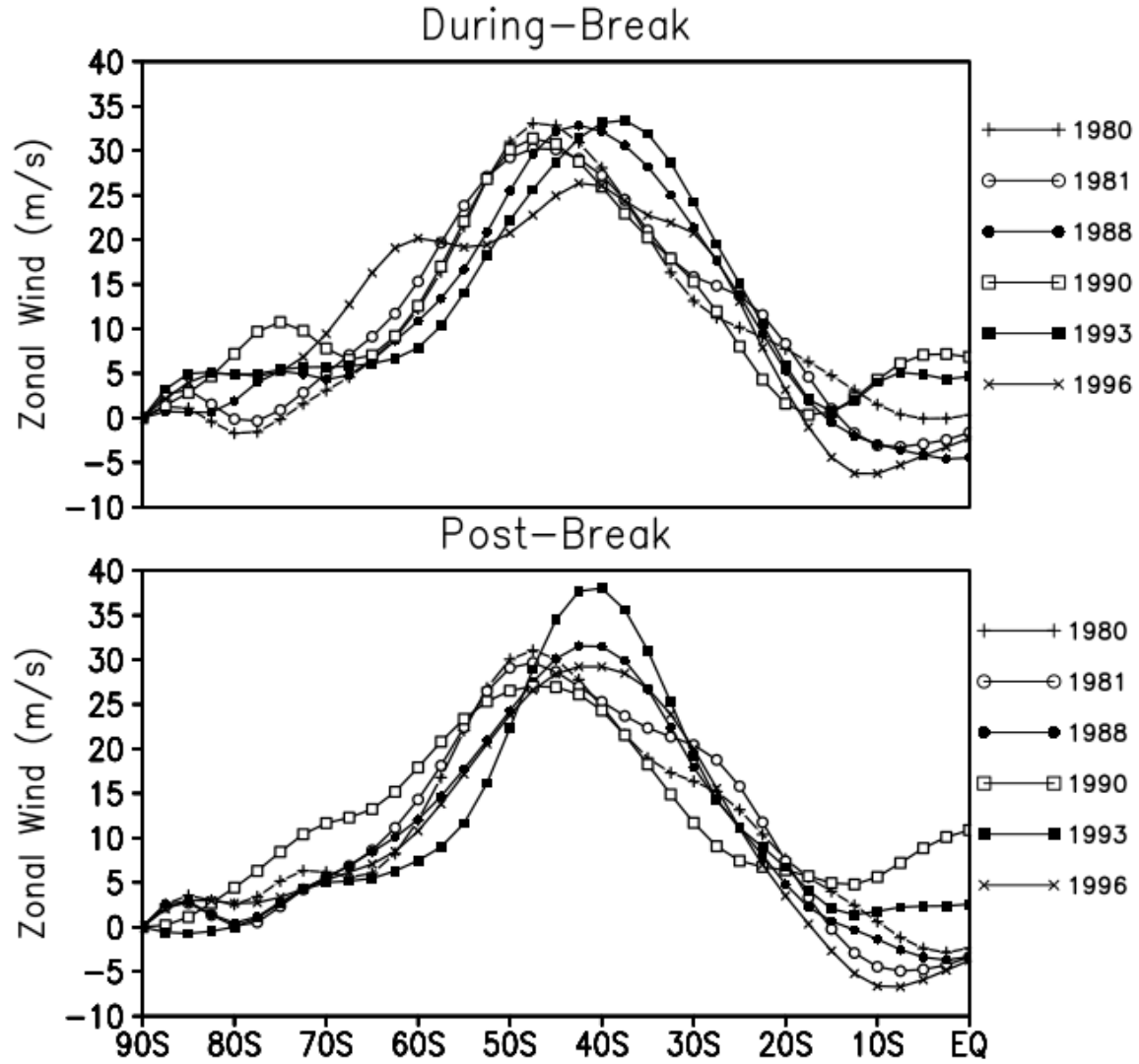
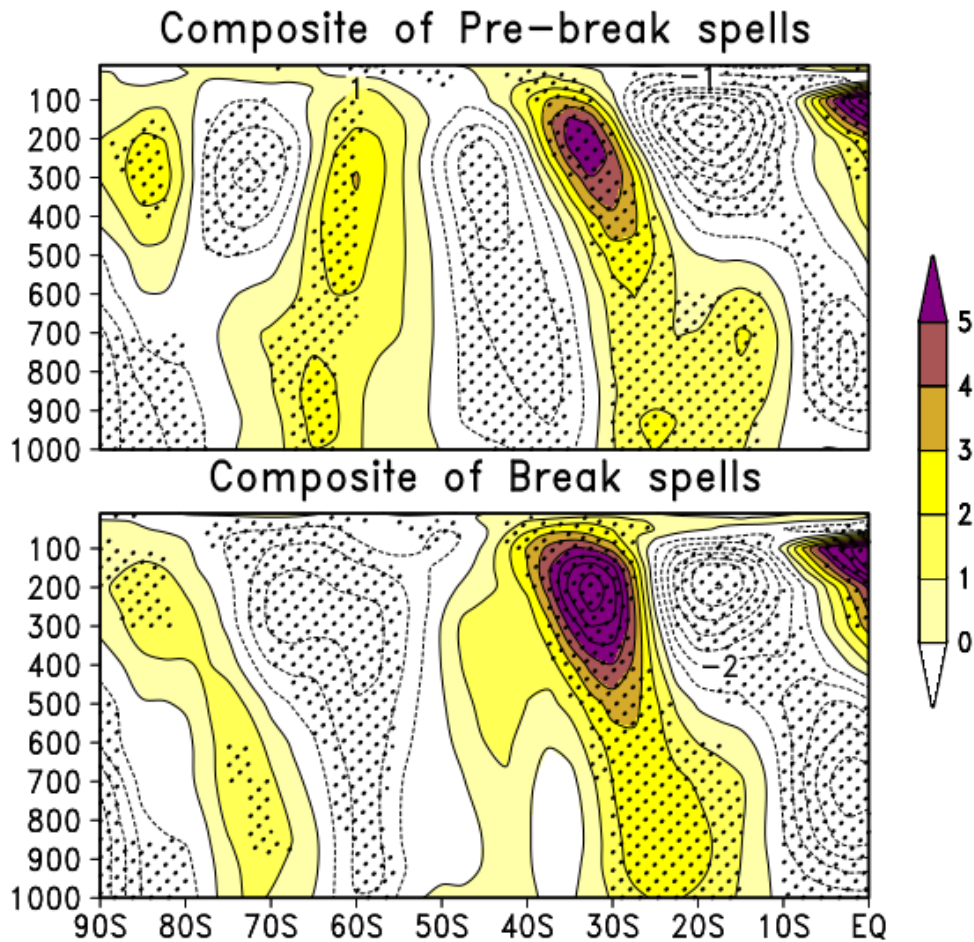


Figure 4.16: (a) (Top) Composite of globally zonal averaged U wind (m/s) for during-break in six different cases. **(b)** (Bottom) Globally zonal averaged U wind (m/s) for after-break periods in six different cases: 18-21 January 1980, 01-04 January 1981, 10-15 January 1988, 16-20 January 1990, 25 December-01 January 1993, 19-24 January 1996.

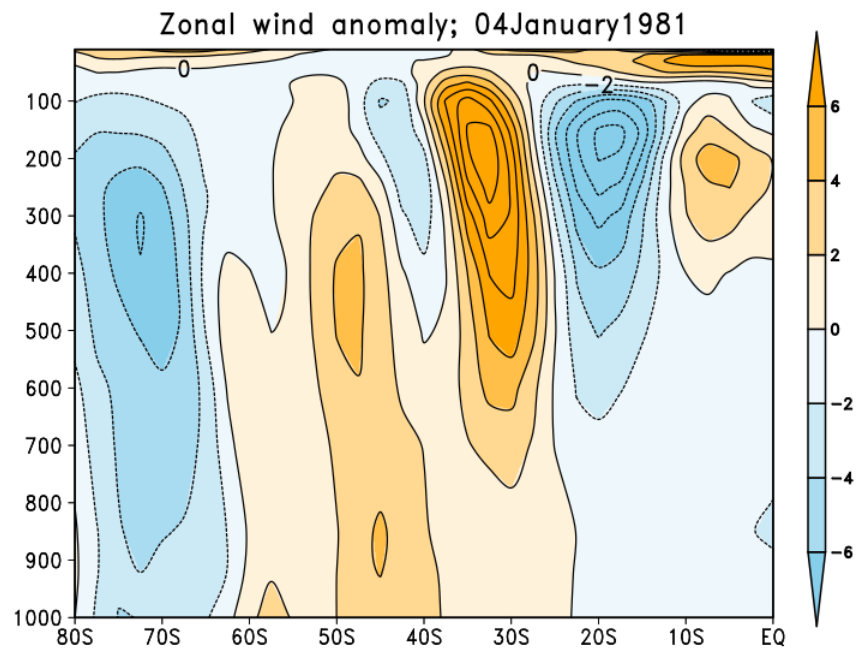
Fig. 4.17 shows the composite structure of zonal wind anomalies. Two jets, one at 55°S and the other at 35°S can be noted around the height of 250 hPa. Also, a tropical easterly jet can be seen

around 15°S at levels above 100 hPa. Similarly, Fig. 4.18 portrays the structure of zonal wind anomalies before the block formation on 4 January 1981. Whereas, the shear between the jets is maximized on 6 January 1981, the day on which the break period starts, indeed the maximized shear can be seen in the composite of during-break (Fig. 4.17b). This occurrence of maximum zonal wind shear during the break is similar to that noted for the Indian case. Further, the strong shear also can be noted in the upper troposphere between the westerly and easterly jets. We noted two regions of convergence, one at 60°S and 400 hPa and the other at 30°S and 250 hPa. Similarly, for the two case studies in January 1980 and 1981 (Fig. 4.19), two large convergence regions are found with a slightly higher magnitude relative to composite EP fluxes during-break period (Fig. 4.20a) and climatology DJF for the 1979-2007 period (Fig. 4.20b). This situation can be seen as similar to that co-occurred over India, the low latitude convergence seems to reduce the zonal wind and maintain the block without being advected, by the zonal current.



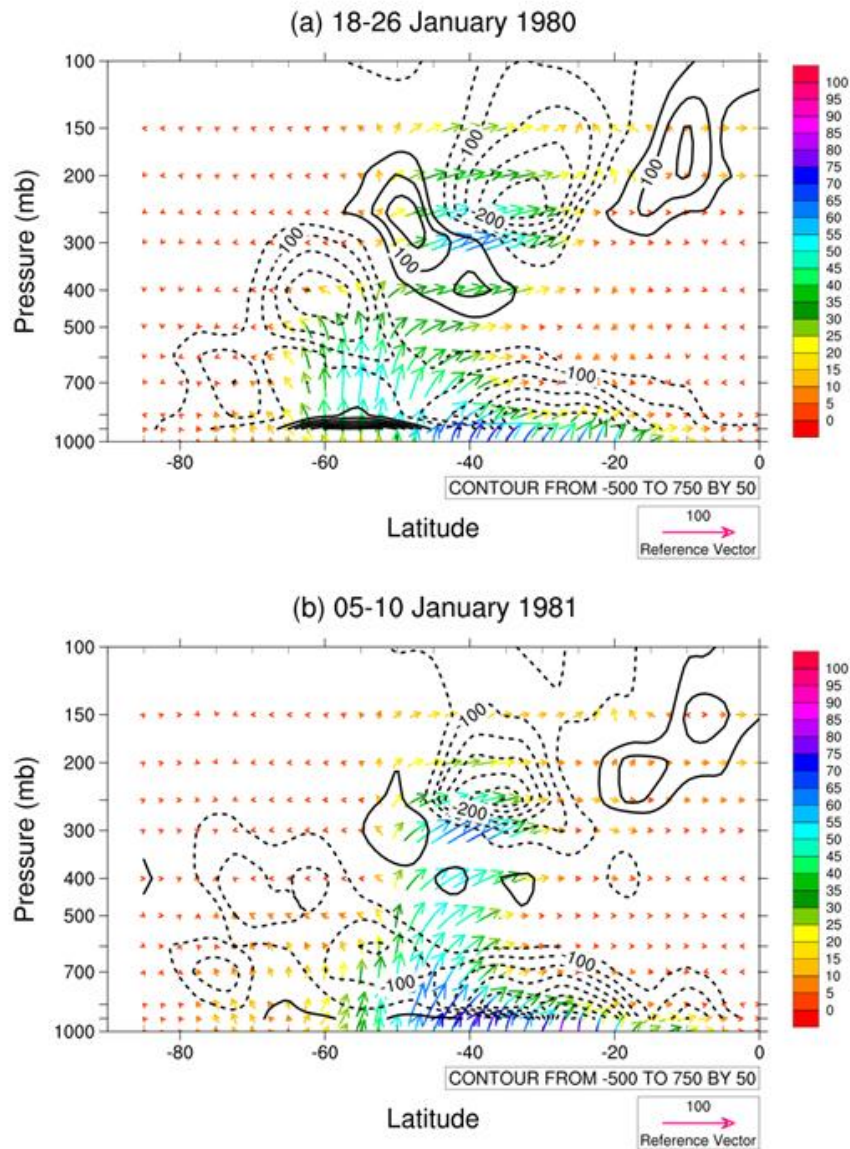
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2060 **Figure 4.17: (a)(Top)** Prebreak composite structure of U wind zonal anomalies of South American
 2061 monsoon (Table 3) (m/s), **(b)(Bottom)** During-break composite structure of zonal wind anomalies
 2062 of South American monsoon (m/s) (Table 3), which are zonally averaged between 60° W to 125°
 2063 W. The hatched region indicates 95% confidence using a two-tailed Student's t test. In both the
 2064 figures positive values are shaded and negative values are shown in contours. Contouring between
 2065 -5 to 0 (m/s) with an interval of 1 m/s.



2066

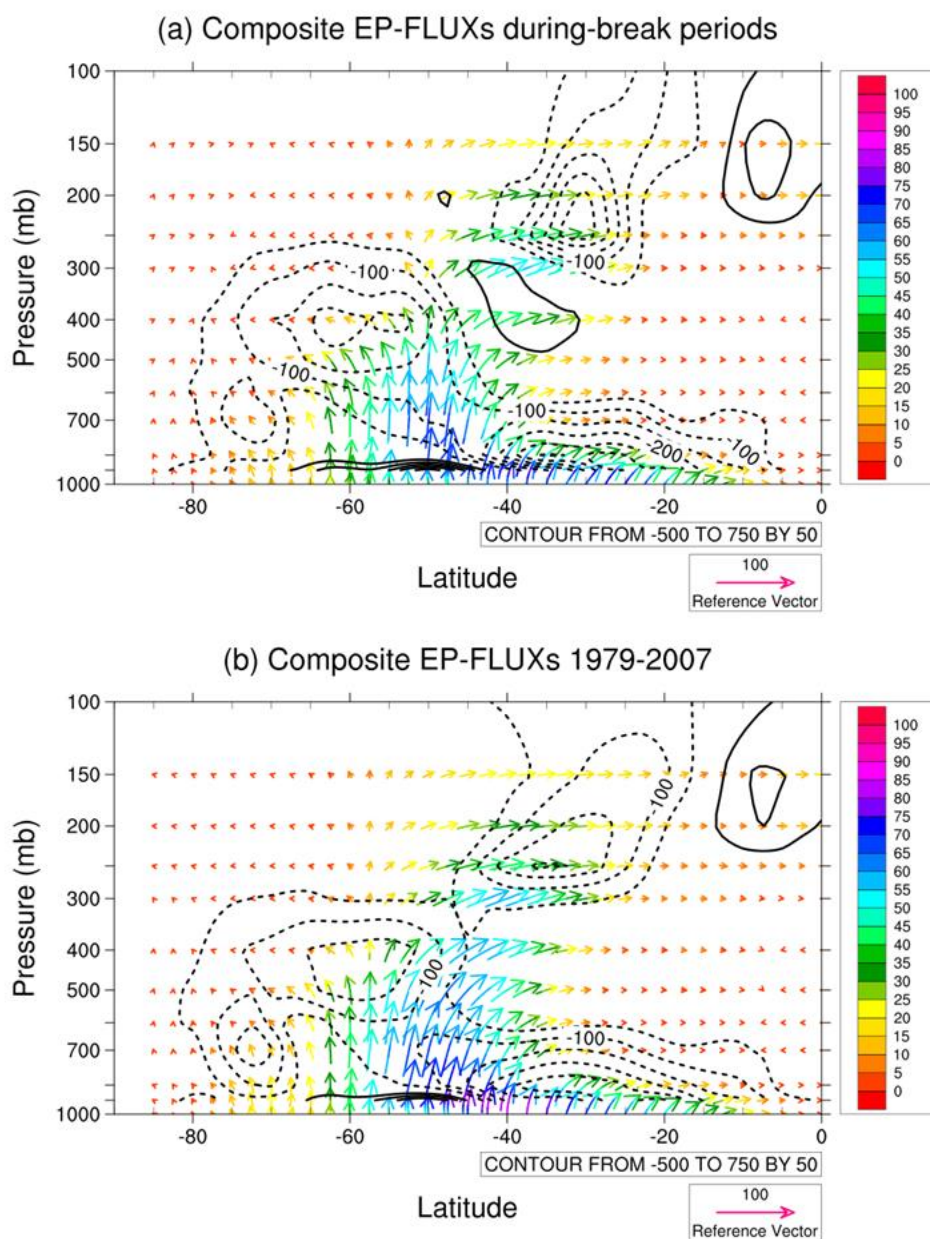
2067 **Figure 4.18:** Structure of zonal wind anomaly on 4 January 1981, which is zonally averaged over
 2068 270° E-330° E.



2069

2070 **Figure 4.19:** Eliassen Palm fluxes for (a) 18-26 July 1980 & (b) 05-10 July 1981. The dotted lines

2071 denote convergence and continues lines denote divergence and contouring at 50 m^3 .



2072

2073 **Figure 4.20:** Eliassen -Palm flux cross-sections of southern hemisphere
 2074 break periods (Table 2.4) of SAM; **(b)** during the season (austral summer) DJF for the 1979-2007
 2075 period. The contour interval is 50 m³. In both the figures dashed lines represent convergence and
 2076 continuous lines represent divergence.

4.4. Summary of the chapter

I observe a similar sequence of break monsoon characteristics as of the Indian summer monsoon, to those seen in the North and South American summer monsoon systems. These similar characteristics can be viewed as a global monsoon feature, in which extratropical eddies play a significant role in the common ground of the basic characteristics of the breaks in these three monsoons, namely, the ISM, SAM, and NAM. This is referred to as the unified theory.

The following chapter examines the revival of active conditions of rainfall after a break period over the core monsoon region of India during boreal summer.

Chapter 5

Understanding the revival of the Indian Summer Monsoon after Breaks

Outline of the chapter

In this chapter, I discuss a mechanism for reviving ISM after a typical break phase. This mechanism is associated with upperlevel barotropic instability caused by an increase of eddies in the mean zonal flow at 200 hPa, which spurs a subsequent synoptic low formation in the head of Bay of Bengal for the revival of ISM.

2119 **5.1 Introduction**

2120 In this chapter, I discuss the revival of active conditions of ISM after a break phase is
2121 facilitated by the formation of synoptic disturbances in the BoB, namely, monsoon depressions,
2122 and low pressure systems that travel toward the northwest from Bay of Bengal into the Indian
2123 region. For example, some of the previous studies discussing about the synoptic disturbance in the
2124 BoB for the active conditions over Indian region during boreal summer, Chen et al. 2005; Sikka
2125 and Dixit 1972; Boos et al. 2015; Sikka and Gadgil 1980.

2126 From a dynamical standpoint, for example, Ramaswamy (1962) emphasise that the relevance of
2127 anomalous southward shift of large-amplitude westerly trough from the midlatitudes into the Indo-
2128 Pakistan region during ISM breaks. Importantly in a case study by Rao (1971), also documents a
2129 manifestation of barotropic instability associated with increased horizontal shear due to the
2130 southward shift of the westerly troughs in the subtropical westerly jet at the midtropospheric level
2131 in the aforementioned break event and a subsequent revival associated with the formation of a
2132 synoptic disturbance. Rao (1971) hypothesized that manifestation of the barotropic instability
2133 during break leads to the formation of disturbances that in turn invigorate the ISM an active phase.
2134 In a recent study, Krishnamurthy and Ajayamohan (2010) have shown that the absence of low
2135 pressure systems, such as lows, depressions, and cyclonic storms, represents the break phase and
2136 their presence represents an active phase of ISM. Therefore, the question remains whether
2137 instabilities generated by large-scale processes lead to subsequent revival of the monsoon through
2138 a barotropic instability mechanism and formation of a synoptic disturbance.

2139 In this chapter, I attempt to answer this question. The availability of reanalysis datasets in the
2140 recent decades is a great opportunity in this sense. Analysis of multiple cases will also help us to

refine any theoretically based thresholds and indices that represent a phenomenon. For example, theory (Kuo 1953; Starr and White 1954; Aihara 1959) suggests that barotropic instability occurs only in disturbances of very long wavelengths. The case study of a break monsoon Rao (1971) suggests that synoptic waves in the subtropical westerly jet in the Indian region with a wavelength greater than (less than) 3000 km are unstable (stable). We revisit this aspect in this study. We present our results and a discussion in section 5.2 & 5.3, followed by conclusions in section 5.4.

5.2 Barotropic instability in the aftermath of breaks

From the works of Starr and White (1954) and Rao (1971), I can suppose that such a break condition will result in barotropic instability, which may in turn manifest as a synoptic disturbance for the revival of ISM. In this context, from Table 2.2, following Rajeevan et al. (2008), we list the dates of various postbreak revival events of ISM. Of the 41 total break events (Table 5.1), 18 revivals occurred with the formation of low pressure in the Bay of Bengal (e.g., Fig. 5.1) and 7 others with the formation of low pressure on land (figures not shown). This result suggests that about 61% of the postbreak revivals are associated with formation of low pressure in the Bay of Bengal or land regions, providing a general support to the hypothesis of Rao (1971) and Raghavan (1973).

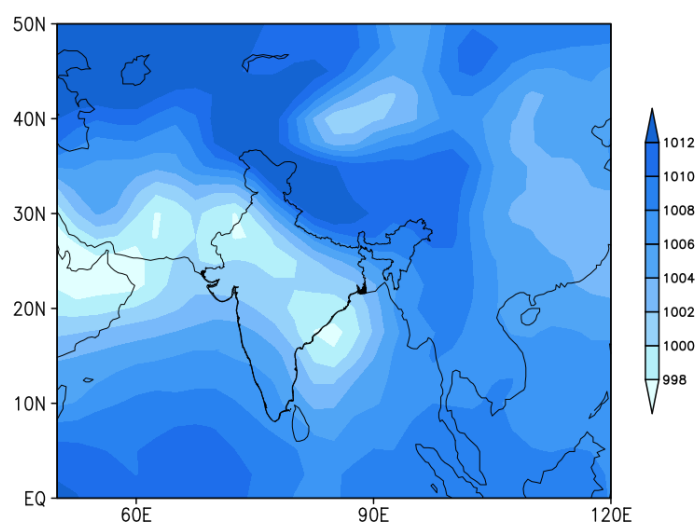


Figure 5.1: Observed Sea Level Pressure (SLP) distribution on 10 Aug., 2000, after a break period.

Table 5.1: Gives the formation of a synoptic disturbance on the particular day after every break event during the 1979-2007 period. '***' represents the revival of ISM without low formation.

Year	Low formation day after every break event	Condition
1979		***
1979		***
1980	23 July	Bay of Bengal
1980	26 August	Bay of Bengal
1981		***
1982	19 July	Bay of Bengal
1983	4 August	Bay of Bengal

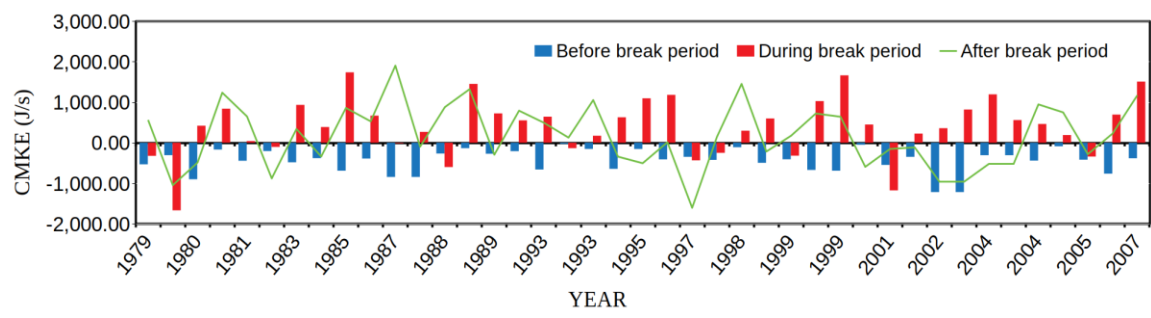
1984	30 July	Bay of Bengal
1985	28 August	Bay of Bengal
1986	9 September	Bay of Bengal
1987	11 August	Land Region
1987	19 August	Bay of Bengal
1988		***
1989	2 August	Bay of Bengal
1989	16 August	Bay of Bengal
1992	18 July	Land Region
1993		***
1993		***
1993	6 September	Bay of Bengal
1995		***
1995		***
1996		***
1997	22 July	Bay of Bengal
1997	16 August	Bay of Bengal
1998	1 September	Land Region
1998	29 July	Land Region
1999		***
1999		***
1999		***
2000	10 August	Bay of Bengal

2001	6 August	Bay of Bengal
2001	11 September	Land Region
2002		***
2002	1 August	Bay of Bengal
2004		***
2004	26 July	Land Region
2004		***
2005	11 September	Bay of Bengal
2005	19 August	Land Region
2007		***
2007	18 August	Bay of Bengal

2164

2165 Now, eddy formation due to barotropic instability would necessitate a conversion of the into, as
2166 shown by equation (2.3). Indeed, this is true in 30 out of the 41 cases (i.e., 73% of postbreak revival
2167 events), as evidenced by the positive values of rate of conversion of mean kinetic energy to eddy
2168 kinetic energy (CMKE) (Table 3) (Fig. 5.2). This indicates that the barotropic instability is the
2169 primary possible large-scale dynamical instability mechanism during the ISM breaks, many times
2170 leading to the formation of synoptic eddies. Another way to ascertain this further is by checking
2171 that there exists a significant negative correlation between the CMKE and wavelength, an
2172 indication of barotropic instability (e.g., Rao 1971). We find a strong correlation of -0.285 (Table
2173 5.2), which is significant at 95% confidence level from a student's two-tailed t test. This significant
2174 correlation confirms that barotropic instability is indeed manifested after the monsoon break events
2175 and is a necessary condition for the revival of Indian summer monsoon after break conditions.

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Figure 5.2: Conversion of Kinetic Energy (J/s) anomaly values of 41 break periods during the 1979-2007 period.

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Table 5.2: Conversion of Mean Kinetic Energy Values (Joule/second) at 200 hPa.

Year	Pre- break period	Break period	After-break period
1979	-533.85	-324.67	568.11
1979	-307.49	-1668.31	-1040.26
1980	-902.14	419.32	-483.19
1980	-168.76	838.84	1241.05
1981	-445.01	41.06	654.83
1982	-206.31	-104.81	-873.12
1983	-483.50	932.99	345.91
1984	-381.30	386.58	-334.05
1985	-691.33	1736.20	864.34
1986	-391.23	667.50	533.51
1987	-845.24	-15.09	1909.51

1987	-845.24	266.40	-92.27
1988	-267.84	-600.93	884.11
1989	-134.70	1450.52	1326.88
1989	-272.84	722.94	-289.01
1992	-209.65	551.18	796.84
1993	-662.03	641.99	500.70
1993	-27.15	-134.37	132.38
1993	-154.66	171.96	1059.52
1995	-645.48	627.13	-336.83
1995	-155.81	1094.81	-499.51
1996	-410.98	1179.61	18.84
1997	-348.92	-435.35	-1599.46
1997	-424.11	-249.73	147.87
1998	-111.57	296.10	1457.89
1998	-496.43	597.56	-211.52
1999	-409.16	-320.18	178.70
1999	-671.99	1026.46	727.02
1999	-691.03	1662.09	646.58
2000	-51.00	447.82	-588.90
2001	-547.99	-1176.18	-147.23
2001	-348.58	224.42	-108.66
2002	-1218.62	357.86	-950.32
2002	-1218.62	818.58	-950.32

2004	-309.27	1192.99	-520.30
2004	-309.27	558.93	-520.30
2004	-442.18	462.54	952.01
2005	-88.06	188.81	754.90
2005	-418.94	-341.93	-275.85
2007	-763.64	693.70	231.90
2007	-384.59	1507.19	1175.04

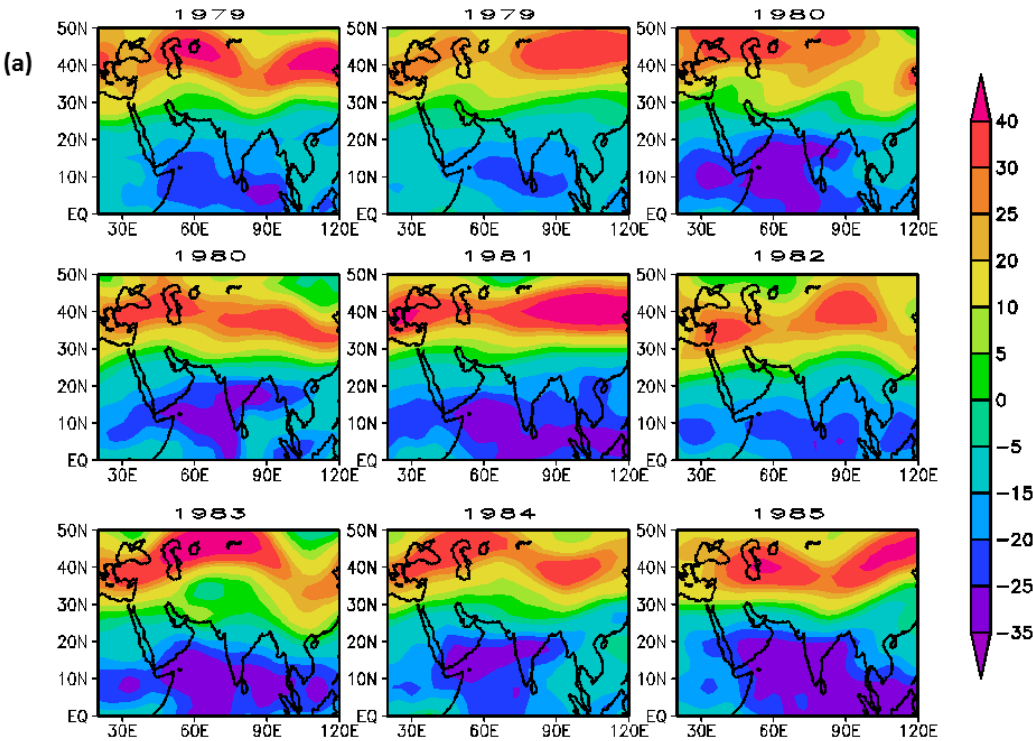
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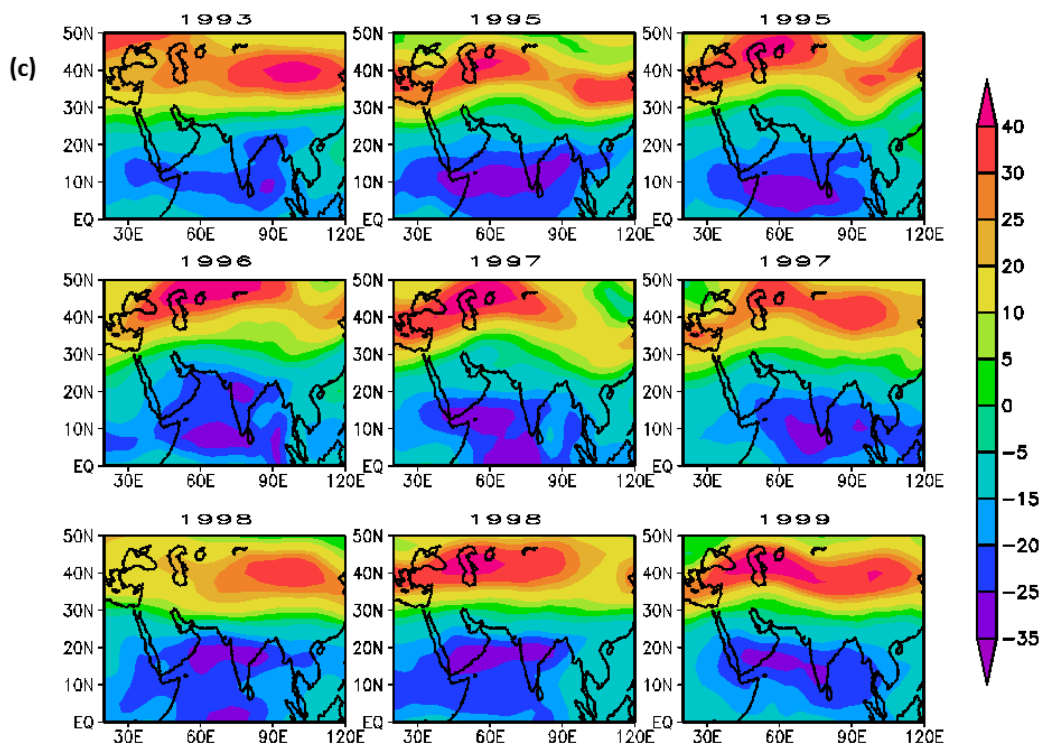
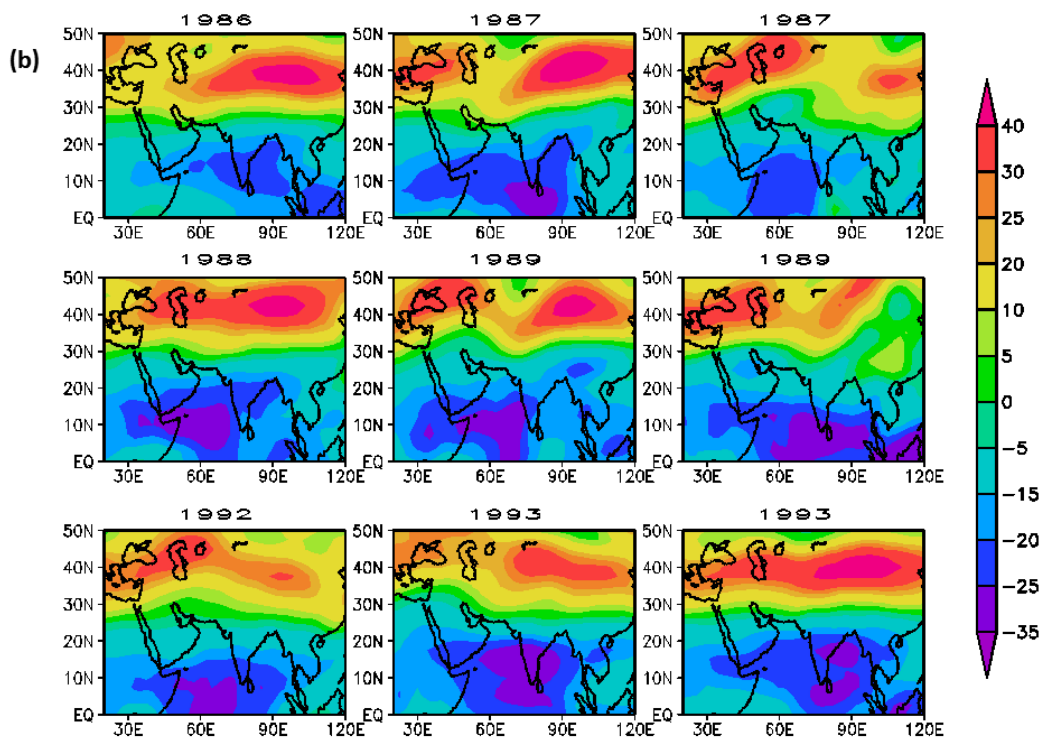
2184 What is the potential mechanism for such manifestation of barotropic instability in these
2185 subseasonal events? As is known, barotropic disturbances derive energy from the mean kinetic
2186 energy. Energy considerations (e.g., Kuo 1951) show that for a disturbance to grow, it must tilt in
2187 a direction opposite to that of the meridional gradient of zonal wind. To be specific, a tilt from
2188 southwest to northeast (SW–NE) in a westerly zonal flow will meet this criterion. That is, waves
2189 with an SW–NE tilt will result in a maximum vorticity to the south [see (3), which is from Kuo
2190 (1949)]. Fig. 5.3 (a-e) shows the zonal wind at 200 hPa of 41 break periods during the 1979-2007
2191 period and Fig. 5.4 (a-e) shows the zonal wind on a peak day of the 41 break periods for the 1979-
2192 2007 period. Whereas, figures 5.3 and 5.4, it is seen most of the break days are also indeed
2193 associated with such an SW–NE tilt in the 200-hPa zonal flow. Such a tilt in the mean 200-hPa
2194 subtropical westerly jet over the Indian region on a typical break day (see Fig. 5.5a as an example,
2195 along with the corresponding geopotential field in Fig. 5.5b) is associated with a northward transfer
2196 of westerly momentum (Kuo 1949). In such a case, the zonally averaged eddy momentum transport
2197 will be positive and is, importantly, conducive to the formation of an eddy disturbance (Fig. 5.6a)

2198 associated with maximum vorticity to its south (Kuo 1949). Truly, the corresponding zonal wind
 2199 structure at 200 hPa shows a southward shift of the westerly jet during the break period and a
 2200 northward shift of the tropical easterly jet (Fig. 5.6b).

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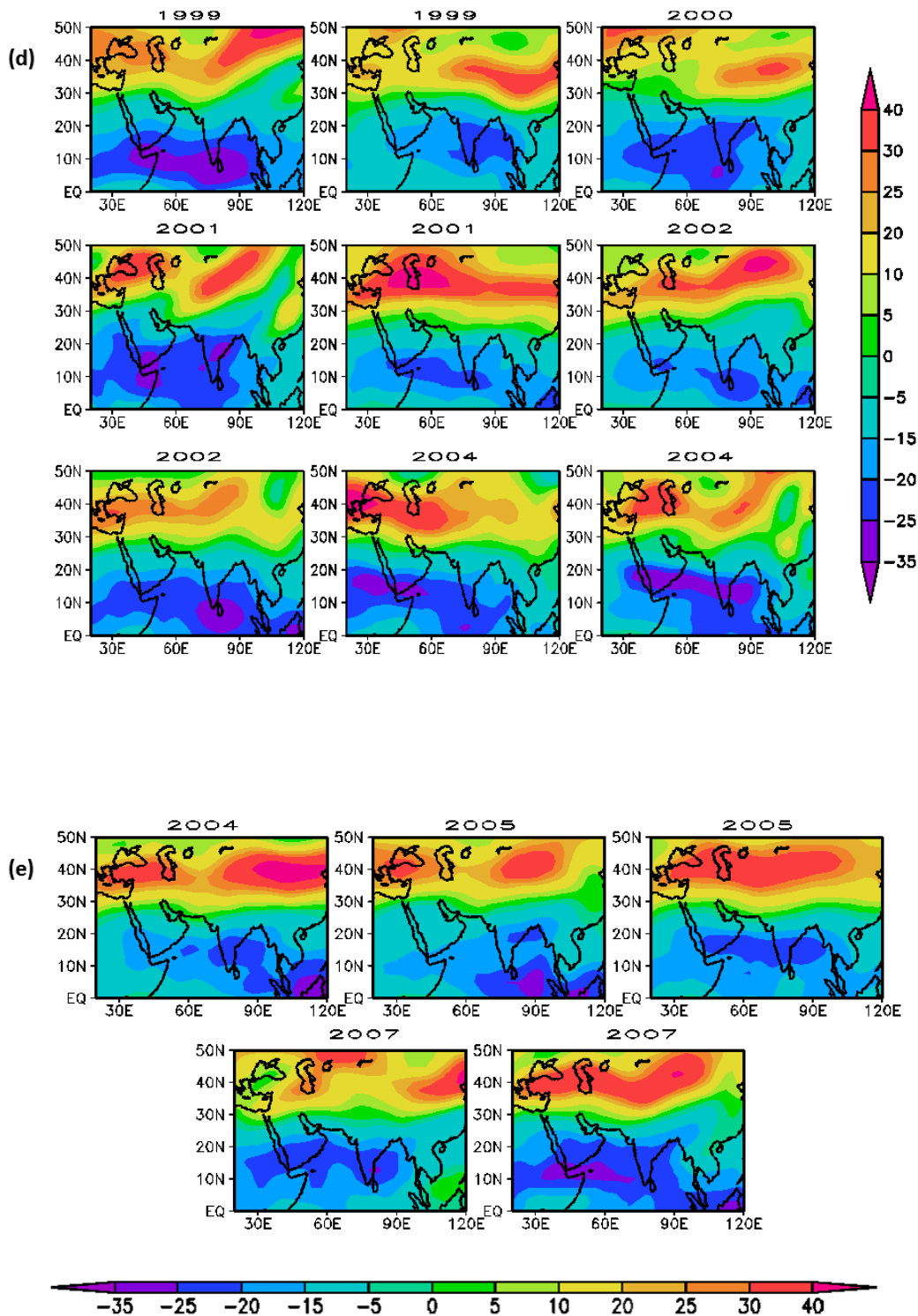
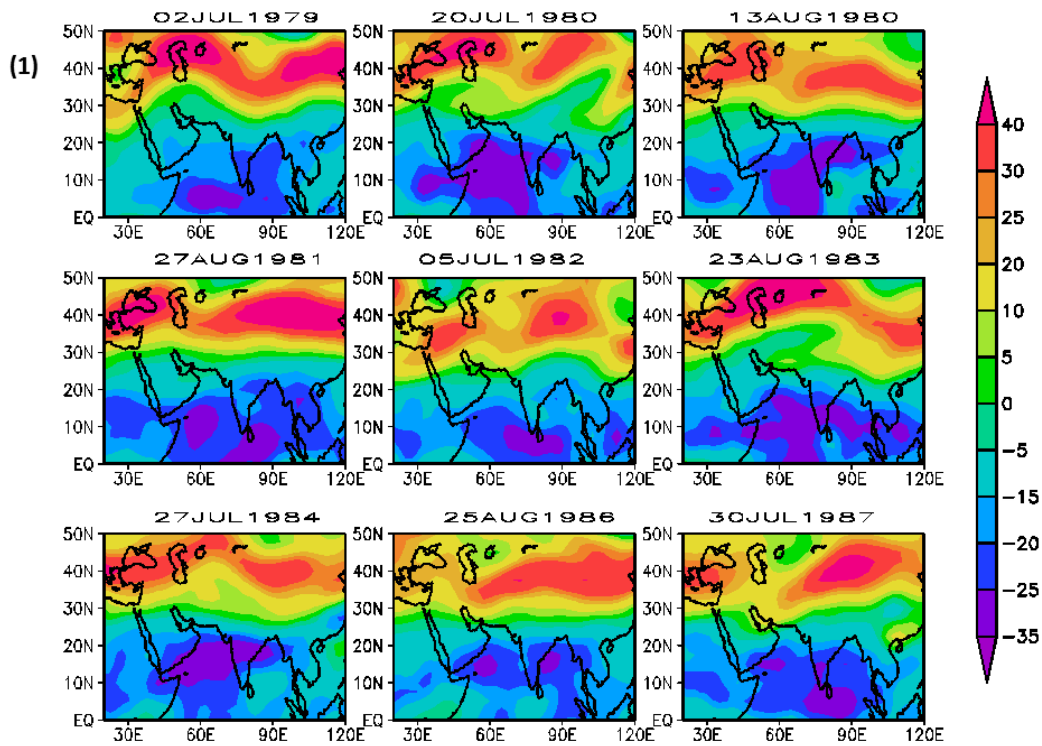
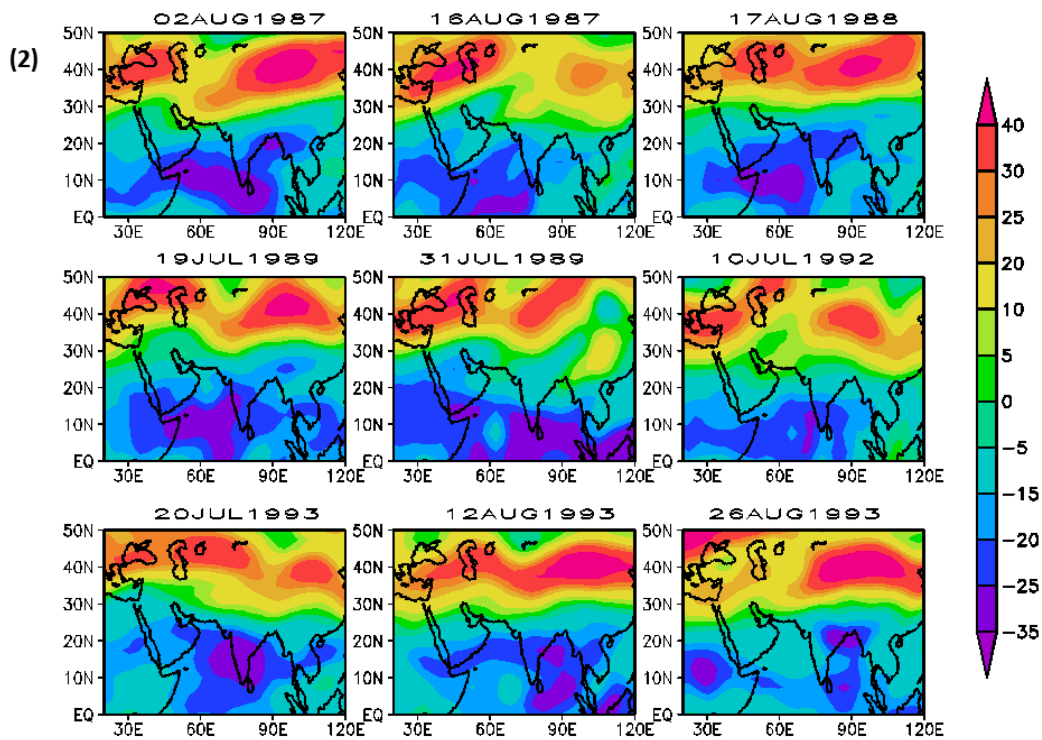


Fig. 5.3 (a-e): Spatial distribution of zonal wind at 200 hPa for 41 break events (Rajeevan et al. 2008) during the 1979–2007 period for the region 0°–50°N, 20°–120°E.

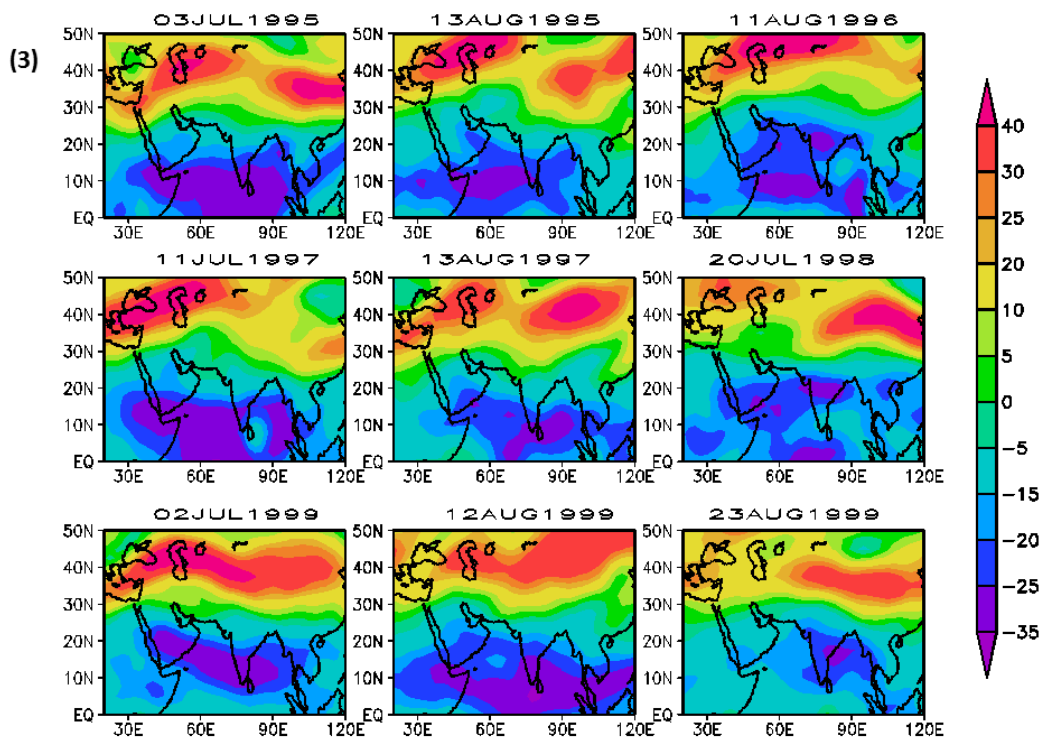


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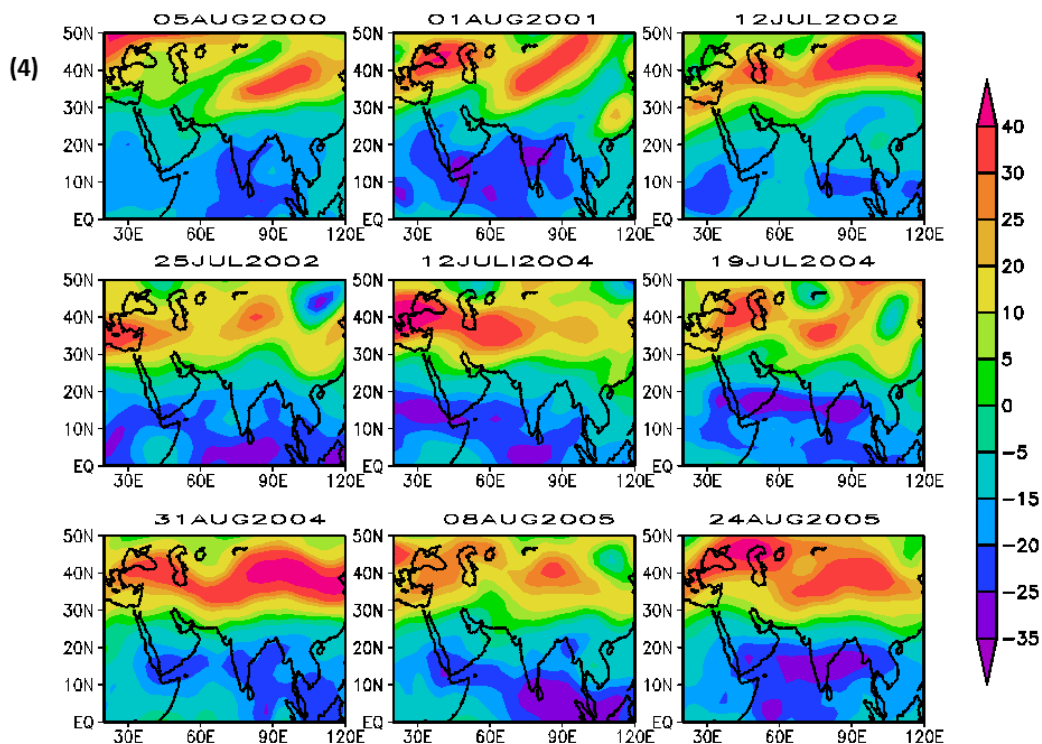
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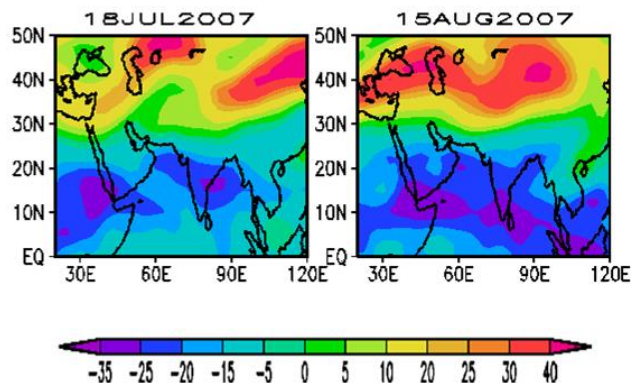


Fig. 5.4 (a-e): Spatial distribution of zonal wind on a peak days of of 41 break events (Rajeevan et al. 2008) during the 1979–2007 period for the region 0°–50°N, 20°–120°E.

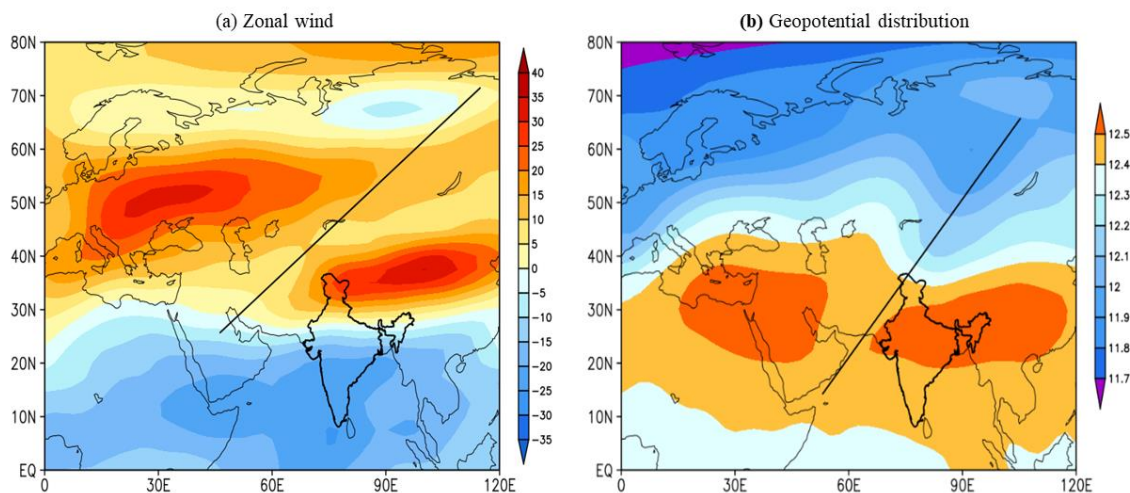


Figure 5.5: (a) Zonal wind at 200hpa on 4 August, 2000, a typical break day; **(b)** the corresponding Geopotential distribution (in Km). The black line shows the NE-SW tilt orientation in both the figures.

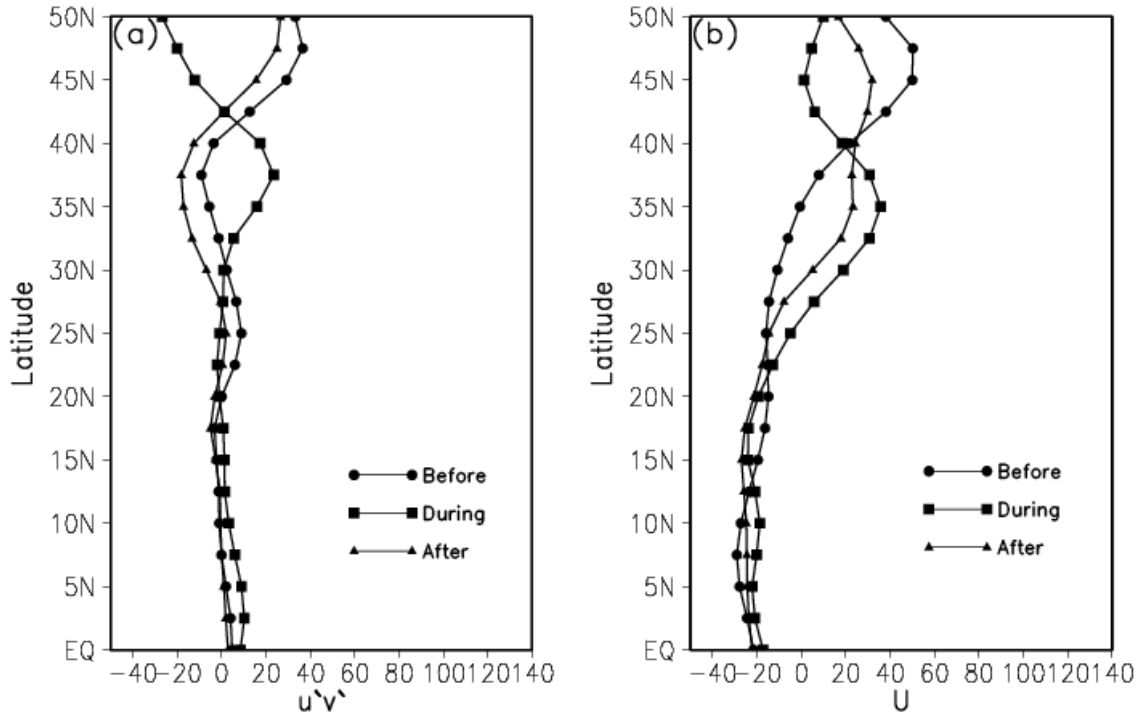


Figure 5.6: (a) Eddy momentum flux transfer during a break (1-9 August); before the break (14-23 July); and after the break (10-15 August), in relation to the 1-9 August break in the year 2000. **(b)** zonal wind profiles (82.5° E) at 200hpa during a break day (4th August); on a day in the prebreak period; (23rd July); on a after-break day (11th August), all for the same event presented in Fig. 5.6(a).

From the point of Rao (1971), it will be instructive to verify that the barotropic instability is a mechanism that would help the aforementioned eddies grow in such situations. To that end, the meridional vorticity distribution of the absolute vorticity ζ in the Indian region composite of during break events is presented in Fig. 5.7a, along with the corresponding all the break events in Fig. 5.7b. Importantly, we see maximum or minimum in absolute vorticity ζ around 29° N in the composite, with the individual values varying between 25° and 30° N. Manifestation of such maximum or minimum values is a necessary condition for the barotropic instability (Kuo 1951) from the individual case also indicates such manifestation (Fig. 5.7b). All this highlights the

importance of the mean seasonal zonal wind structure, with westerlies to the north and easterlies to the south of the Indian subcontinent, in facilitating such a dynamical instability manifested by the breaks.

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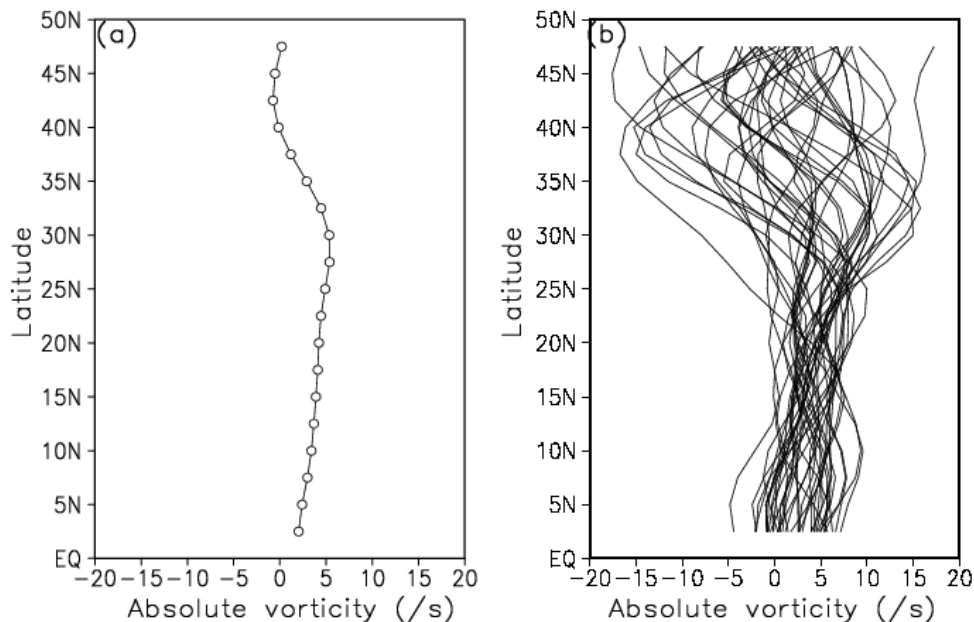


Figure 5.7: (a) Meridional distribution of the composite absolute vorticity (200 hPa), obtained by compositing it over all the break periods for the period of 1979-2007. **(b)** Same as Fig. 5.7(a) but composited over each break periods during 1979-2007.

5.3 Wavelength threshold for manifestation of a postbreak synoptic disturbance

Ramaswamy (1962) and Rao (1971) claim from their individual case studies a decrease in channel width ($D/2$) between subtropical westerly and tropical easterly jets that manifest as a dynamical instability. We revisit this aspect by computing the $D/2$ during the break events in the study period.

Our results (Table 5.3 and Fig. 5.8) show that 32 out of 41 break events (78%) indeed show a decrease in channel width. From this, we can deduce that a dynamical instability during the breaks is facilitated either as a result of a transient southward shift of the westerlies over the northern portions of the subcontinent and/or a transient northward shift of the tropical easterly jet stream over the peninsular region. Such a decrease in the channel width in the zonal width can also manifest with a weakening (strengthening) of the upper-level westerlies (easterlies) in the Indian region.

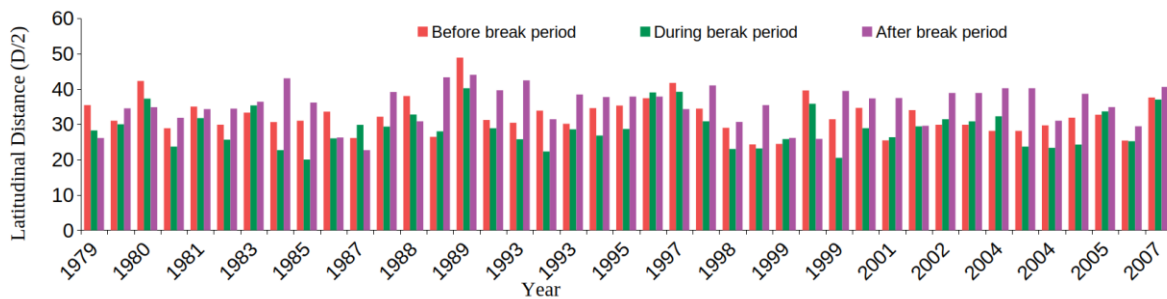
Table 5.3: The channel width ($D/2$) between the subtropical westerly jet and tropical easterly jet ($^{\circ}$) at 200 hPa for 41 break periods during the 1979-2007 period.

Year	Prebreak period	Break period	After-break period
1979	35.43	28.25	26.11
1979	31.00	30.00	34.50
1980	42.25	37.25	34.83
1980	28.88	23.67	31.83
1981	35.00	31.75	34.29
1982	29.88	25.63	34.43
1983	33.29	35.33	36.38
1984	30.63	22.67	43.00
1985	31.00	20.00	36.14
1986	33.57	26.00	26.25
1987	26.10	29.83	22.67

1987	32.14	29.33	39.14
1988	38.00	32.75	30.83
1989	26.43	28.00	43.29
1989	48.86	40.20	44.00
1992	31.22	28.88	39.63
1993	30.43	25.75	42.43
1993	33.86	22.29	31.43
1993	30.14	28.57	38.43
1995	34.57	26.80	37.71
1995	35.29	28.67	37.83
1996	37.38	39.00	37.83
1997	41.71	39.20	34.29
1997	34.43	30.83	41.00
1998	29.00	23.00	30.67
1998	24.29	23.14	35.43
1999	24.43	25.80	26.14
1999	39.57	35.80	25.86
1999	31.43	20.50	39.43
2000	34.63	28.89	37.33
2001	25.43	26.33	37.43
2001	34.00	29.40	29.57
2002	29.86	31.43	38.86
2002	29.86	30.82	38.86

2004	28.13	32.25	40.20
2004	28.13	23.67	40.2
2004	29.70	23.33	31.00
2005	31.86	24.25	38.63
2005	32.71	33.63	34.86
2007	25.38	25.20	29.43
2007	37.57	37.00	40.57

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2276

2277 **Figure 5.8:** Latitudinal distance (degrees) between westerlies and easterlies at 200 hPa for various
2278 phases associated with observed break events.

2279

2280 Theory (Kuo 1953; Syōno and Aihara 1957) shows that barotropic instability occurs only in zonal
2281 waves of wavelength shorter than a critical wavelength L_c [see chapter 2]. Rao (1971), from his
2282 sole case study, estimates L_c of the upper-level westerly jet stream in the Indian region to be ~3000
2283 km. However, given that it was only a single case and the relatively poor quality of the upper-air
2284 data during that period, we use the reanalyzed gridded datasets for multiple monsoon break cases
2285 to revisit this important finding by Rao (1971). Our analysis using (1) (Table 5.4) shows that (i)
2286 wavelengths in the upper-level westerlies north of Indian region during the summer monsoon reach
2287 a minimum value during breaks as compared to a few days prior to and after the event and (ii) the

critical mean value of the aforementioned wavelength, obtained by averaging it over all break events, is 7411 km. The minimum L_c we find is just 5127 km (Fig. 5.9).

Table 5.4: Critical wavelength (L_c Km) for prebreak, break and after-break periods.

Year	Prebreak period	Break period	After-break period
1979	9082	7310	6693
1979	7947	7690	8844
1980	10831	9549	8929
1980	7402	6067	8160
1981	8972	8139	8789
1982	7658	6569	8826
1983	8533	9057	9324
1984	7851	5810	11023
1985	7947	5127	9265
1986	8606	6665	6729
1987	6691	7648	5810
1987	8240	7519	10034
1988	9741	8395	7904
1989	6775	7178	11096
1989	12524	10305	11279
1992	8004	7402	10158

1993	7800	6601	10876
1993	8679	5713	8057
1993	7727	7324	9851
1995	8862	6870	9668
1995	9045	7349	9698
1996	9581	9997	9698
1997	10693	10049	8789
1997	8826	7904	10510
1998	7434	5896	7861
1998	6225	5933	9082
1999	6262	6614	6702
1999	10144	9177	6628
1999	8057	5255	10107
2000	8876	7405	9570
2001	6518	6750	9595
2001	8716	7536	7580
2002	7654	8057	9961
2002	7654	7900	9961
2004	7210	8267	10305
2004	7210	6067	10305
2004	7613	5981	7947
2005	8166	6216	9901
2005	8386	8620	8935

2007	6505	6460	7544
2007	9631	9485	10400

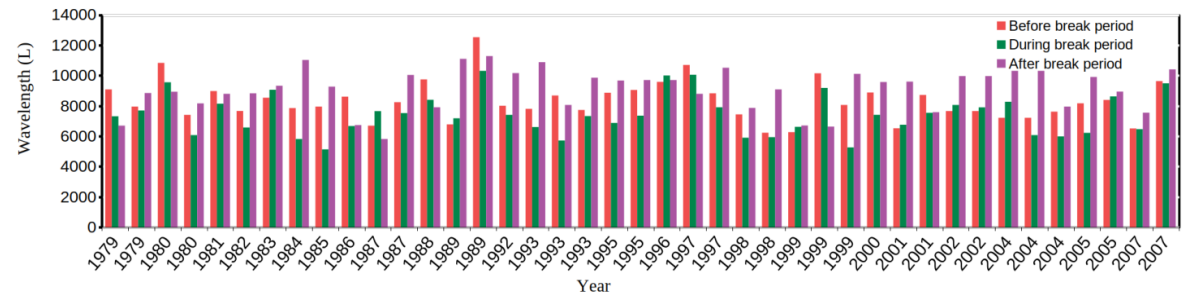


Figure 5.9: Wavelength anomalies (Km) of the zonal wind at 200 hPa, averaged over the span of each break event

5.4 Summary of the chapter

In this chapter, using the atmospheric circulation datasets from the NCEP–NCAR Reanalysis 2 (Kanamitsu et al. 2002) for the period 1979–2007, I explore the potential role of monsoon break conditions in subsequent revival of the monsoon through formation of a synoptic disturbance in the Indian region. I find that barotropic instability manifests in the Indian region during monsoon breaks in 61% of the cases. Such a revival is found to be associated with a reduction of the zonal width between the upper-level subtropical westerlies and tropical easterlies. Further, the anomaly correlation between the wavelength of zonal winds in the Indian region and the local rate of conversion of mean kinetic energy values for the study period is -0.285 , statistically significant at 95% confidence level, which confirms the role of barotropic instability for formation of the post break synoptic disturbance. During the monsoon break period there is no

2311 rainfall over most of the country, and therefore the succeeding disturbances are not generated by
2312 the condensational heating. Thus, the argument that generation of monsoon depressions and
2313 synoptic disturbances due to the break-induced barotropic instability is reasonable. I also find that
2314 the mean wavelength of westerlies during boreal monsoon events north of the Indian region, which
2315 leads to the revival of the monsoons, is about 7400 km. While Rao (1971) suggests a threshold
2316 wavelength of 3000 km from his single case study, my analysis of the 41 cases (adopted from
2317 Rajeevan et al. 2008) suggests an apparent threshold to be above 5000 km.

2318 The following chapter accesses the revival of active conditions of rainfall after a break period over
2319 the core monsoon region of India during boreal summer using the WRF model.

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Chapter 6

Assessment of a dynamical mechanism for the revival of active summer rainfall conditions over India following a break phase

Outline of the chapter

This chapter mainly focuses on various WRF model sensitivity experiments, which I have carried out to confirm the active role of the dynamical mechanism for the revival of active monsoon after a break phase, as proposed in Chapter 5. In this regard, we have carried out a variety of simulations to see the sensitivity of the simulated break to active transition to changes in the various meteorological parameters.

6.1 Introduction

Chapter 5 proposes, using nearly 30-years of reanalysis data, that, the monsoon revival after a break is caused by the barotropic instability of the zonal current. They found that during the peak of monsoon break, the tropical Easterly Jet, which is normally located around 9° N, invariably shifts northward. Also, the Subtropical Westerly Jet normally located around 40° N, shifts southward. These meridional shifts in the position of the jet streams increase the zonal wind shear, which results in the generation of barotropic instability. This instability results in the formation of a synoptic disturbance at least in about 61% of the cases, thereby reviving to the active monsoon phase.

To ascertain that the formation of barotropic instability associated with the closing of jet streams is indeed important for the subsequent revival of the monsoon, we have carried out few sensitivity experiments using the Weather Research and Forecasting (WRF) model version 3.8 (Skamarock et al., 2008) for a typical break case. The results are discussed in this chapter. We also verify the role of convection is also important in this process.

In the next section 6.2, I provide a description of the model configuration detailed experimental setup, and other datasets used in this study. I discuss the experimental results in section 6.3 and finally provide conclusions in the last section 6.4.

6.2 WRF model description and sensitivity experiments performed

The WRF model version 3.8 is a mesoscale numerical weather prediction system. It has both hydrostatic and non-hydrostatic options with the latter needing a fully compressible condition. The model has a sigma coordinate system designed for atmospheric research and operational forecasting needs. Several studies have been used this regional model to study the monsoon

problems from weather scale (e.g., Srinivas et al., 2017; Mohanty, Rao, 2014), intra-seasonal variability of ISM (e.g., Taraphdar et al. 2010; Kolusu et al., 2014; Chen et al., 2018), and even the relevance of interannual variations and background changes for extreme rainfall events (e.g., Boyaj et al., 2017, 2020). Various authors used the WRF model to successfully simulate and diagnose the break monsoon over the Indian region. For example, Taraphdar et al., (2010) used the WRF model to estimate the predictability of active and break phases of ISM from 2001 to 2009. Taraphdar et al., (2010) reported that the predictability is more for the break phases than the active phases of the monsoon over the Indian region.

We configured the model with two nested domains at 60 and 20 km horizontal resolutions, and 30 vertical levels. The mother domain covers the large-scale region encompasses 20°S-60°N, 10°E-160°E, the and nested domain 0°-50°N, 40°E-120°E region (Fig. 6.1). The selected physics options follow those of Rao et al. (2014), Boyaj et al., (2020), given that the simulation is mainly for the Indian region. Some of the physics schemes chosen are shown in Table 6.1.

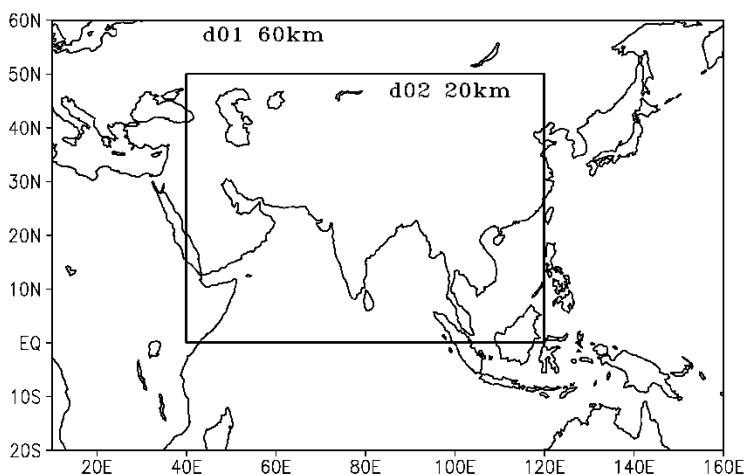


Figure 6.1: Schematic figure of the WRF model domain chosen, two nested domains of 60 and 20 km horizontal resolutions, and 30 vertical levels. The mother domain covers the large-scale

2393 region encompasses 20°S-60°N, 10°E-160°E, the and nested domain 0°-50°N, 40°E–120°E
 2394 region.

2395

2396 **Table 6.1:** List of physics schemes/parametrizations used in the sensitivity experiments of WRF
 2397 model

Sr. No.	Physics schemes	Citation
1	Goddard Ensemble scheme for microphysics	Tao et al., 2016
2	Short-wave radiation scheme	Dudhia et al., 1989
3	Rapid Radiation Transfer Model (RRTM) for long-wave radiation	Mlawer et al., 1997
4	Yonsei University (YSU) non-local scheme	Hong et al., 2006
5	Planetary boundary layer and NOAH scheme	Tewari et al., 2004
6	The Kain-Fritsch (KF) scheme	Kain, 2004

2398

2399 The initial and lateral boundary meteorological conditions for the model are obtained from the
 2400 National Centers for Environmental Prediction-Final Analysis (NCEP-FNL) data, available for
 2401 every six hours at 1°x1° horizontal resolution (NCEP, 2000). From the break case we selected,
 2402 from those suggested by Rajeevan et al. (2010), is from 1August to 9August, 2000 (hereafter
 2403 referred to as B2000).

2404 The experiments carried out in this study using the WRF model are tabulated below (Table 6.2).
 2405 We have carried out three experiments. In all the experiments, the model simulation spans from
 2406 01 to 15 August 2000, and the model initial boundary conditions start at 0000 UTC on 01 August
 2407 2000. In all the experiments, the lower boundary conditions are obtained from NCEP-FNL data.

The first experiment is what we refer is the control experiment (hereafter CTL). In this experiment, the lateral boundary conditions (LBC) for the outermost domain are adapted from the NCEP-FNL datasets. The second experiment is referred as the Daily Climatological Winds (hereafter DCW) experiments, respectively (Table 6.2). The DCW experiment is similar to that as the CTL, except that the LBCs are from the daily NCEP reanalysis-II climatological winds (generated 1980 to 2000). The third experiments are similar to DCW experiments, except that the convection scheme is switching off (DCWNC).

Table 6.2: A list of Experiments carried out using WRF model.

Sr. No.	Brief experiment details	Acronym
1	Real simulation using NCEP-FNL data as LBC's, from 31 July to 15 August 2000	CTL
2	Same as CTL but the experiment is forced with daily climatology of NCEP-II reanalysis zonal and meridional winds (1980-2000) as initial conditions.	DCW
3	Same as DCW experiment with the cumulus convection scheme switched off	DCWNC

6.3 Results and Discussion

6.3.1 Wavelength characteristics and energetics exchange for the revival of active conditions over India

We recall from chapter 5 that the critical wavelength (L_c) above which barotropic instability can occur during monsoon breaks is equal to or greater than ~ 5000 km. The L_c for the B2000 break

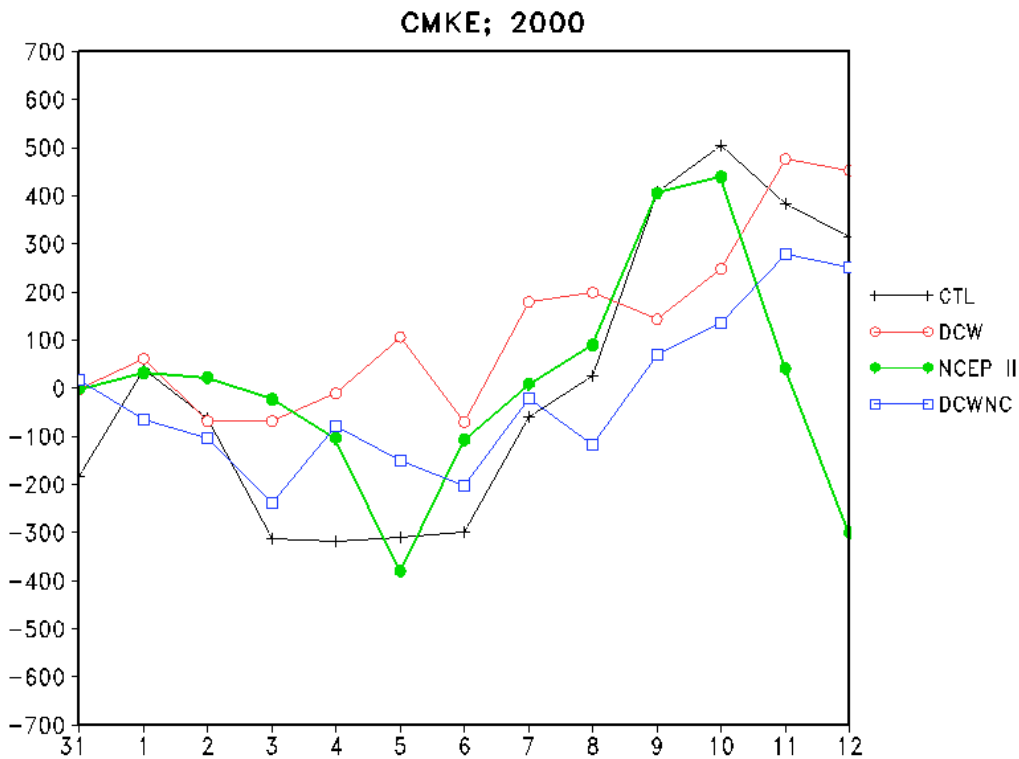
event from the NCEP/NCAR reanalysis II is 7405 km (Table 6.3). During the break, the average latitudinal distance between the two upper level jets was 30°. The corresponding values simulated in the CTL experiment with latitudinal distance of 30°, and L_c of 7690 km, which are well in agreement with the values from the NCEP-II reanalysis. The DRW experiment simulates a significant latitudinal distance between the SWJ and TEJ during the B2000 break event, particularly in comparison to NCEP II reanalysis data. Interestingly, the latitudinal distance simulated in the DRW is 24°. We believe that the relatively better results from the CTL simulation, compared to the NCEP II reanalysis data, whereas the FNL data sets are relatively coarser. Furthermore, all of the experiments show L_c values greater than the average (~5000 Km), indicating that there is a basis for barotropic instability at 200 hPa.

Table 6.3: Latitudinal distance ($D/2$) (°) and Critical wavelength (L_c) (Km) during the break period B2000 in the NCEP II reanalysis data and conducted sensitivity experiments.

Year	Lat. Dist. (°), B2000	Wavelength (Km), B2000
NCEP II reanalysis	29	7405
CTL	30	7690
DCW	32	8203
DCWNC	31	8175

As discussed so far, horizontal atmosphere, the barotropic instability develops due to an increase in horizontal shear of the zonal flow. From the context of Lorenz cycle (1955), this involves the conversion of mean kinetic energy to eddy kinetic energy of the zonal flow, a term we represent by C_k . In this context, in Fig. 6.2, I illustrate the evolution of the C_k at 200 hPa over the CMR

2444 through the break phase B2000 from the NCEP II reanalysis as well as that from all the simulation
 2445 experiments. Its evolution from the NCEP reanalysis suggests a significant fall in the C_k during 1-
 2446 5 August, 2000 in association with the peaking of the break, after which its magnitude increases;
 2447 it peaks on 8 August. The CTL experiment also broadly captures the general weakening phase and
 2448 its intensification later. The DCW and DCWNC experiments, which are forced with daily
 2449 climatological winds do not simulate the fall in the C_k well. On the other hand, Fig. 6.2 also shows
 2450 that the evolution of the C_k from the simulation is qualitatively reasonable.

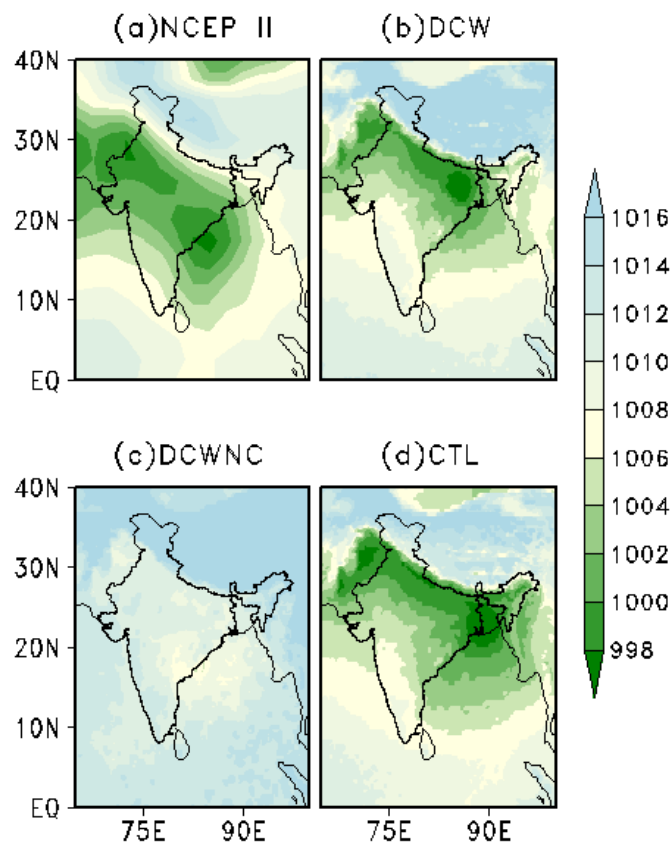


2451
 2452 **Figure 6.2:** Daily mean C_k , the rate of conversion of mean kinetic energy to eddy kinetic energy
 2453 (J/s) during break event from 31 July-09 August 2000 over core monsoon region of India, from
 2454 the NCEP II reanalysis data, and from the CTL, DCW, and DCWNC experiments. The abscissa
 2455 shows the dateline from 31st July-12th August 2000 and the ordinate shows the rate of conversion
 2456 of energy.

2457

2458 The Fig. 6.3a shows the SLP from the NCEP II reanalysis data on 10th August 2000, which is the
2459 day after the demise of the B2000 break event. In this figure, we see the formation of a low in the
2460 north BoB. It is widespread and close to the eastern Indian coast, and located in the southern
2461 portion of the MT, which is aligned along the CMR. The CTL experiment simulates the low (Fig.
2462 6.3d) with an intensity similar to that in Fig. 6.3a at the head of the BoB, with though the MT has
2463 been simulated slightly North of its observed position. However, the DCW experiment shows a
2464 weak low and MT (Fig. 6.3b). Despite DCWNC experiment (Figures 6.3c) do neither simulate the
2465 low nor the MT.

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Figure 6.3: Sea level pressure on 10 August 2000 over the Indian region from (a)NCEP II reanalysis data, (b) DCW, (c) DCWNC, and (d) CTL experiments

6.3.2 Eddy momentum flux and rainfall characteristics during a break phase.

The energy to grow the barotropic disturbances is derived from the mean kinetic energy. Energy considerations show that for a disturbance to grow, it must tilt in a direction opposite to that of the meridional gradient of zonal wind (Kuo 1949). That is, waves with an SW–NE tilt will result in maximum vorticity to the south (Kuo 1949), and are associated with a northward transfer of westerly momentum. In such a case, the zonally averaged eddy momentum transport will be positive and is, importantly, conducive to the formation of an eddy disturbance associated with maximum vorticity to its south (Kuo 1949). The transport of eddy momentum flux ($\overline{u'v'}$) is associated with large-scale eddies. Fig. 6.4 depicts the eddy momentum flux for the B2000 break event over the CMR from the NCEP II and all simulations. The evolutions of the simulated eddy momentum flux from the CTL experiment is comparable to that from the NCEP II (Fig. 6.4), though the magnitudes of the eddy momentum flux simulated by the DCWNC, DCW are much weaker.

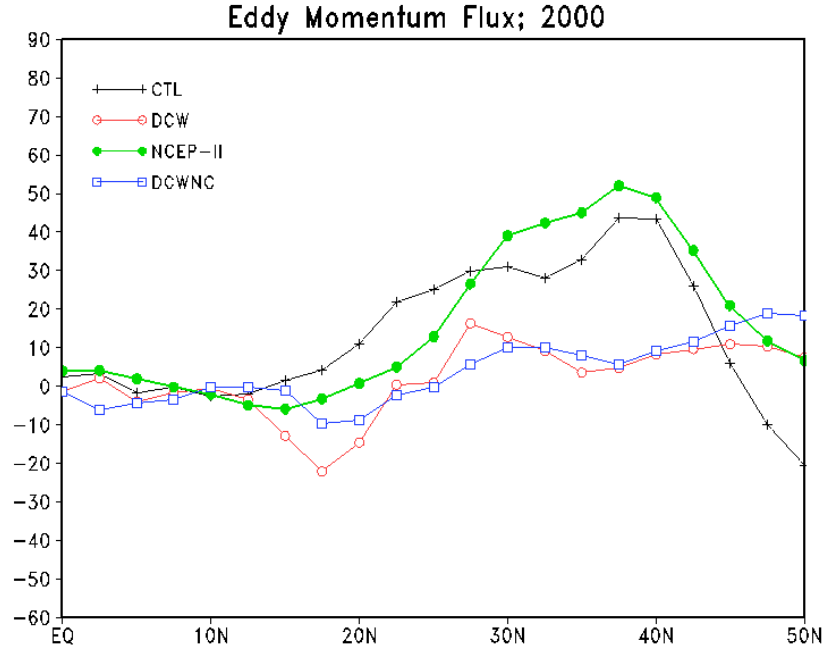


Figure 6.4: Evolution of the eddy momentum flux (m^2/s^2) through the break event 01-09 August 2000 over the core monsoon region of India using NCEP II, DCW, DCWNC, and CTL. The abscissa shows the dateline from 31st July-12th August 2000 and the ordinate shows the rate of conversion of energy.

As per Kuo (1949), the occurrence of neutral waves requires that the absolute vorticity (Z) of zonal flow possess an extreme value ($\frac{dz}{dy} \equiv \beta - U'' = 0$) at some point. This is the condition for barotropic instability (CBI) of the zonal Rossby waves. If no such critical point exists in the velocity profiles of waves, there can be no neutral waves with phase velocity in the range of the minimum and maximum zonal magnitude of the Rossby waves. Unstable waves manifest the growth of the eddies, which are associated with the transport of momentum flux depending upon the orientation of the troughs of vacillating flow. Kuo (1951) further extended the “*the study of wave motions from the very long and slowly moving or retrograding waves into the realm of*

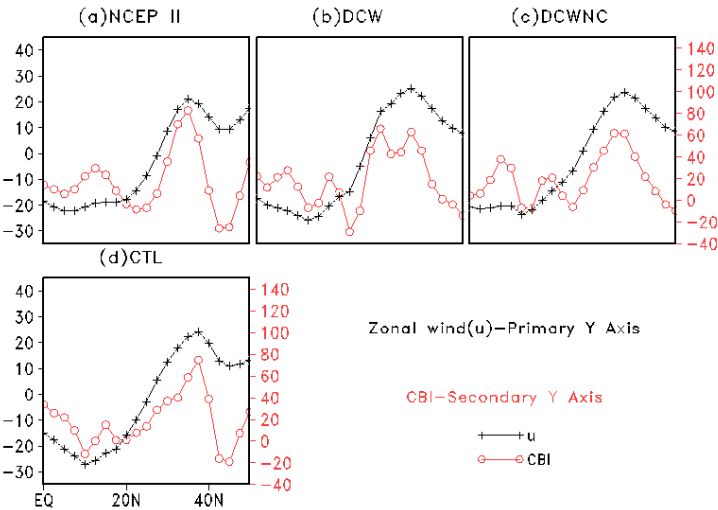
2498 ordinary waves and cyclone waves and it is found that, for nondivergent barotropic motion, the
2499 condition for the presence of neutral and amplified waves with a phase velocity whose value is
2500 between the maximum and minimum wind velocity in the belt is the existence of critical points
2501 where the absolute vorticity has an extreme value. If no such point exists, then all perturbations
2502 must be damped. When this condition is satisfied, both amplified (unstable) and neutral waves can
2503 be expected. The waves moving with a velocity equal to the current velocity at the critical point is
2504 neutral while those with a velocity less than this value but greater than the minimum wind velocity
2505 will be amplified. The amplification will be greatest when the phase velocity is the intermediate
2506 between the latter two values; therefore, both fast and slowly moving waves will have little
2507 amplification. The degree of instability will also depend upon the sharpness of the velocity profile.
2508 When the wave is unstable, the trough line will be directed from southeast toward northwest to the
2509 south of the point of the minimum absolute vorticity and from southwest toward northeast to the
2510 north of the point of maximum vorticity.”

2511 In this sense, Fig. 6.5 shows the zonal wind (u) and condition for barotropic instability (CBI)
2512 during a break event B2000. In chapter 5, I show the composite CBI or absolute vorticity (chapter
2513 5, Fig. 5.6 & Fig. 5.7) using NCEP II zonal wind data and declining meridional width between
2514 SWJ and TEJ. In comparison to NCEP II reanalysis data, CTL, DCW, and DCWNC experiments
2515 show a weak zonal wind magnitude on peak days of break phase, along with weaker vorticity.
2516 Importantly, the unstable waves are located to the north of the point of maximum vorticity and the
2517 westerly waves orient along NE-SW direction (Kuo 1951), indeed, the geopotential height
2518 distribution at 200 hPa on B2000 break event show the trough oriented along NE-SW direction
2519 (Fig.6.5). Fig. 6.6 shows the 200 hPa geopotential height distribution for the average of break
2520 period B2000 and Fig. 6.7 shows the 200 hPa geopotential height difference between the NCEP II

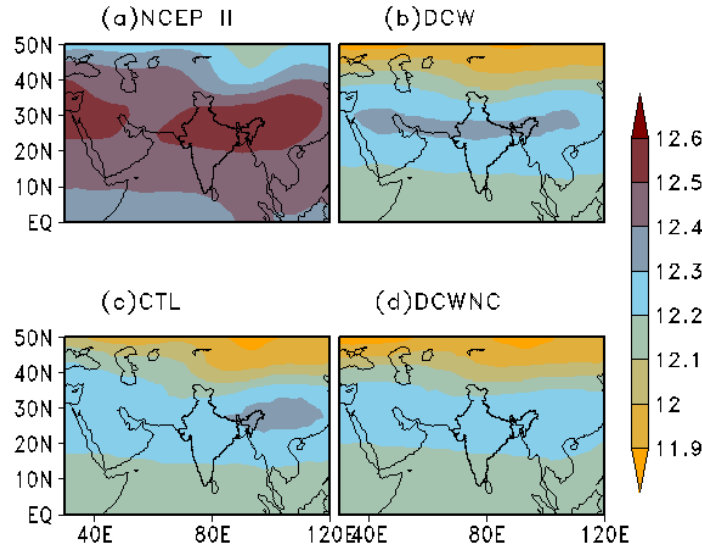
2521 and conducted sensitivity experiments, for the break B2000. The observed westerly jet trough
 2522 orientation along the NE-SW direction is seen in all the experiments with a minor spatial shift.

2523 Similarly, I note the trough NE-SW orientation in the zonal wind distribution at 200 hPa during
 2524 the break B2000 in all the sensitivity experiments conducted with a slight spatial shift and changes
 2525 in the magnitude to the SWJ (Fig. 6.8). Fig. 6.9 shows the zonal wind difference between the
 2526 reanalysis data and model experiment results. In comparison to the reanalysis data, the experiments
 2527 CTL, and DCW seem to show a good shifting of SWJ and TEJ towards each other during the
 2528 B2000 break event (Fig. 6.9). Notably, the experiments DCWNC show weaker magnitudes of
 2529 zonal winds and meridional distribution of SWJ and TEJ. In addition to the geopotential height
 2530 distribution, all the experiments of the zonal wind distribution at 200 hPa also demonstrate the
 2531 westerly jet trough orientation along the NE-SW direction (figures 6.8 & 6.9).

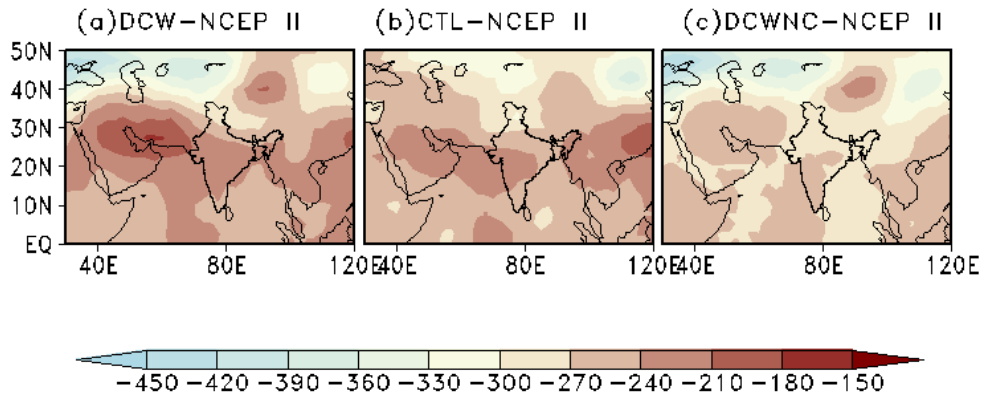
2532 Thus, all the model results show a weak rainfall situation over CMR. Fig. 6.10 shows the daily
 2533 rainfall over the CMR of India for the B2000 break event. Apart from rainfall IMD gridded data,
 2534 all the experiments show an overestimated daily rainfall during the break period B2000.



2536 **Figure 6.5:** Zonal wind (u; Black) and Condition for barotropic instability (CBI) (/s) for break
 2537 event 01-09 August 2000 over core monsoon region of India using (a)NCEP II, (b) DCW, (c)
 2538 DCWNC, (d) CTL, (e) DNW, and (f) DNWNC

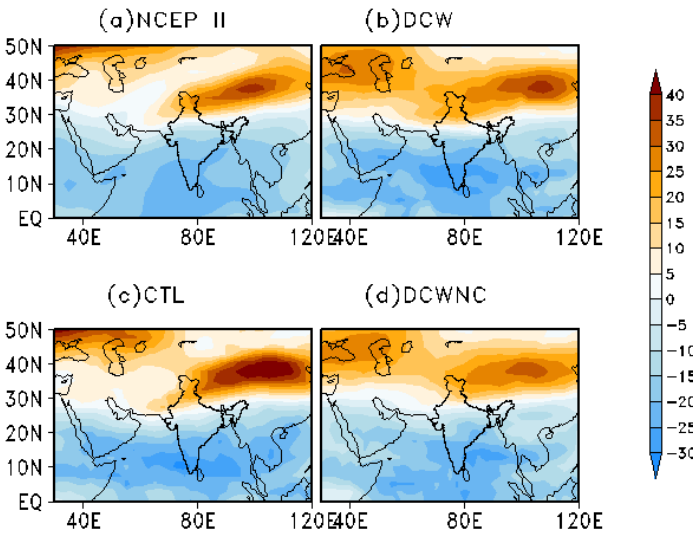


2539 **Figure 6.6:** Geopotential height (m) at 200 hPa on break event 01-09 August 2000 of (a) NCEP
 2540 II, (b) DCW, (c) DNW (d) CTL, (e) DCWNC, and (f) DNWNC

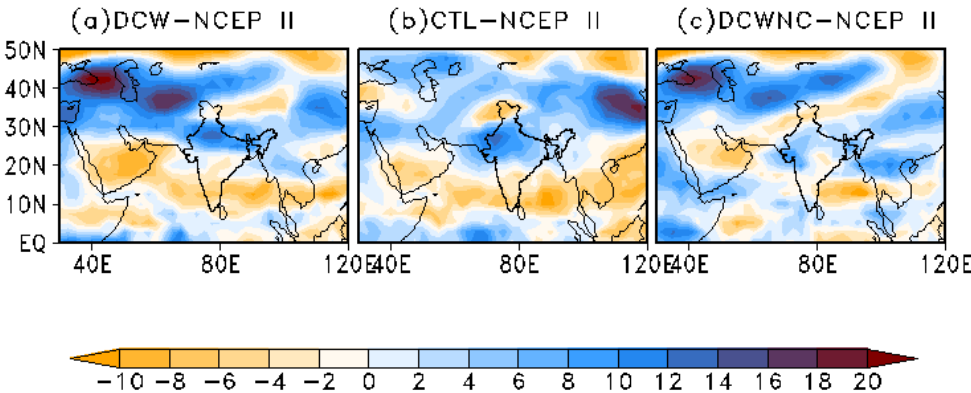


2542 **Figure 6.7:** Geopotential height (m) at 200 hPa on break event 01-09 August 2000 of (a) difference
 2543 between DCW-NCEP II, (b) difference between DNW-NCEP II, (c) difference between CTL-
 2544

2545 NCEP II, (d) difference between DCWNC-NCEP II, and (e) difference between the DNWNC-
 2546 NCEP II



2547
 2548 **Figure 6.8:** Zonal wind (m/s) at 200 hPa on break event 01-09 August 2000 of (a) NCEP II, (b)
 2549 DCW, (c) DNW (d) CTL, (e) DCWNC, and (f) DNWNC



2550
 2551 **Figure 6.9:** Zonal wind (m/s) at 200 hPa on break event 01-09 August 2000 of (a) difference
 2552 between DCW-NCEP II, (b) difference between DNW-NCEP II, (c) difference between CTL-
 2553 NCEP II, (d) difference between DCWNC-NCEP II, and (e) difference between the DNWNC-
 2554 NCEP II

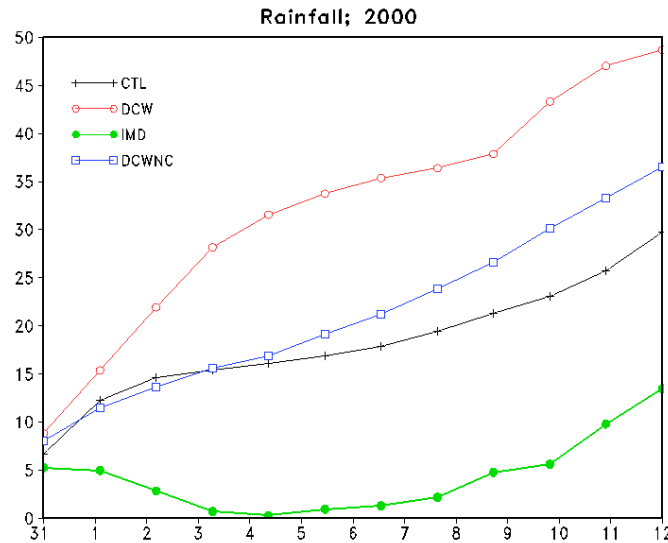


Figure 6.10: Daily Rainfall over core monsoon region of India, using NCEP II, DCW, DNW, CTL, DCWNC, and DNWNC. The abscissa shows the dateline from 31st July-12th August 2000 and the ordinate shows the rate of conversion of energy.

6.4 Summary of the chapter

As the study on monsoon break case i.e., 01-09 August 2000, is assessed using the WRF model to ascertain the changes in the upper atmosphere during a break phase for the revival of summer monsoon over India. The primary goal of the study is to see the characteristic changes in the zonal flow of the upper atmosphere with the conditions variable initial conditions provided in the WRF model, which is the key component to the development of dynamical shear at 200 hPa and subsequently followed by the formation of a low in the head of the Bay of Bengal for the revival of ISM.

Our study suggests the experimental results have shown a qualitative similarity to the NCEP II reanalysis data. The experiment CTL (real run of the model using NCEP-FNL as LBCs), have proven with the significant energy exchange between the SWJ and TEJ, as the jets move close to

each other tend to develop a horizontal shear. The associated geopotential height and zonal wind spatial distribution at 200 hPa during the break period B2000 of the experiments DCW, DCWNC, and CTL experiments show a good agreement with the reanalysis data.

The horizontal shear developed is associated with a synoptic low formation in the head of BoB for the revival of ISM. The experiment DCWNC with KF convection scheme turned off, using SLP daily distribution, do not show a low formation in the BoB after a break phase, which is understandable. The experiments DCW and CTL show a low formation with a minor spatial shift towards the Himalayan region or the coastal regions of India.

Nonetheless, apart from upper atmosphere characteristics, the model results of the rainfall over the CMR of India are pretty vague. This could be due to the model's initial conditions, which are developed through the assimilation of limited observational data and ocean features. From this context, it is evident that the inclusion of the observational ocean characteristics/data into the model is very much necessary, this needs further analysis to be carried out using a coupled model.

Chapter7

Summary, conclusion and future scope of the thesis

Outline of the chapter

In this chapter, I discuss the entire summary and conclusion of the research conducted, as well as a brief outline of the thesis's future scope.

2615 **7.1 Conclusion**

2616 This primary focus of the thesis is to the break monsoon characteristics of ISM and a comparison
2617 of these characteristics with those over the Americas, and the role of breaks in the ISM in its
2618 subsequent revival. For this, using observed rainfall datasets and NCEP/NCAR II reanalysed
2619 datasets, over the broad period from 1979 to 2007, I study 41 known break cases of ISM, which
2620 were identified by Rajeevan et al. (2010).

2621 I find that the formation of a blocking high over sub-tropical west Asia a few days earlier
2622 is crucial for the manifestation of break monsoon conditions over the India region. While there has
2623 been a single case study earlier (Raman and Rao, 1971) that mentioned about the relevance of a
2624 blocking high for the formation of a break monsoon event, the study is essentially of historic
2625 relevance given the limited data that was used in that study. Interestingly, my results from the
2626 analysis of these 41 cases are in conformation with a conclusion from this single case study. My
2627 study shows that the formation of these blocking highs preceding the breaks is due to a quasi-
2628 resonant amplification of planetary waves. Importantly, I also discover similar conditions
2629 associated with break monsoon conditions in the NAM and SAM, respectively. The above
2630 discussion on how the extratropical eddies potentially play a major role in the common ground of
2631 the basic characteristics of the breaks in these three monsoons i.e., ISM, SAM, and NAM is
2632 indicative of a potential global monsoon feature. This motivates us to propose as a unified theory
2633 for the development of breaks during summer monsoon.

2634 In addition, I find a dynamical mechanism involved in the revival of the summer monsoon after a
2635 break. During the peak of significant breaks, the southward shift of the subtropical westerly jet
2636 stream and northward shift of a stronger-than-normal easterly jet facilitate local generation of

2637 barotropic instability mechanism, which apparently leads to formation of a synoptic disturbance.
2638 Indeed, my other computations of energetics and correlation analysis, suggest an increase in the
2639 eddy kinetic energy at the expense of the mean kinetic energy during the breaks, which is also
2640 apparently with the formation of the synoptic disturbances in many cases. Indeed, formation of
2641 such a disturbance is critical to the subsequent revival of the summer monsoon in 61% of the
2642 observed break-to-active revivals.

2643 To verify this hypothesis, I have carried out a few sensitivity experiments using the WRF model.
2644 My sensitivity experiments yield energetics and zonal wind changes quantitatively and
2645 qualitatively comparable to those from the analysis of the NCEP reanalysis 2 data, and thereby
2646 support the afore-proposed mechanism for the revival of active conditions after a break phase.

2647 Thus, the transformation from break to active conditions is also a barotropic process (Rao, 1971,
2648 and Govardhan et al., 2017 which has been reported in this thesis). In this context, the entire life
2649 cycle of formation of the monsoon breaks and reversal to normal monsoon can therefore be seen
2650 as a barotropic adiabatic cycle. This is somewhat similar to the index cycle in a barotropic
2651 atmosphere of the middle and polar latitudes, confirming the hypothesis put forth by Arakawa,
2652 (1961). This conceptual similarity will be elucidated later. Rao (1971) have first pointed out that
2653 the study by Arakawa (1961) is relevant for break-normal cycles. What is more, this concept seems
2654 to apply to the SAM and NAM as well. For example, Webster and Curtin (1975), using constant
2655 density balloon data, suggested that the 18-23 day variations in the SH seem to be a barotropic
2656 interaction between the midlatitudes mean westerlies and ultralong perturbations. They also
2657 mention the potential relevance of the findings from the study by Arakawa (1961) on index cycles
2658 in a barotropic atmosphere. The real test would be to see whether these variations in the zonal
2659 index occur in the observational data over the monsoonal regions. The divergence of eddy

momentum in the region of the jet, decelerating the mid-latitude jet implies, a decrease of zonal kinetic energy and a corresponding increase of eddy kinetic energy or barotropic instability, very similar to the Indian break monsoon case (Rao, 1971 and Govardhan et al., 2017; also see equation (2.2)). Note that the regional variations can also be viewed as a regional manifestation of global waves as suggested by Mishra (2018) for the ISM. Fig. 7.1 illustrates the time latitude cross section of 200 hPa zonal wind averaged for the period June to September 2000. As discussed earlier, I see an increase in zonal wind shear during the break period. Fig. 7.1, which shows the longitudinally averaged 200 hPa zonal winds, shows a 50m/s per 20° latitude meridional shear of zonal winds between 20°N and 40°N on 1st July. While on 1 August, the meridional shear of horizontal wind was 50m/s per 10° latitude, between these latitudes, which is twice the value recorded for 1 July. As shown in chapter 5, this increase in shear generates barotropic instability.

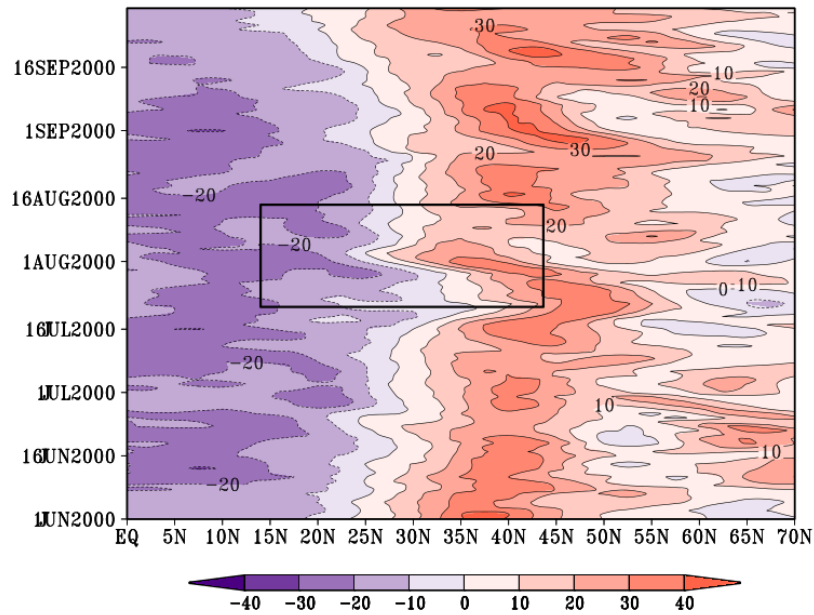


Figure 7.1: Structure of daily U (zonal) wind averaged along 60° E - 95° E over the Indian region, during the period 1 June – 30 September 2000. Black box highlights the Sub-tropical Westerly jet

2674 and Tropical Easterly jet moving close to each other during a break period 01-09 Aug 2000 (case
2675 study). Contouring between -40 to 40 (m/s) with an interval of 10 m/s.

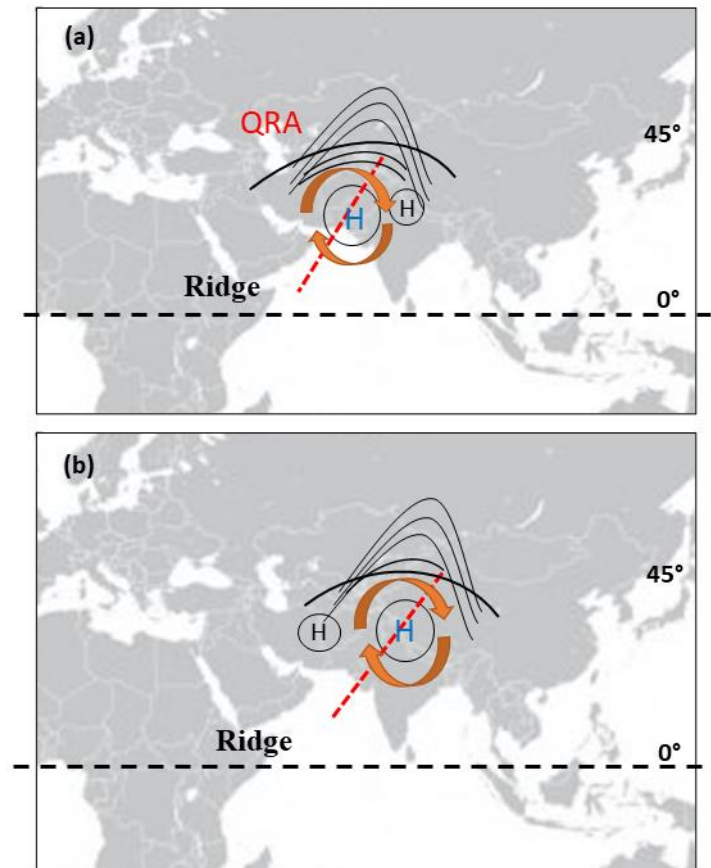
2676 One potential reason for the higher latitude westerly jet and lower latitude easterly jet coming
2677 closer to one another, in all the three monsoonal regions, is the formation of the blocking high, as
2678 conjectured by Ramaswamy (1956, 1962) and confirmed by Fig. 7.1. Similar to the index cycles,
2679 during a break, I find that the upper level sub-tropical westerlies and tropical easterlies shift closer.
2680 This is analogous to the changes in the westerlies for mid-latitude index cycles. In my case, apart
2681 from the blocking, it is still not known exactly what makes these westerly and easterly jets come
2682 closer and later move farther. At this moment, I only conjecture that such shifts in the upper level
2683 zonal winds in the context of monsoons are closely related to the barotropic processes, as suggested
2684 by Chapter 5. As noted earlier, the index cycle variations in the westerly jet can be periodic
2685 Arakawa (1961). However, the variations in the Tropical Easterly jet are irregular. This can be
2686 understood in a general sense; the variations in the westerlies, extending almost across the globe
2687 are more related to the mid-latitude process. The variations in the easterlies may be associated with
2688 the changes in the Tibetan plateau, and even due to changes in the monsoon regions, and North
2689 Indian Ocean. Therefore, the shifts in both jet streams can be purely due to internal variability, and
2690 in some cases, also due to forcing from a co-occurring phenomenon such as an ENSO, IOD,
2691 Atlantic zonal mode, etc., which can modulate the jet streams (e.g. Guan and Yamagata, 2003
2692 GRL; . These aspects need to be explored further in detail. Before the break, all waves with
2693 wavelengths above 10,000 km are only unstable. This value is manifested through the $2D/\sqrt{3}$ limit
2694 Rao (1971). That is, as there are hardly any waves with some lower wavelength, prior to a break,
2695 practically there is no instability during the normal conditions. However, during the break the

corresponding limit is 5000 km; this value is the lowest found by Govardhan et al., (2017; also see chapter 5) cases. Thus, this jet-induced barotropic instability can generate monsoon disturbances.

A careful examination of Fig. 7.1 shows clear periods of high westerlies in the latitudinal band 30°N-50°N. In the period 1 June- 1 October 2000, there are six varying intervals of high westerlies of about 30 m/s. Also, one can note from 1st September the westerlies in lower latitudes, with the change of the seasons. Further, westerlies increase in speed from 1st September. During the break period, starting at the end of July, the easterly jet (-20m/s) moves northward and the westerly jet (20m/s) southward. In the first week of August, I see a strong shear. Thus, in general, I see regular westerly jet cycles, reminiscent of index cycles between 30°N-50°N. The easterly cycles are very irregular. Fig. 7.1, thus shows the existence of “*Monsoon Index Cycle*”, with growth and decay of westerlies over the monsoon region. It would be interesting to speculate the origin of these “*Monsoon Index Cycle*”. One possibility, as mentioned earlier, is the barotropic and adiabatic interactions as suggested by Arakawa (1961). Breaks in monsoon occur in one or more cycles. As noted earlier, a typical break is preceded by the formation of blocking high, which occurs due to the penetration of midlatitudes baroclinic wave as shown by Raman and Rao (1981). In Fig. 7.1, I note the southward shift of westerlies in the last week of July 2000. This southward extension of westerlies favours the penetration of midlatitudes baroclinic wave, since easterlies will not permit propagation of a midlatitude baroclinic wave (Charney, 1969). Overall, I note the occurrence of “*Monsoon Index Cycles*” at least in this monsoon season of 2000. A simple schematic, which elucidates the same, is shown in Fig. 7.2. It would be interesting to see whether these cycles occur in other seasons over India and in other monsoon regions. Considering the regularities in westerlies, interestingly it contributes to the predictability of block formation and consequently breaks.

2719 We should also note that that the breaks in the monsoon could occur due to various other factors
 2720 also. For example, a recent study by Umakanth et al., (2019) suggests the importance of high-
 2721 frequency latent heat changes over the central Indian region, foothills of the Himalayas, and Indo-
 2722 China region for the manifestation of breaks through triggering of meridional wave train trough
 2723 Artic. As mentioned in the introduction, the monsoon breaks can be manifested due to the MJO.

2724



2725

2726 **Figure 7.2:** A simple schematic picture of a barotropic adiabatic cycle at the upper level. The bold
 2727 curved arrows represent the Blocking high. The thin angular continuous lines above the Blocking
 2728 high represent the planetary Rossby waves amplification. The dotted line is the ridgeline and is

inclined along the NE-SW direction. **(a)** Prebreak situation: the genesis of Blocking high due to the QRA **(b)** Break situation: the sustenance of the Blocking high over Indian region due to the deceleration of the zonal wind as shown in the prior Eliassen-Palm flux analysis.

Of course, the Indian summer monsoon variability is controlled by several other factors and drivers. For example, both monsoons and MJO are modulated by the ENSO (e.g., Pai et al., 2016). Pai et al. (2016) show anomalously high number of the ISM break events, which are longer than normal, during the summers when El Niños occur. On the other hand, La Niñas, which are traditionally associated with anomalously surplus ISM rainfall, do not seem to have any association with the frequency of the active cases of the ISM. Importantly, we know that a new type of ENSO, named as ENSO Modoki has been occurring since late 1970s (Ashok et al., 2007; Marathe et al., 2015). In this context, I have carried out a preliminary analysis to document the relative impacts of ENSO types on the active and break monsoon events. I find that both canonical and modoki El Niños anomalously increase the break events, there is a difference in the associated circulation pattern and the distribution of negative rainfall anomalies between them. Interestingly, in a general conformation with Pai et al. (2016), the both La Niña types do not indicate a significant association with active event statistics. Further details are provided in Appendix II, and have been published as part of a book chapter (Feba et al., 2020). As it is, as the ENSO impacts and associated circulation patterns last throughout the season, and importantly act from a relatively remote region, any effect of ENSO must be through the ENSO-associated circulation changes in the Indian theatre. Indeed, formation of a synoptic disturbance, which revives the monsoon from breaks, depends on various other factors such as local SST and moisture availability. The monsoon can also revive as a result of large-scale circulation changes that are forced by external drivers, without

the explicit formation of a synoptic disturbance. In such cases, the manifested instability may be different. From this context, the other large-scale processes likely explain the remaining cases of the break–active transitions that are not explained by the formation of the synoptic disturbances. We should also be mindful that the mechanism for the revival proposed by us is not necessarily out of the purview of the MJO. For example, a particular phase of an MJO co-occurring during a break induced by a blocking high, may potentially cause an anomalous northward shift of the TEJ and expedite the revival of ISM after a break phase. These possibilities need to be explored.

7.2 Future Scope

Finally, there is also a need to explore the synergies and distinctions in the various mechanism proposed for the active and break cycles.

1. The relative roles of the impact of the midlatitudes versus the role of MJO during breaks and active conditions of ISM.
2. To observe the future changes of instabilities associated with intraseasonal variability of ISM due to the increasing GHGs, aerosols, and land use land cover changes.

Appendix-I

Barotropic instability problem

These are two ways of studying the development of disturbances, namely,

1. Eigen value problem (Dynamic Meteorology by Holton (2004))
2. The initial value problem (Kuo 1953; also see Chapter 6 of Tropical meteorology: An Introduction by Krishnamurti (2013))

Here we have adopted the initial value problem. The symbols/notations representing various variables/parameters in the appendix are listed below (Table A1).

In order to estimate the energy exchange between the basic zonal current and a superimposed disturbance in a barotropic, non-divergent and frictionless atmosphere, we use the barotropic vorticity equation in the form.

$$\frac{d}{dt}(f + v_E) = 0 \quad (1)$$

where

$$\frac{d}{dt} = \frac{\partial}{\partial t} + u \frac{\partial}{\partial x} + v \frac{\partial}{\partial y}$$

$$f = 2\Omega \sin(\phi) ; \text{ Coriolis force term}$$

$$\phi - \text{Latitude}$$

$v_E = \nabla^2 \psi$ - Relative vorticity

ψ - Stream function

u and v are the zonal and meridional components of the horizontal velocity vector, and can be expressed as

$$u = \frac{-\partial \psi}{\partial y}; v = \frac{\partial \psi}{\partial x}$$

As can be understood, x & y are the co-ordinate axes taken positive towards east and north respectively.

Linearization of equation (1) yields

$$\frac{\partial}{\partial t} \nabla^2 \psi + U \frac{\partial}{\partial x} \nabla^2 \psi + \frac{\partial \psi}{\partial x} \left(\beta - \frac{\partial^2 U}{\partial y^2} \right) = 0 \quad (2)$$

U is the mean zonal current and ψ is the stream function for the perturbation flow.

$$\beta = \frac{df}{dy} \text{ --- Rossby factor}$$

A typical solution for equation (2) will be

$$\psi = A(y, t) \sin(kx) + B(y, t) \cos(kx) \quad (3)$$

Where $k = \frac{2\pi}{L}$ is the wave number, and L the Wavelength

Substituting solution (3) in the equation (2) and equating the coefficients of $\sin(kx)$ and $\cos(kx)$ terms, we get the following equations:

$$\frac{\partial^2}{\partial y^2} \left(\frac{\partial A}{\partial t} \right) - k^2 \frac{\partial A}{\partial t} = -U \left(k^3 B - K \frac{\partial^2 B}{\partial y^2} \right) + kB \left(\beta - \frac{\partial^2 U}{\partial y^2} \right) \quad (4)$$

$$\frac{\partial^2}{\partial y^2} \left(\frac{\partial B}{\partial t} \right) - k^2 \frac{\partial B}{\partial t} = U \left(k^3 A - K \frac{\partial^2 A}{\partial y^2} \right) - kA \left(\beta - \frac{\partial^2 U}{\partial y^2} \right) \quad (5)$$

(4) and (5) are two unknown equations in two unknowns, $\frac{\partial A}{\partial t}$ and $\frac{\partial B}{\partial t}$ and so form a closed system of equations.

From the prescribed initial values of u , A , B , and $\frac{\partial^2 U}{\partial y^2}$, and with proper boundary conditions, we can find solutions for $\frac{\partial A}{\partial t}$ and $\frac{\partial B}{\partial t}$.

Initial conditions

$$A_0 = 0 \text{ and } B_0 = a \sin y, \quad l = \frac{\pi}{D} \quad (6)$$

where D is the channel width, and suffix 'o' represents the initial value. As pointed by Platzman (1952), it is desirable to take initial conditions in such a way as to make the first derivative of perturbations kinetic energy zero. As would be shown later specifically in equation (10), the above condition (6) will fulfil the requirement.

Boundary conditions-- Meridional direction

$$A=0 \text{ at } y=0 \text{ and } y=D; \quad \frac{\partial A}{\partial t}=0 \text{ and } \frac{\partial B}{\partial t}=0 \text{ at } y=0 \text{ and } y=D \quad (7)$$

In the X -direction we assume that the disturbance quantities have cyclic periodicity at intervals of one wavelength L . If Q is any disturbance quantity, then $Q(x,y)=Q(x \pm L,y)$. Thus it is sufficient to

consider the domain of integration as the area bounded by one wavelength 'L' in the X- direction and distance D in the y- direction to evaluate various kinds of energies.

Time tendency of Amplitudes

Amplitudes A and B after a time Δt are given by the Taylor's series

$$A(\Delta t) = A_o + \left(\frac{\partial A}{\partial t}\right)_o \Delta t + \frac{1}{2} \left(\frac{\partial^2 A}{\partial t^2}\right)_o \Delta t^2 \pm \dots \quad (8)$$

$$B(\Delta t) = B_o + \left(\frac{\partial B}{\partial t}\right)_o \Delta t + \frac{1}{2} \left(\frac{\partial^2 B}{\partial t^2}\right)_o \Delta t^2 \pm \dots \quad (9)$$

If Δt is sufficiently small, the above series can be truncated after the second derivative. This will no doubt introduce some error in the forecasted amplitudes. Nevertheless, it is not an essential shortcoming as shown by the results.

With initial conditions (6), (5) becomes

$$\frac{\partial^2}{\partial y^2} \left(\frac{\partial B}{\partial t}\right) - k^2 \left(\frac{\partial B}{\partial t}\right) = 0 \quad (10)$$

It can easily be shown from (10) and (7) that $\left(\frac{\partial B}{\partial t}\right)_o = 0$, everywhere,

Equations for $\left(\frac{\partial^2 A}{\partial t^2}\right)_o$ and $\left(\frac{\partial^2 B}{\partial t^2}\right)_o$ can be obtained by differentiating (4) and (5) with respect to time. They take the form

$$\frac{\partial^2}{\partial y^2} \left(\frac{\partial^2 A}{\partial t^2} \right) - k^2 \frac{\partial^2 A}{\partial t^2} = -U \left(k^3 \frac{\partial B}{\partial t} - K \frac{\partial^2}{\partial y^2} \left(\frac{\partial B}{\partial t} \right) \right) + k \frac{\partial B}{\partial t} \left(\beta - \frac{\partial^2 U}{\partial y^2} \right) \quad (11)$$

$$\frac{\partial^2}{\partial y^2} \left(\frac{\partial^2 B}{\partial t^2} \right) - k^2 \frac{\partial^2 B}{\partial t^2} = U \left(k^3 \frac{\partial A}{\partial t} - K \frac{\partial^2}{\partial y^2} \left(\frac{\partial A}{\partial t} \right) \right) - k \frac{\partial A}{\partial t} \left(\beta - \frac{\partial^2 U}{\partial y^2} \right) \quad (12)$$

Initial conditions (6) are used to obtain (11) and (12) since $\left(\frac{\partial B}{\partial t} \right)_0 = 0$ from equations (11) and (7)

it can easily be shown that $\left(\frac{\partial^2 A}{\partial t^2} \right)_0 = 0$ everywhere, so (8) and (9) reduce to

$$A(\Delta t) = \left(\frac{\partial A}{\partial t} \right)_0 \Delta t \quad (13)$$

$$B(\Delta t) = B_0 + \frac{1}{2} \left(\frac{\partial^2 B}{\partial t^2} \right)_0 \Delta t^2 \quad (14)$$

so after time Δt , ψ is given by

$$\psi(\Delta t) = A(\Delta t) \sin(kx) + B(\Delta t) \cos(kx)$$

$$\psi(\Delta t) = R_\psi \cos(kx - \delta\psi) \quad \text{where } R_\psi = [A^2(\Delta t) + B^2(\Delta t)]^{1/2}$$

$$\text{and } \tan(\delta\psi) = \frac{A(\Delta t)}{B(\Delta t)} \quad (15)$$

Thus the amplitude and phase of ψ wave can be found after time Δt from (15)

Initial change of kinetic energy

The rate of change of kinetic energy may be regarded as the rate of amplification of the disturbances. If it is positive, kinetic energy tends to increase with time, and disturbance is said to be unstable. If it is negative, kinetic energy tends to decrease, and the disturbance is said to be stable or damping. If the rate of change of kinetic energy is zero, the kinetic energy remains constant, and the disturbance is said to be neutral.

The kinetic energy of the disturbance is given by

$$K_r = \int_0^D \int_0^L \frac{u^2 + v^2}{2} dx dy \quad (16)$$

But

$$u = \frac{-\partial \psi}{\partial y} = - \left[\frac{\partial A}{\partial y} \sin(kx) + \frac{\partial B}{\partial y} \cos(kx) \right] \quad (17)$$

$$v = \frac{\partial \psi}{\partial x} = -k[A \cos(kx) - B \sin(kx)] \quad (18)$$

Inserting (17) and (18) into (16) we get

$$K_r = \frac{\pi}{2k} \int_0^D \left[\left(\frac{\partial A}{\partial y} \right)^2 + \left(\frac{\partial B}{\partial y} \right)^2 + k^2 (A^2 + B^2) \right] dy \quad (19)$$

Differentiating (19) with respect to time and using (4), (5) and (7) we get

$$\frac{\partial K_r}{\partial t} = -\pi \int_0^D U \left[A \frac{\partial^2 B}{\partial y^2} - B \frac{\partial^2 A}{\partial y^2} \right] dy \quad (20)$$

The equation for the time change of the zonal wind is

$$\frac{\partial U}{\partial t} = \frac{-\partial}{\partial y} \overline{uv} \quad (21)$$

where the overbar denotes a zonal average.

Multiplying (21) by U and integrating over the region we get the equation for the time change of zonal kinetic energy as

$$\frac{\partial}{\partial t} K_{r_z} = - \int_0^D \int_0^L U \frac{\partial}{\partial y} \overline{uv} dx dy \quad (22)$$

Where K_{r_z} is the zonal kinetic energy given by

$$K_{r_z} = - \int_0^D \int_0^L \left(\frac{U^2}{2} \right) dx dy$$

Using (17) and (18)

$$\overline{uv} = \frac{k}{2} \left[B \frac{\partial A}{\partial y} - A \frac{\partial B}{\partial y} \right] \quad (23)$$

Using (23) and (22) becomes

$$\frac{\partial}{\partial t} K_{r_z} = \pi \int_0^D U \left[A \frac{\partial^2 B}{\partial y^2} - B \frac{\partial^2 A}{\partial y^2} \right] dy \quad (24)$$

It is seen from (20) and (24) that the right hand side of (24) is the same as the right hand side of (20) but with opposite sign. Thus this term represents the interaction between the zonal and perturbation kinetic energies.

In view of our initial conditions,

$$\left(\frac{\partial K_r}{\partial t} \right)_0 = \left(\frac{\partial K_{r_z}}{\partial t} \right)_0 = 0 \quad (25)$$

Thus, as pointed out, earlier our initial conditions are such that the first derivative of perturbation kinetic energy is made equal to zero. So we have to consider the second derivative of perturbation kinetic energy K_r , in order to find out the initial change of kinetic energy, then

$$\frac{\partial^2}{\partial t^2} K_{r_z} = -\pi \int_0^D U \left[\left(\frac{\partial A}{\partial t} \right)_o \frac{\partial^2 B_o}{\partial y^2} - B_o \frac{\partial^2}{\partial y^2} \left(\frac{\partial A}{\partial t} \right)_o \right] dy \quad (26)$$

Initial conditions are used to get (26)

Now, we will study the stability properties of different zonal currents with initial disturbance

$$\psi = a \sin(ly) \cos(kx), \text{ i.e., } B_o = a \sin(ly) \text{ and } A_o = 0 \quad (27)$$

The actual forms of the zonal current will be selected in such a way as to study different aspects of the problem.

(i) the zonal current U is given by

$$U = c \cos(2ly) \quad \text{Where } l = \frac{\pi}{D}$$

We shall discuss this symmetric mean zonal current. This profile has two inflections points (where $\frac{\partial^2 u}{\partial y^2} = 0$) midway between the axis of the flow and the walls. Kuo (1949) found that the presence of flex points plays an important role in the barotropic stability problem.

We now need to solve equation (4) for the above prescribed zonal wind profile and B_0 (given by equation 27). β is given as

$$\beta = \frac{2\Omega \cos(\phi)}{R} = \frac{\Omega}{R} (1 + \cos(2\phi))^{1/2} = \frac{\Omega}{R} (1 + \alpha \cos(ly)) \quad (28)$$

where R is the radius of the earth and $\alpha < 1$,

With the prescribed expressions for U , B and β , equation (4) is solved with the boundary conditions (7) to give

$$\left(\frac{\partial A}{\partial t}\right)_0 = \frac{-E}{k^2 + l^2} \sin(ly) - \frac{F}{k^2 + 4l^2} \sin(2ly) - \frac{G}{k^2 + 9l^2} \sin(3ly) - \dots \quad (29)$$

Where

$$E = \frac{ka\Omega}{R} + \frac{k^3ac}{2} - \frac{3kac l^2}{2}$$

$$F = \frac{ka\Omega\alpha}{2R}$$

$$G = \frac{3}{2}l^2kac - \frac{k^3ac}{2}$$

with the expressions for $\left(\frac{\partial A}{\partial t}\right)_o$, B_o and U , the integral in (26) is evaluated to give

$$\left(\frac{\partial^2 K_r}{\partial t^2}\right)_o = \frac{ka^2c^2l^2\pi D}{k^2+9l^2}(3l^2 - k^2) \quad (30)$$

$$\begin{aligned} \left(\frac{\partial^2 K_r}{\partial t^2}\right)_o &= 0 & \text{when } L &= \frac{2D}{\sqrt{3}} \\ &> 0 & \text{when } L &> \frac{2D}{\sqrt{3}} \\ &< 0 & \text{when } L &< \frac{2D}{\sqrt{3}} \end{aligned} \quad (31)$$

Thus, the neutral wavelength $L < \frac{2D}{\sqrt{3}}$ separates the stable shorter waves and unstable longer waves. It is to be noted that the terms due to earth's rotation will not appear in (30). So earth's rotation will not contribute to $\left(\frac{\partial^2 K_r}{\partial t^2}\right)_o$ with the symmetric profile for U considered. $\left(\frac{\partial^2 K_r}{\partial t^2}\right)_o$ is maximum at a wavelength 2.1 D and so is the most unstable disturbance.

Table A1: A complete list of the symbols/notations representing various variables/parameters in the appendix.

Symbol	Definition
f	Coriolis parameter
ϕ	Latitude
$v_E = \nabla^2 \psi$	Relative vorticity
ψ	Stream function for perturbation flow
u	Zonal wind
v	Meridional wind
U	Mean zonal wind
$\beta = \frac{df}{dy}$	Rossby factor
$k = \frac{2\pi}{L}$	Wave number
t	time
L	Wavelength
D	Channel width
Q	Any disturbance quantity
A, B	Amplitude
K_r	Perturbation Kinetic energy
K_{r_z}	Zonal kinetic energy
R	Radius of Earth
dx & dy	Incremental zonal & meridional distances used in integration/differentiation
a, α, c	

Ω	Wavelength
ω	Vertical 'p' velocity
Δ	Angular speed of the earth
Δ^2	Del operator applied to a quantity which varies on an isobaric surface
	Laplacian operator

Appendix - II

Intraseasonal Variability of ENSO Modoki - ISM Teleconnections

The intraseasonal variability of the Indian Summer Monsoon (ISM) rainfall has is characterized by periods of wet (active) and dry (break) rainfall activity mainly prominent over the monsoon core region of central India. It is manifests as a 30-60-day oscillation (Gadgil and Joseph 2003; Goswami and Mohan 2001; Goswami and Xavier 2003; Keshvamurthy and Sankar Rao 1992; Krishnamurti and Bhalme 1976; Krishnamurti and Subrahmanyam 1982; Sikka and Gadgil 1980; Yasunari 1979). Goswami and Chakravorty elaborate on these aspects of active and break spells of monsoon in their recent review (Goswami and Chakravorty 2017).

Pai et al. (2016) showed that the composite rainfall patterns of active days during El Niños in the 1901-2014 period show only a slight difference from those during La Niñas, except for slightly stronger negative rainfall anomalies along the foothills of Himalayas during El Niños. However, in the case of break monsoon, the positive composite rainfall anomalies along the foothills of Himalayas during La Niña years are both stronger and extending more westwards (North-eastern states) compared to that of El Niño years.

Motivated by this, in my study, I carried out composite analysis of intraseasonal summer monsoon rainfall anomalies for each type and phase of ENSOs during the 1951-2019 period was obtained by averaging daily rainfall anomalies during the break monsoon and active monsoon periods identified by Rajeevan et al. (2006; 2010), and are shown in Figures A2.1 and A2.2, respectively. TableA2.1 lists the aggregate number of break and active monsoon days in all the canonical El

Niño years, those in El Niño Modoki years, and those in two types of La Niña years respectively (Marathe et al., 2015), during the 1951-2007 period. The El Niño and El Niño Modokis are seen to be predominantly associated with higher number of break conditions over India (Figures A2.1). Interestingly, unlike its seasonal signature (e.g., Figure 5, Ashok et al., 2019), the El Niño Modoki events are associated with negative rainfall anomalies along the monsoon trough during the break periods (Figure A2.1). However, the association between the type and phase of ENSO events with the active monsoon condition is not clear, unlike for the break events. Heuristically, I would expect that the number of active days is more during La Niñas. But we also see many active days during El Niños as well. Of course, the relatively high rainfall during El Niño Modokis over the core monsoon region may be due to the fact that the anomalously negative seasonal rainfall anomalies during these events are more concentrated in peninsular India, with a positive association in the core monsoon region (Fig. 3.4; also see Ashok et al., 2007 and 2019). There is an argument that propensity of the break and active cycles in any monsoon season decide whether the relevant seasonal monsoonal rainfall is below or above normal (Goswami and Chakravorty, 2017; Lau et al., 2012 and references therein).

Table A2.1 List of total number of active and break monsoon days (and spells) during El Niño, El Niño Modoki, La Niña Modoki and La Niña years during the period of 1951-2007.

Event	Years	Total No. of active rainfall days over CMR	Total average rainfall over CMR (mm)	Total No. of break rainfall days over CMR	Total average rainfall over CMR (mm)

All	1951-2007	414	7.67	404	-4.06
El Niño Modoki	1967, 1977, 1991, 1994, 2002, 2004	36	8.84	52	-2.25
El Niño	1957, 1965, 1972, 1982, 1987, 1997	40	5.24	74	-4.29
La Niña Modoki	1975, 1983, 1998, 1999	19	7.37	32	-2.12
La Niña	1970, 1973, 1988, 2007	43	8.39	25	-3.05

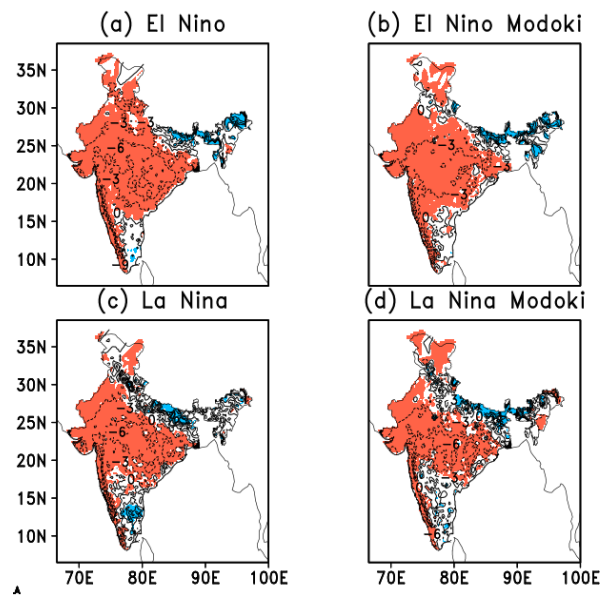


Figure A2.1: (clockwise) Composite of rainfall anomalies during break monsoon days in the El Niño, El Niño Modoki, La Niña Modoki & La Niña years (in contours, at 5 mm/day intervals).

The shaded region indicates 95% confidence using two tailed Student's t test (orange - negative; sky-blue - positive).

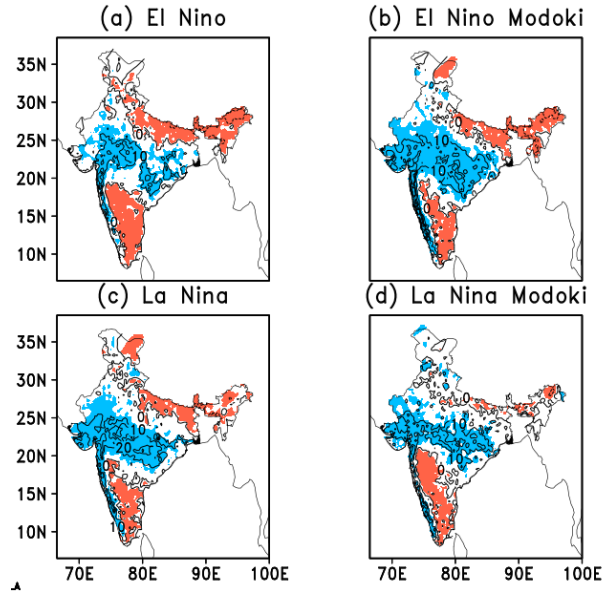


Figure A2.2: (clockwise) Composite of rainfall anomalies during active monsoon days of Indian region in the El Niño, El Niño Modoki, La Niña Modoki & La Niña years (in contours, at 10 mm/day intervals). The shaded region indicates 95% confidence using two tailed Student's t test (orange - negative; sky-blue - positive).

We see from Figure A2.3(b) that the highest number of break days in a year are seen in 2002, associated with a strong El Niño Modoki. The 1965 El Niño seems to be also associated with the third-highest number of break monsoon days. When the total number of break days are concerned (Table A2.1), the canonical El Niño events are more proficient. During years with a warm-phased event in the tropical Pacific, break spells of longer length are more common, such as the El Niño

years during 1951-2007 showing longer break spells compared to the active spells & El Niño Modoki in 2002 & 2004, rather than during the La Niña years. All this suggests that the ENSOs affect the intraseasonal variability of the ISM through the modulation of background seasonal circulation. Moreover, there should be a significant role of internal variability, as evidenced by that the second-highest number of breaks has occurred in 1966 (Table A2.1) when neither a strong warm ENSO nor a positive IOD have co-occurred (Figures not shown).

Interestingly, we also see from Table A2.1 that the number of active days during either of these two El Niños is not negligible. The number of active days associated with La Niñas during 1951-2007 is higher than those associated with the La Niña Modokis. The highest number of active days during summer monsoon occurred in 2006 when no discernible ENSO event was observed. These are likely associated with the co-occurring strong positive IOD event (Cai et al., 2009; Kucharski et al., 2020). The above-normal active cycles during the strong canonical El Niño in 1997 and those during the strong El Niño Modoki during 1994 can be attributed to the opposing effect of the co-occurring strong positive IOD events in years such as 1997 (Ashok et al., 2001). This of course needs to be verified through modelling experiments. In short, we observe that both El Niños and El Niño Modokis apparently facilitate higher than normal break days in Indian summer monsoon, the former particularly so. However, an analogous unique impact of the La Nina's on active cycles is not that evident, which supports the argument of Pai et al. (2016).

Figure A2.1 depicts a composite analysis of daily rainfall anomalies during break conditions associated with the two types of El Niños, El Niño Modoki, and corresponding La Niñas. A similar analysis for the active cases is shown in Figure A2.2. We see negative rainfall anomalies over most of India, except the anomalously positive rainfall anomalies along the Himalayan foothills. I do not see a big difference between the signatures of two types of El Niños, except that the

canonical events seem to introduce higher deficit in daily rainfall during the breaks. Notably, the impact of El Niño Modokis on the seasonal summer monsoon rainfall along the monsoon trough (e.g., Fig. 5 of Ashok et al., 2019) is apparently opposite to that during the break monsoon days. In addition, while the El Niño Modokis and El Niños are associated with a positive rainfall anomalies in the break monsoon distribution and negative rainfall anomalies during active conditions over the Northeast India, these indications are subject to the data quality in the r region (Soraisam et al., 2018). The surplus rainfall footprints during active cases associated with all types ENSOs seem to be only confined to the monsoon trough region (Fig. A2.2), unlike those during the corresponding break conditions. The La Niña Modokis do not seem to be as proficient as the La Niñas in causing surplus rainfall in the central Indian region during the active phase of ISM (Figures A2.1 and A2.2, lower panels).

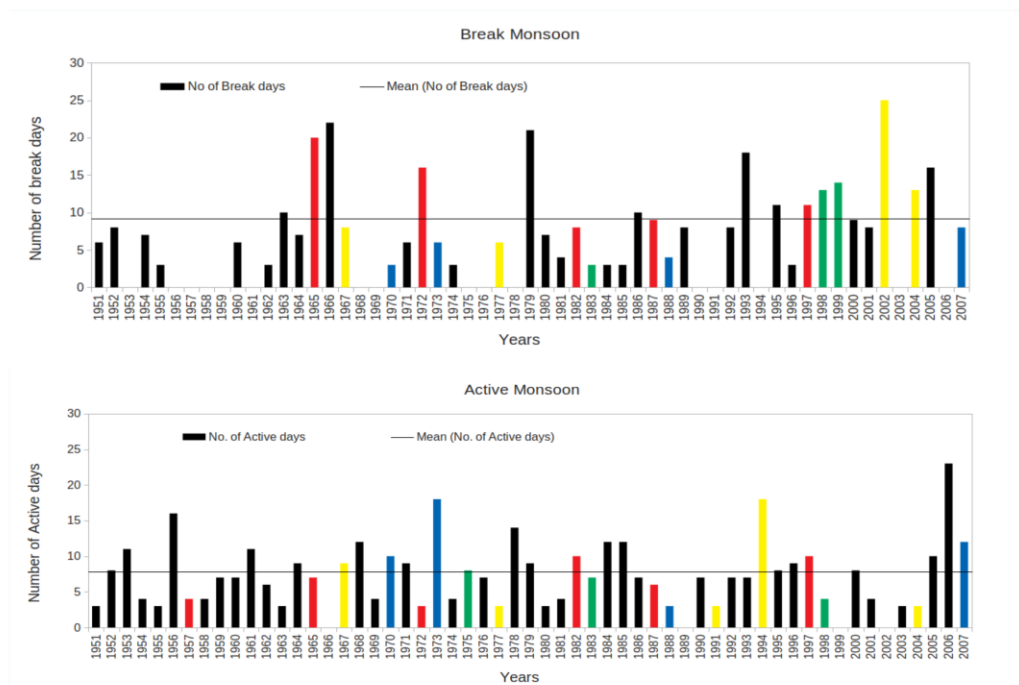


Figure A2.3: (a - Top) Histogram of break days during 1951-2007 based on Rajeevan et al., (2008) criterion. The El Niño, El Niño Modoki, La Niña, La Niña Modoki are designated by the Bars in Red, Yellow, Blue and Green, respectively. (b - Bottom) Same as Figure 3.7 (a) but for active days.

Figures A2.4 a and A2.4 b shows the composite of daily anomalous 850 hPa velocity potential and the divergent wind vectors over all the break and active cases during the 1951-2007 period. Notably, during the breaks, I see a relatively small zone anomalous convergence confined over the central tropical pacific (Fig. A2.4 a). On the other hand, during the active monsoon conditions, the zone of convergence extends well over the western pacific region (Fig. A2.4 b); this looks like an enhanced seasonal mean Walker circulation, with a strong signal over the tropical western pacific. In conformation with the active conditions, we find a convergence over the Indian sub-continent too, along with convergence from the equatorial Indian Ocean (Fig. A2.4 b). This is more prominent over the eastern portion, importantly, over the equatorial Pacific region at the level of 850 hPa, the magnitude of the divergent vectors is strong during the active phase of the Indian summer monsoon, with an opposite signature during the breaks. (Fig. A2.4 b).

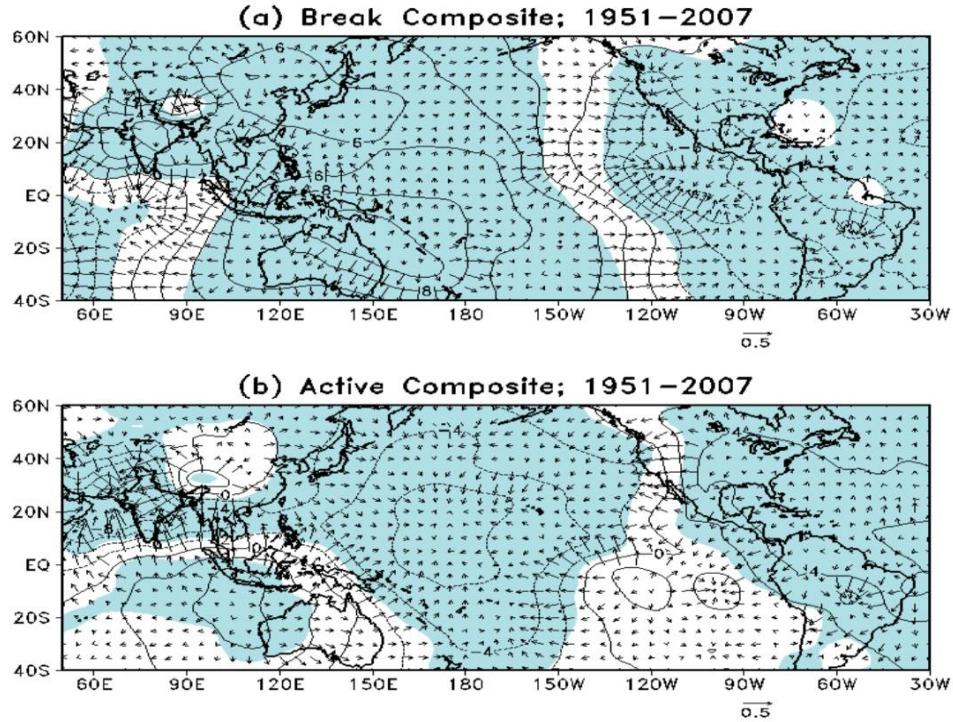


Figure A2.4: 850hPa Velocity potential anomalies (in contours, at 2 m²/s intervals) and divergent (convergent) wind vectors for all the (a) break and (b) active cases during the 1951–2007 period. The shaded region indicates 95% confidence using two tailed Student’s t test.

Figure A2.5 illustrates the composites of the daily anomalous 850 hPa velocity potential and the divergent wind vectors for summer monsoon breaks during all phases of co-occurring ENSOs and ENSO Modokis. Figure A2.5a shows that during canonical El Niños, I find a strong zone of anomalous divergence over the western Indian sub-continent. An anomalous zone of convergence over the northeast Indian region is also shown. These anomalous signatures explain the dynamics behind the rainfall anomalies during breaks associated with the canonical El Niños (Fig. A2.1). During the El Niño Modokis as well, we still see anomalous divergence signature, statistically significant at 95% confidence level. The composites of the active cases during all phases of ENSO

and ENSO Modoki indicate anomalous convergence mainly over the northern Head Bay of Bengal (Fig. A2.6), except in case of La Niña Modoki where the anomalous convergence is more prominent over the Arabian Sea. We also observe a strong convergence zone (Fig. A2.6 a) in the tropical eastern pacific case of El Niño, and in northern tropical central Pacific region in the case of El Niño Modokis with a slight weakened convergence over the western equatorial Pacific region (Fig. A2.6 b) and vice versa in the cases of La Niña (Fig. A2.6 c) and La Niña Modoki (Fig. A2.6 d) respectively in these locations. . The anomalous convergence signal over the Indian regions fall slightly short of statistical significance during active cases associated with canonical La Niñas.

The above composite analysis, essentially a preliminary effort, when combined with Table A2.1 suggest that changes in seasonal circulation may play potential role in the active-break cycles. There seems to be a relatively stronger association between the El Niños – be it canonical or Modokis -and breaks. However, a similar relatively stronger relation between active conditions with La Niñas, which I expect because of the propensity of La Niñas to provide a favourable background condition for the active conditions, is not that conspicuous. These subtle conjectures need to be supported by further observations using multiple observations -particularly station level observations, and a more exhaustive dynamical analysis. We would also note from Figures A2.5 and A2.6 that the large scale circulation changes associated with monsoon break and active conditions due to ENSOs have their signatures as far as Australia, indicating the complex dynamics that may be involved, and need further attention. The analysis is also subject to the fact that we have not factored the impacts from other climate drivers such as the IOD, Atlantic Zonal Mode, and Madden-Julian Oscillation into our analysis.

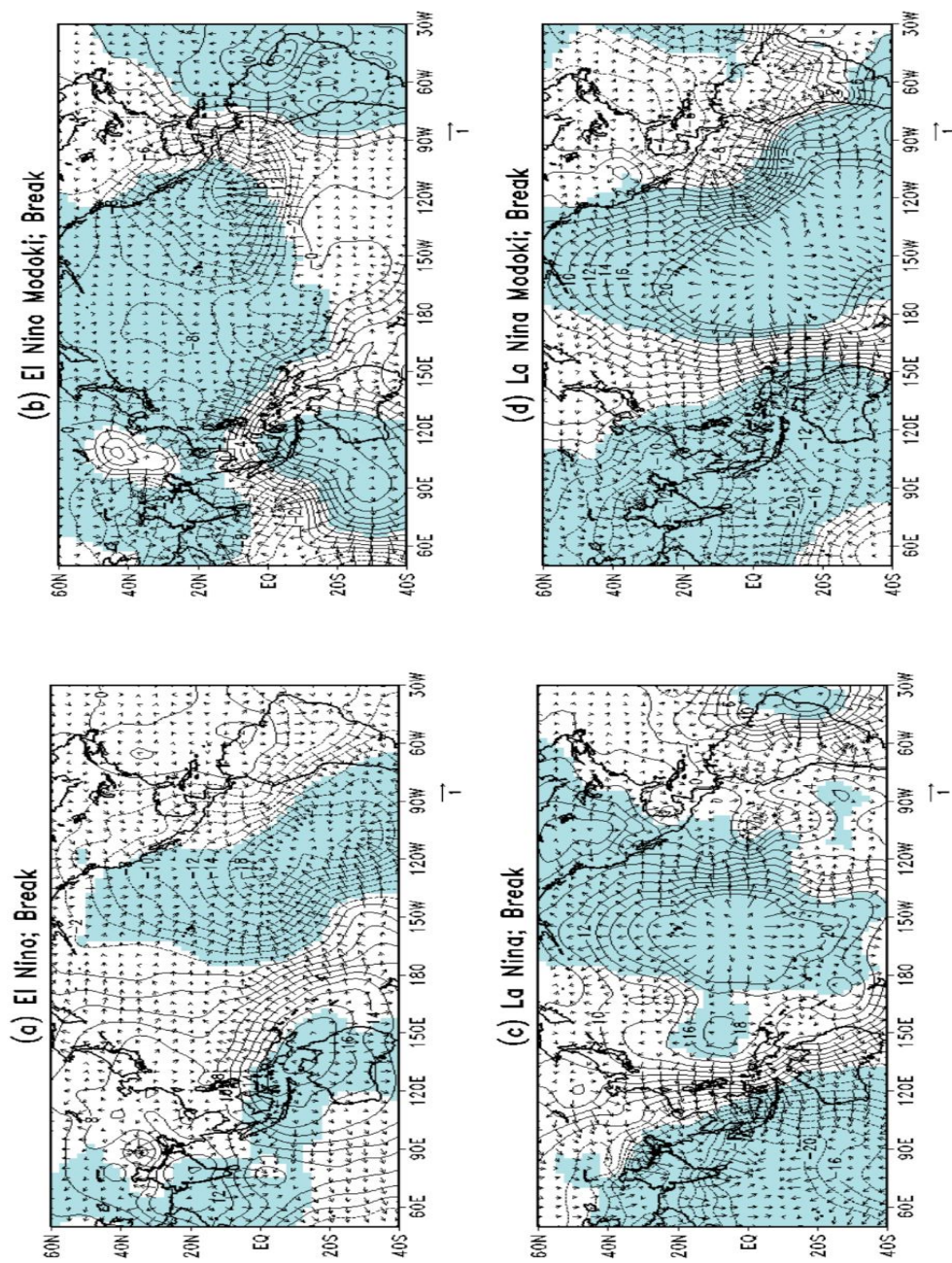


Figure A2.5: 850 hPa Velocity potential anomalies (in contours, at 2 m²/s intervals) and divergent (convergent) wind vectors for the break cases in (a) El Niño, (b) El Niño Modoki, (c) La Niña and

(d) La Niña Modoki years. The shaded region indicates 95% confidence using two tailed Student's t test.

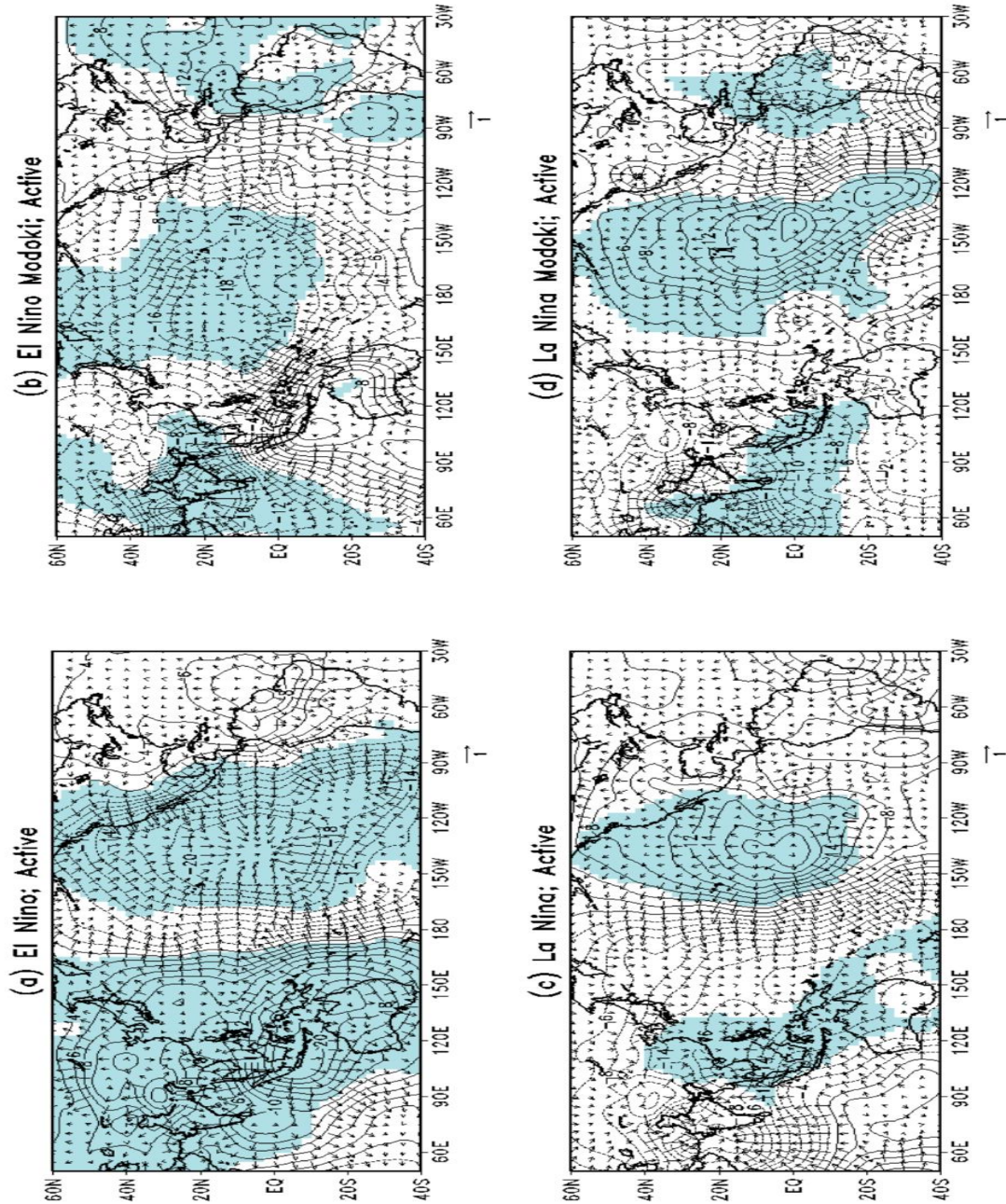


Figure A2.6: 850 hPa Velocity potential anomalies (in contours, at 2 m²/s intervals) and divergent (convergent) wind vectors for the active cases in (a) El Niño, (b) El Niño Modoki, (c) La Niña and

(d) La Niña Modoki years. The shaded region indicates 95% confidence using two tailed Student's t test.

ENSO & ENSO Modoki Index

We use the standard Niño3 and Niño4 indices (area-averaged SST anomaly over the regions 5°N-5°S, 150°W-90°W and 5°N-5°S, 160°E-150°W respectively) to define the canonical ENSO events. For ENSO Modoki, we use the ENSO Modoki Index (EMI) as defined by Ashok et al. (2007) and Marathe et al. (2015).

$$EMI = [SSTA]_A - 0.5*[SSTA]_B - 0.5*[SSTA]_C$$

with SSTA area-averaged over the regions,

A (165°E-140°W, 10°S-10°N), B (110°W-70°W, 15°S-5°N), and C (125°E-145°E, 10°S-20°N), respectively.

Criterion for Break/Active spells

In evaluating the potential impact of the ENSO Modoki on the active and break monsoon events, the rainfall anomalies are obtained by subtracting the daily rainfall climatology for the period 1951-2019 from the daily mean value. We follow the well-accepted Rajeevan et al. (2006) criterion to identify the active and break spells/phase of the ISM, which is based on daily rainfall data. According to the criterion, the active (break) episodes are identified during the peak monsoon months (July to August) as the period through which the standardized daily rainfall anomaly in the interested area exceeds (is less than) one standard deviation. If this continues for at least 3 consecutive days, then these are called as Active (Break) spells. Here, we applied the criterion to identify the active and break episodes in the monsoon core zone limited to the region 18°N-28°N, 65°E-88°E.

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